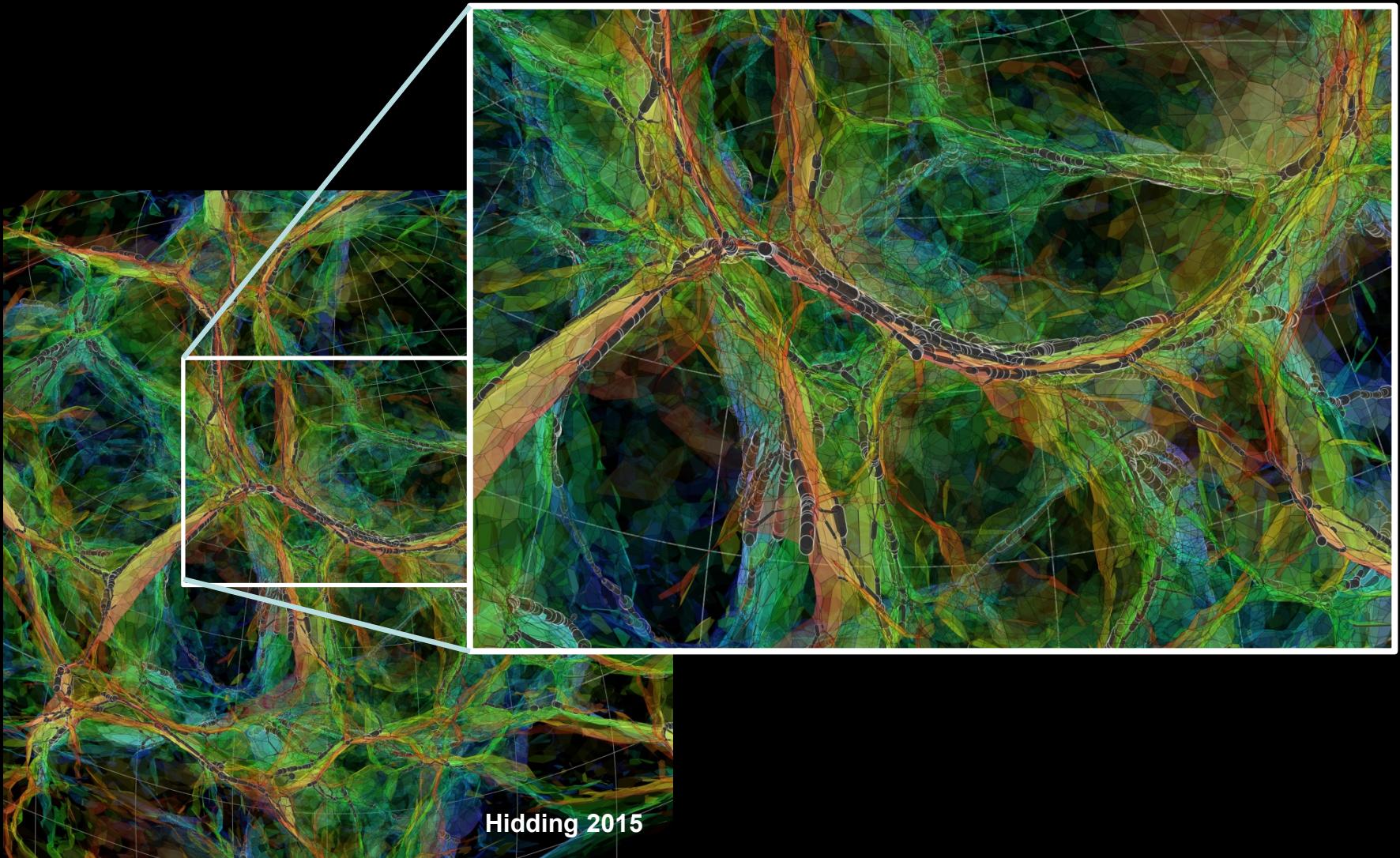


the Cosmic Web: Complexity and Connectivity of the largest structure in the Universe

Rien van de Weijgaert,
2nd workshop on Topological Methods in Data Analysis, Heidelberg, Oct. 2020

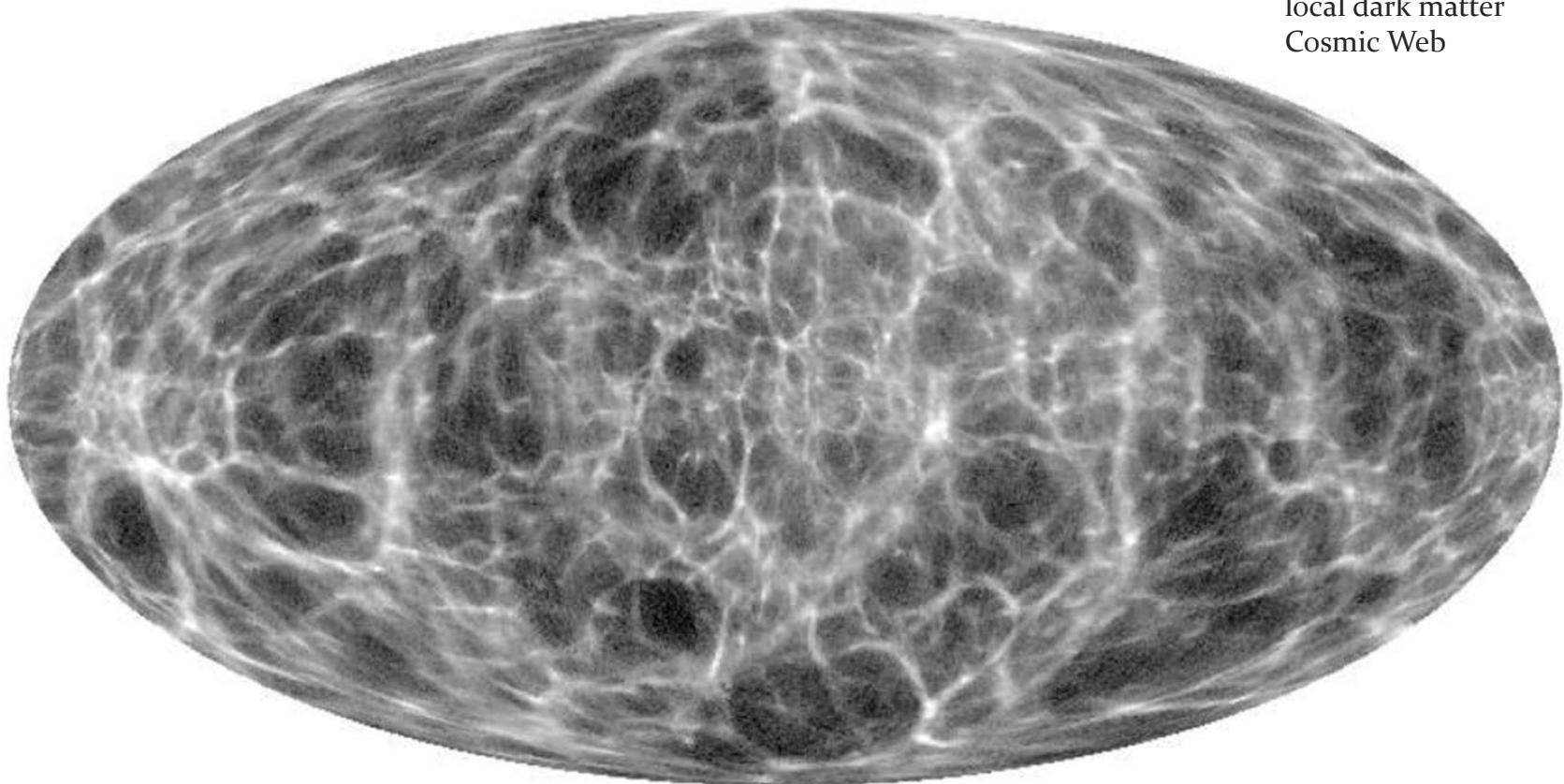
Pisces-Perseus Supercluster



local Cosmic Web: 2MRS

most detailed reconstruction
of the

local dark matter
Cosmic Web



1.0 6.0

Courtesy: Francisco Kitaura

Cosmic Web

S. Rieder 2014

on scales of ~5-100s Mpc

complex weblike pattern

in which
matter, gas & galaxies
aggregate in

- compact clusters,
 - elongated filaments
 - flattened sheets
- around
- cosmic voids

The Cosmic Web

Physical Significance:

- **Manifestation mildly nonlinear clustering:**
Transition stage between linear phase
and fully collapsed/virialized objects
- **Weblike configurations contain**
cosmological information:
eg. *Void shapes & Alignments*
- **Cosmic environment within which to understand**
the formation of galaxies.

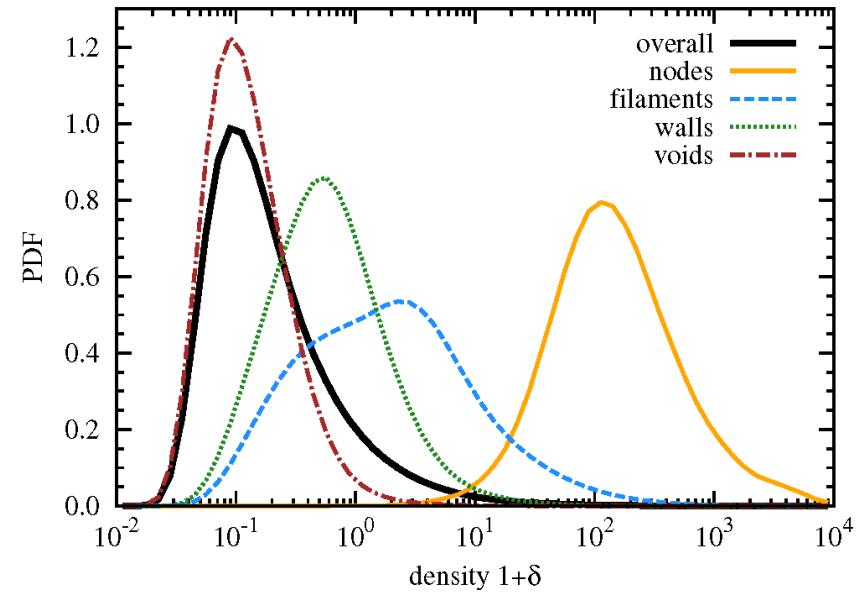
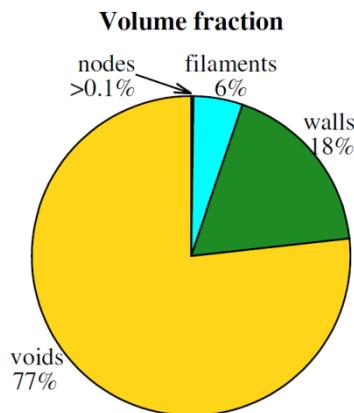
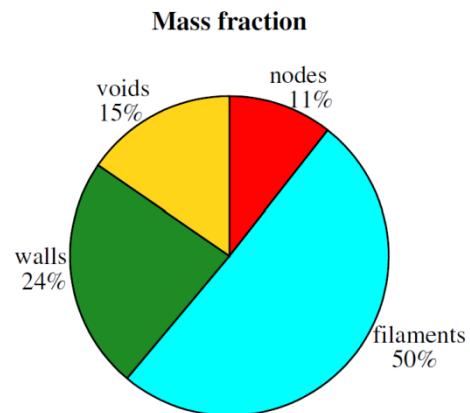
Cosmic Web:

Density-Morphology Connection

Mass & Volume content
Web morphologies



Density distribution
Individual morphologies



Cautun et al. 2014

Cosmic Structure Formation

Cosmic Structure Formation



- Gravitationally driven growth & amplification
- Primordial Gaussian density & velocity fluctuations
- first nonlinear stage formation process:
 - contraction & collapse into walls and filaments
 - connecting at compact high-density nodes
 - expansion & evacuation low-density regions
- Cosmic Web: first cosmic structure emerging from primordial universe

M. Aragon-Calvo

Cosmic Web

Setting the Scene

Cosmic Web Characteristics

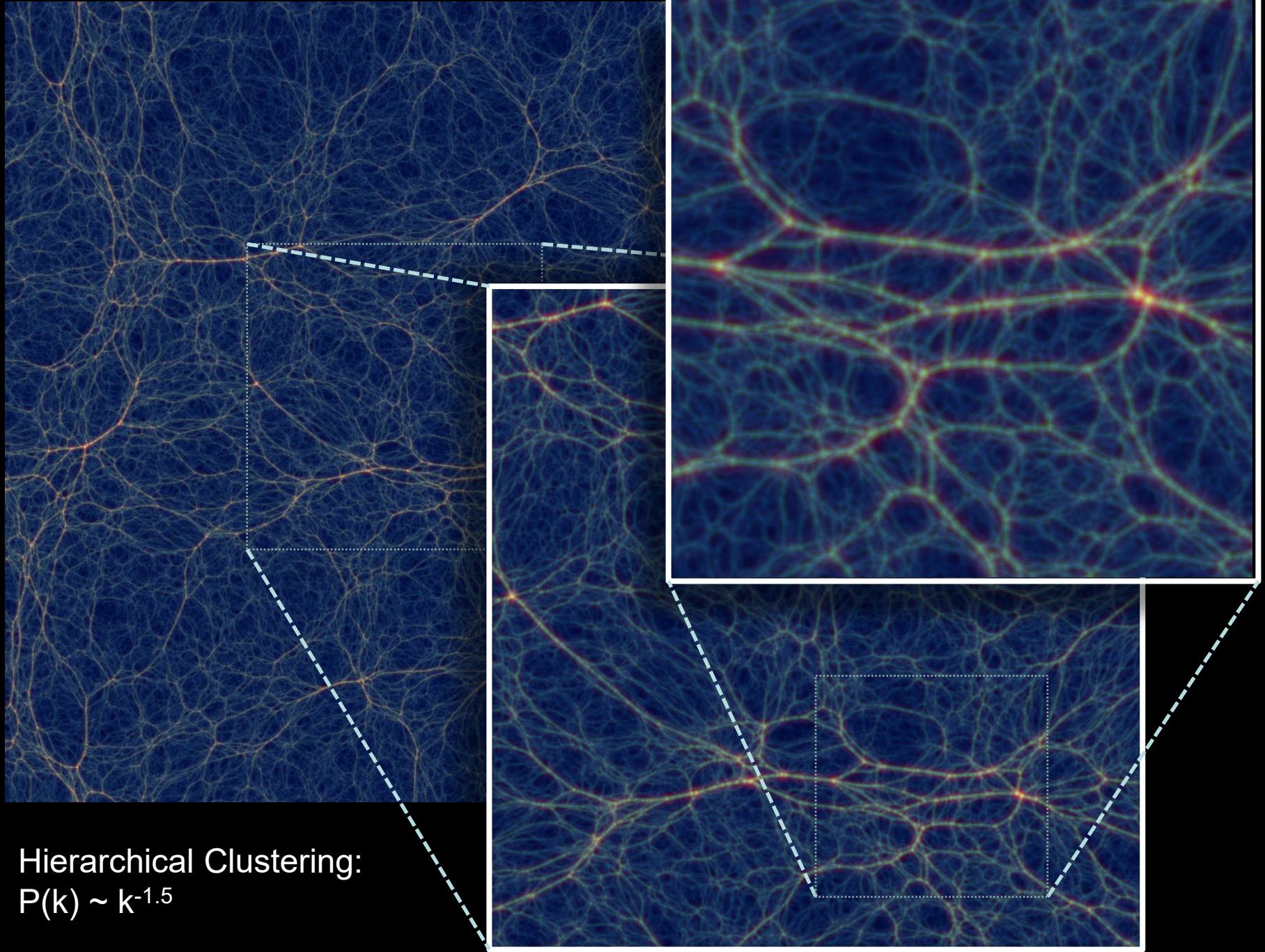
- **anisotropic structure:**
 - filaments dominant structural feature
 - elongated
 - flattened
 - sheets/walls
- **multiscale nature**
 - structure on wide range of scales
 - structures have wide range of densities
- **overdense-underdense asymmetry**
 - voids: underdense, large & roundish
 - filaments & walls: overdense, flattened/elongated
 - clusters: dense, massive & compact nodes
- **complex spatial connectivity**
 - all structural features connected in a complex, multiscale weblike network

Multiscale Cosmic Web

MMF/Nexus+ tracing of filaments

inherent multiscale
character of filamentary web

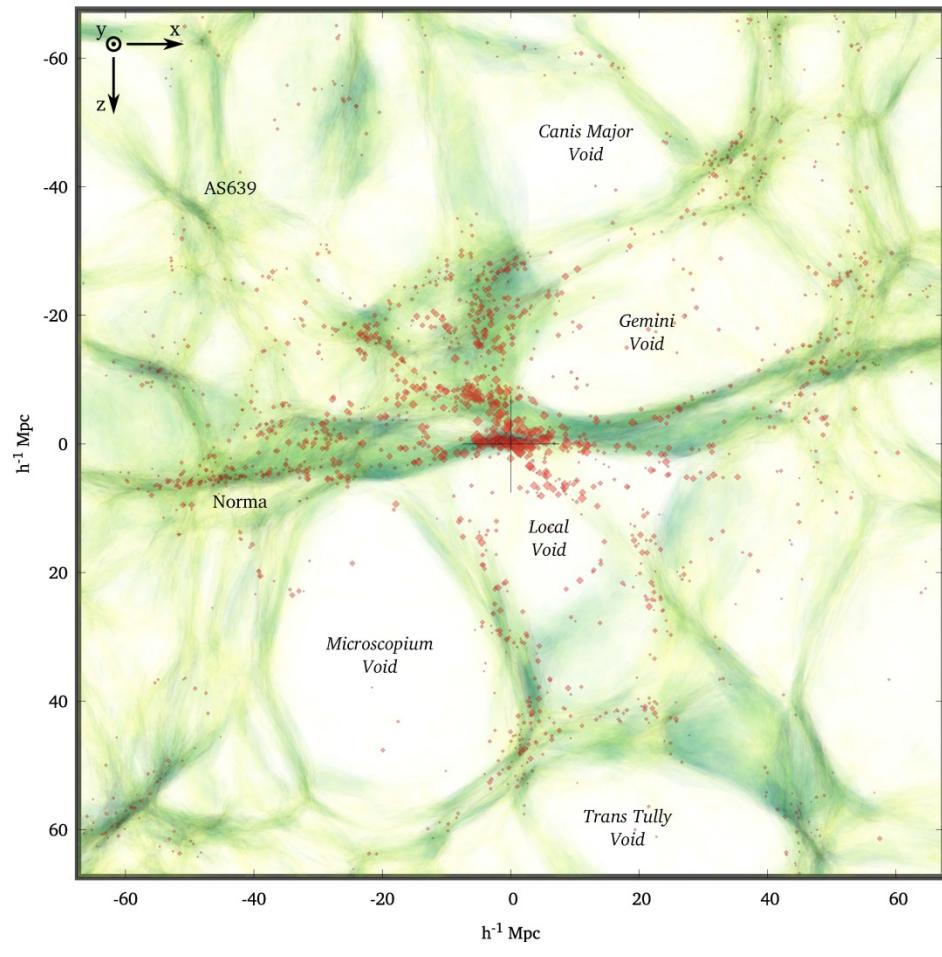
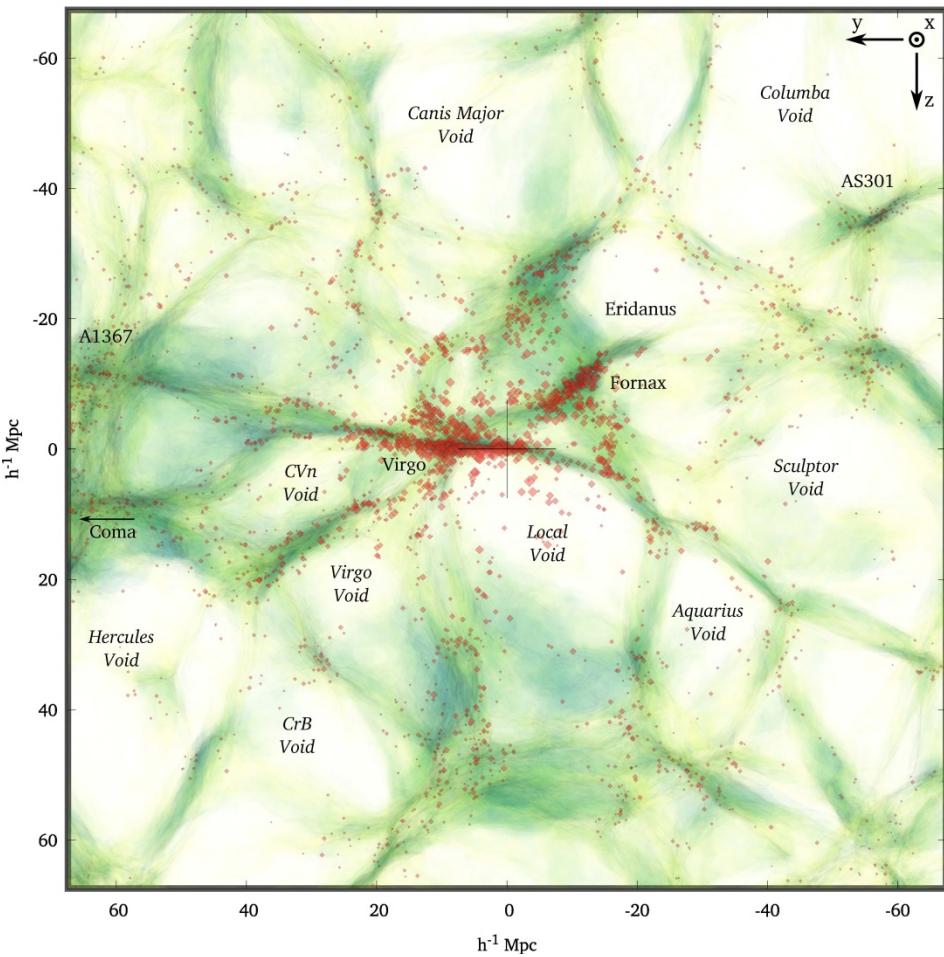
Hidding, Cautun, vdW et al. 2018



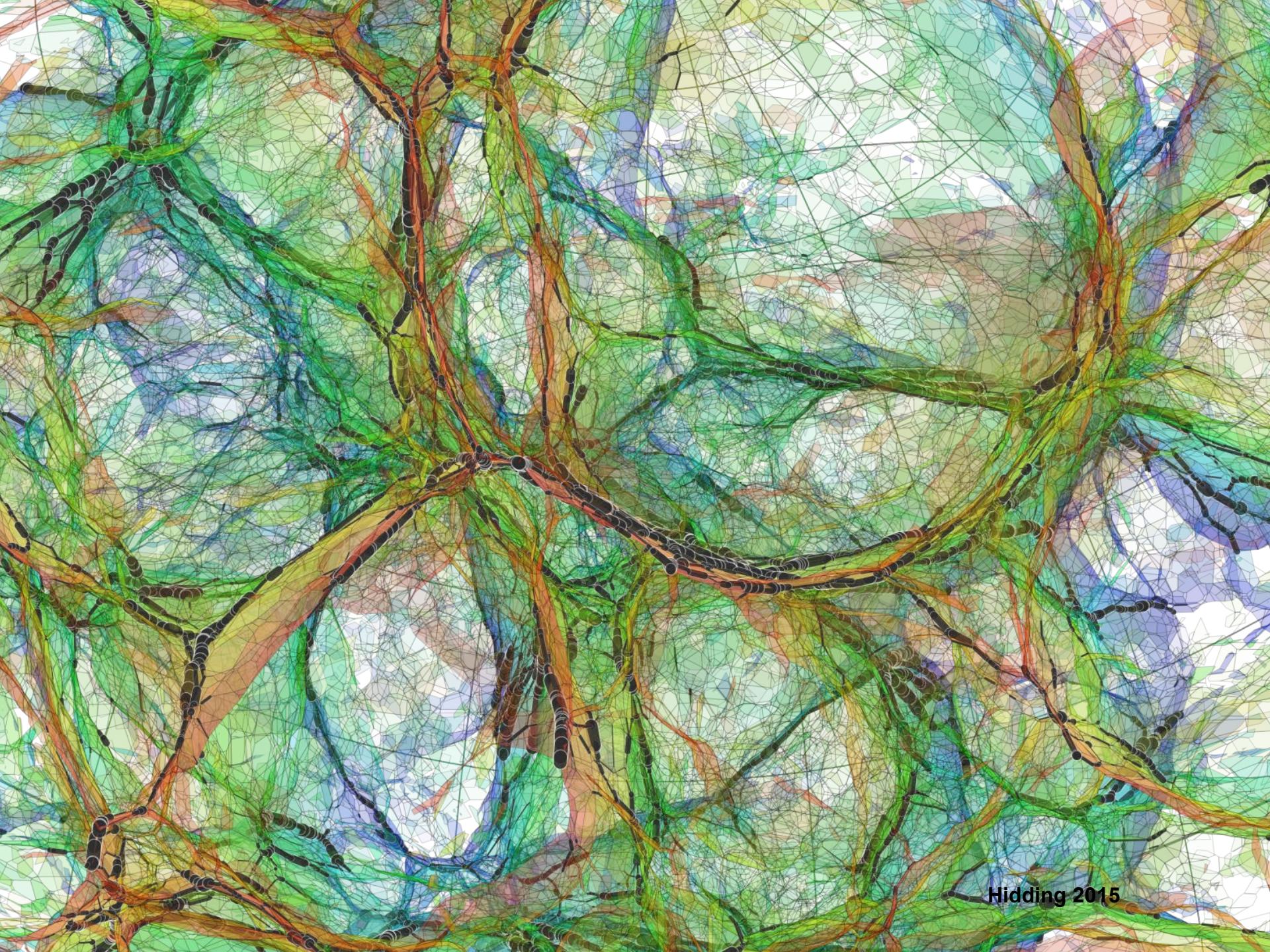
Void Population Local Universe

mean KIGEN-adhesion reconstruction (2MRS)

Hidding, Kitaura, vdW & Hess 2016/2017



Cosmic Web Connectivity



Hidding 2015

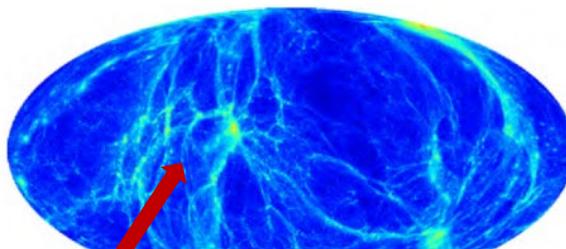
CGV: on walls & filaments

- Mollweide sky projection matter distribution around CGV halos
- CGV halos embedded in walls
- Walls dominate void infrastructure
- substantial fraction in filaments (embedded in walls)
- active dynamical evolution of wall-filament goes along with active void galaxy halo evolution

merging system of
Intravoid walls

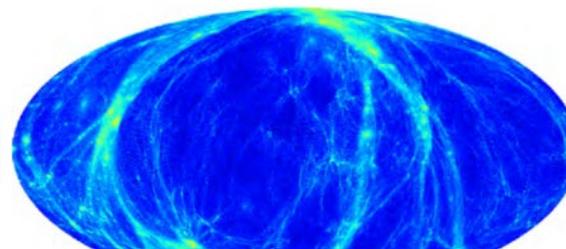
Rieder et al. 2013

CGV_D



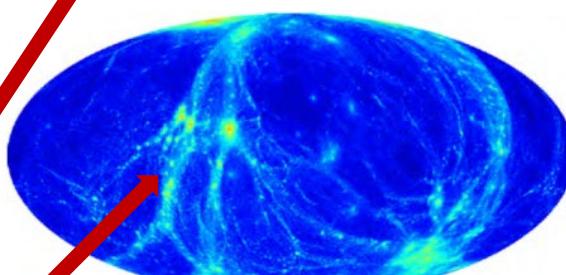
(a) CGV-D_a, $z = 3.7$

CGV_G

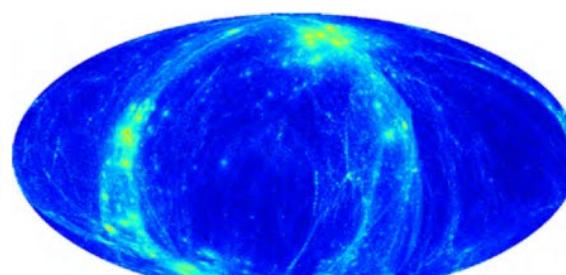


(b) CGV-G_a, $z = 3.7$

$z = 3.7$

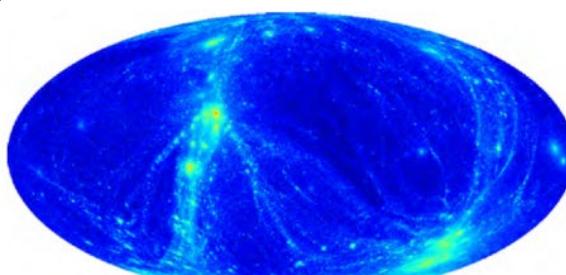


(c) $z = 1.6$

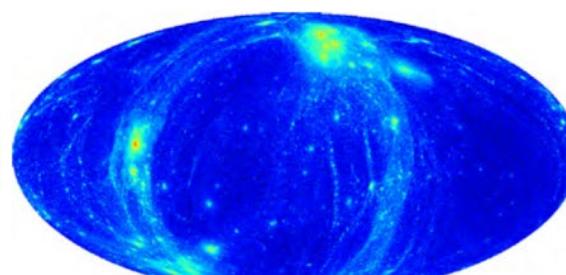


(d) $z = 1.6$

$z = 1.6$

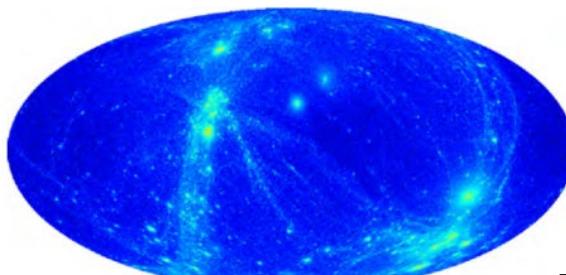


(e) $z = 0.55$

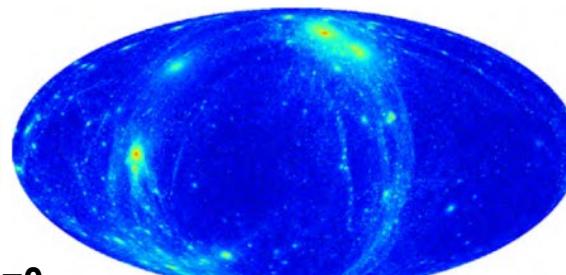


(f) $z = 0.55$

$z = 0.55$



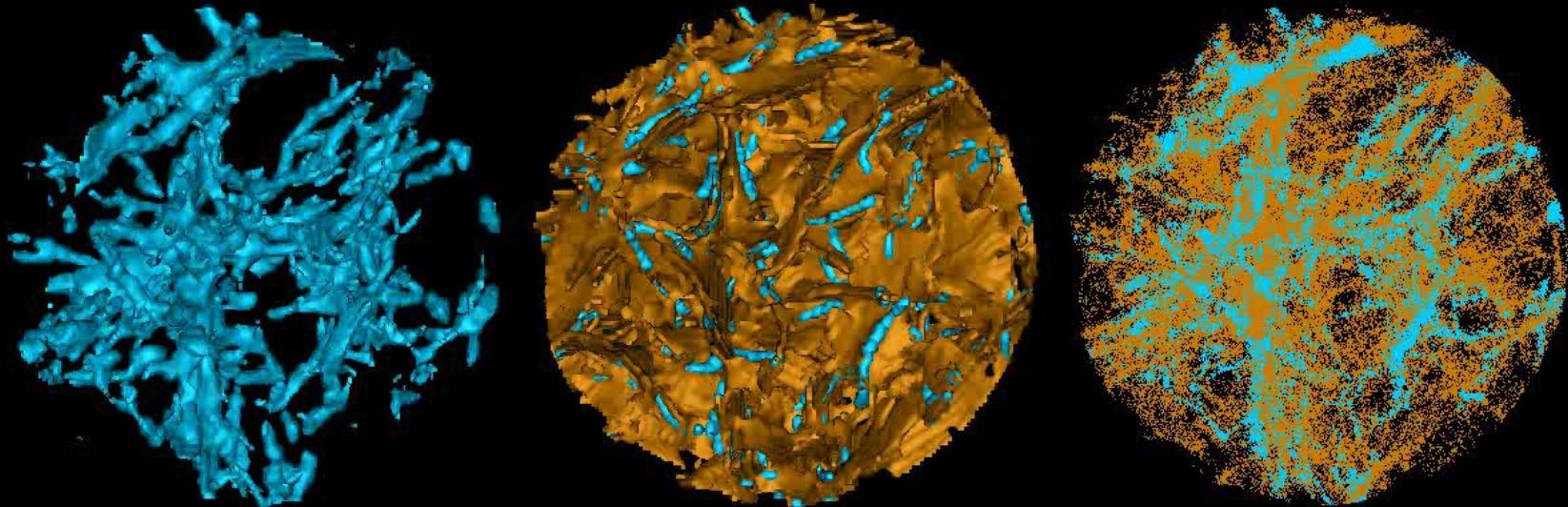
(g) $z = 0$



(h) $z = 0$

$z = 0$

The Cosmic Web



MMF/Nexus
Cautun et al. 2013, 2014

Stochastic Spatial Pattern

- Clusters,
 - Filaments &
 - Walls
- around
- Voids

in which matter & galaxies

have agglomerated

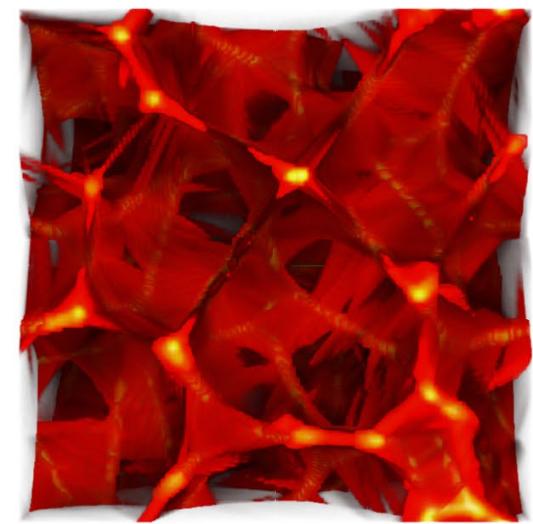
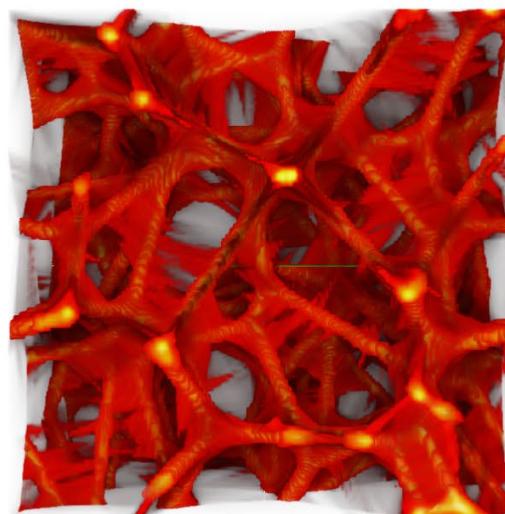
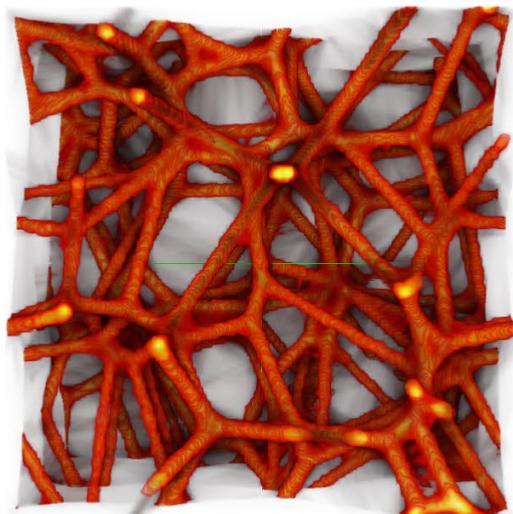
through gravity

Multiscale Topology & Filtrations

Topological Hierarchy: Excursion Sets & Filtrations

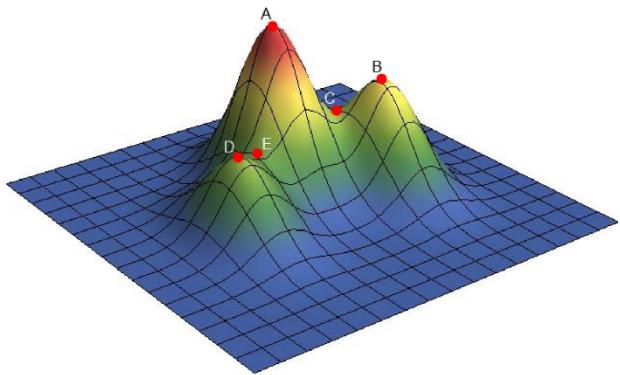
Superlevel Sets

$$\begin{aligned}\mathfrak{M}_v &= \left\{ \vec{x} \in \mathfrak{M} \mid f_s(\vec{x}) \in [f_v, \infty) \right\} \\ &= f_s^{-1}[f_v, \infty)\end{aligned}$$

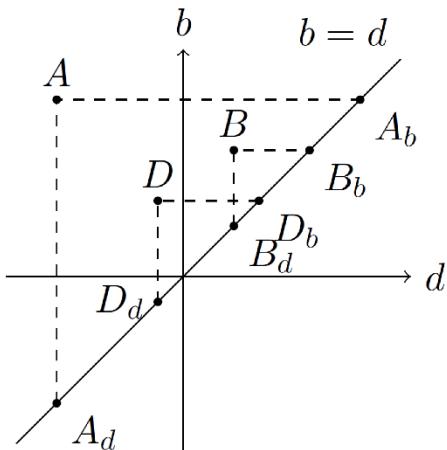


Pranav et al. 2013a

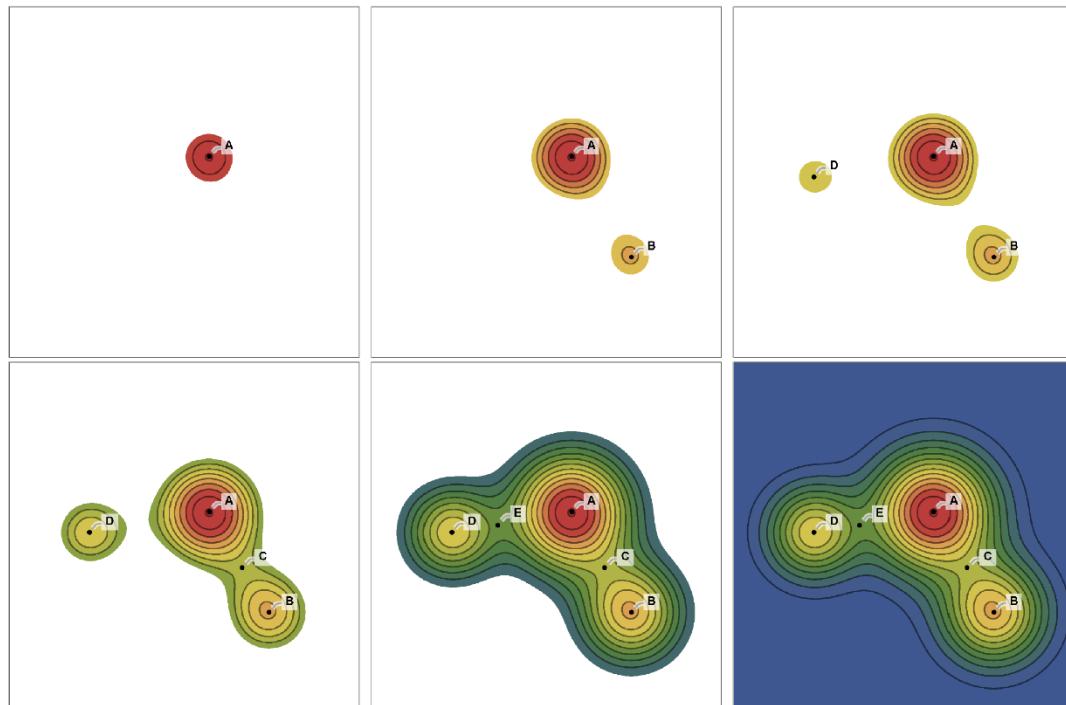
Topological Hierarchy: Persistence



(a) Realization of random field.



(b) Sketch of persistence diagram.



Persistence Diagram

Hierarchy
Field Singularities

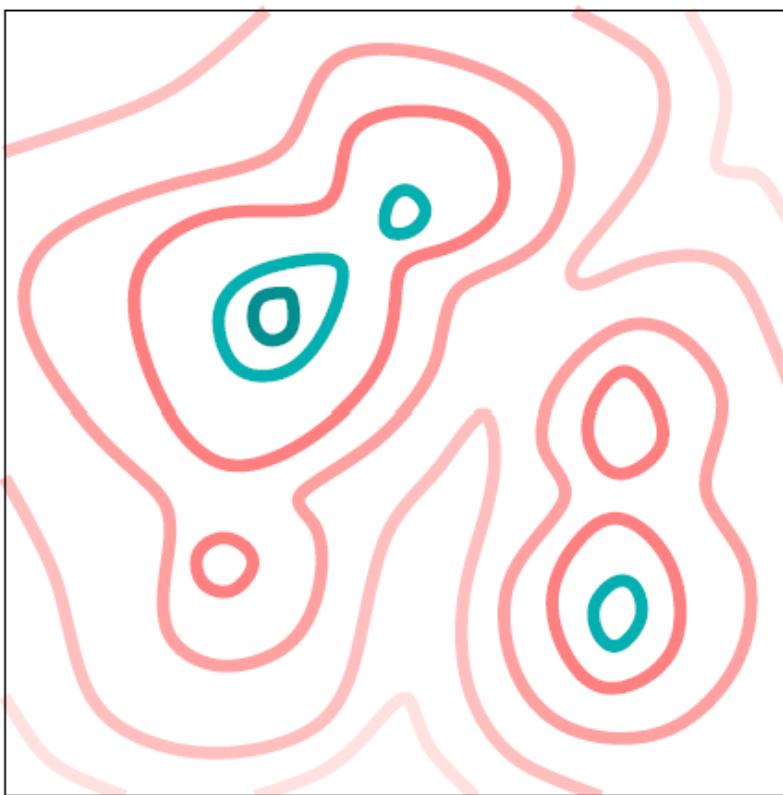
Persistent Pairs:

- Birth level singularity
- Death level singularity

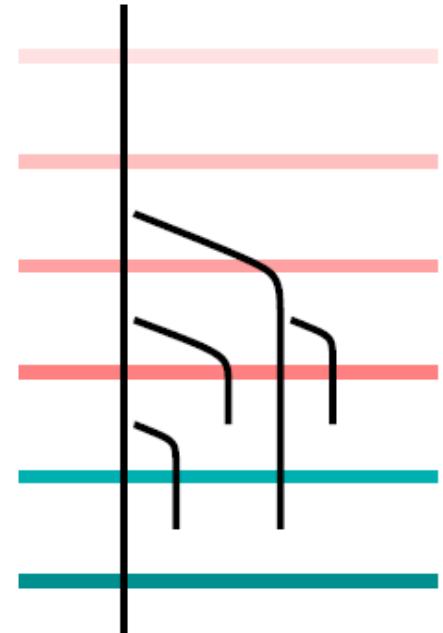
Feldbrugge, van Engelen,
vdW et al. 2020

Topological Hierarchy

Persistent Homology:
“Cycling” over field filtration



Topology Tree



Edelsbrunner & Harer 2010

field value filtration
tree hierarchy

Topology & Geometry

Geometry & Topology

- Conventional Cosmological Topology Measure:

(Reduced) Genus

- # holes - # connected regions
- (Gott et al. 1986; Hamilton et al. 1986; Choi et al. 2010)

- Complete quantitative characterization of local geometry in terms of

Minkowski Functionals

- Minkowski Functionals:
 - Volume
 - Surface area
 - Integrated mean curvature
 - Genus/Euler Characteristic
- (Mecke, Buchert & Wagner 1994)

Genus, Euler & Betti

- For a surface with c components, the genus G specifies # handles on surface, and is related to the Euler characteristic χ via:

$$G = c - \frac{1}{2} \chi(\partial M)$$

where, according to the Gauss-Bonnet theorem

$$\chi(\partial M) = \frac{1}{2\pi} \oint \left(\frac{1}{R_1 R_2} \right) dS$$

- Euler characteristic 3-D manifold & 2-D boundary manifold:

$$\chi(M) = \frac{1}{2} \chi(\partial M)$$



$$\chi(\partial M) = 2(\beta_0 - \beta_1 + \beta_2)$$

the Rule of Euler

from: Robert Adler

SIMPLICIAL TOPOLOGY

Simplices, complexes,
cycles, numbers of simplices,
Betti numbers

$$\sum_k (-1)^k \# \{k\text{-dimensional simplices}\}$$

$$\sum_k (-1)^k \beta_k$$

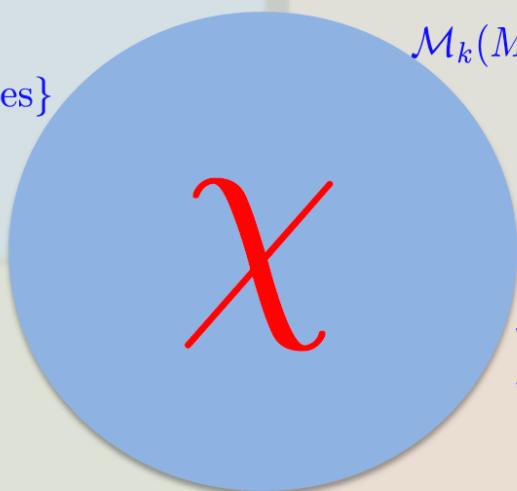
ALGEBRAIC TOPOLOGY

Homology, homotopy,
dimensions of groups,
Betti numbers, persistence

INTEGRAL GEOMETRY

Convexity, convex ring
kinematic formulae
Minkowski functionals

$$\mathcal{M}_k(M) = c_{dk} \int_{\text{Graff}(d,d-k)} \chi(M \cap V) d\mu_{d-k}^d(V)$$



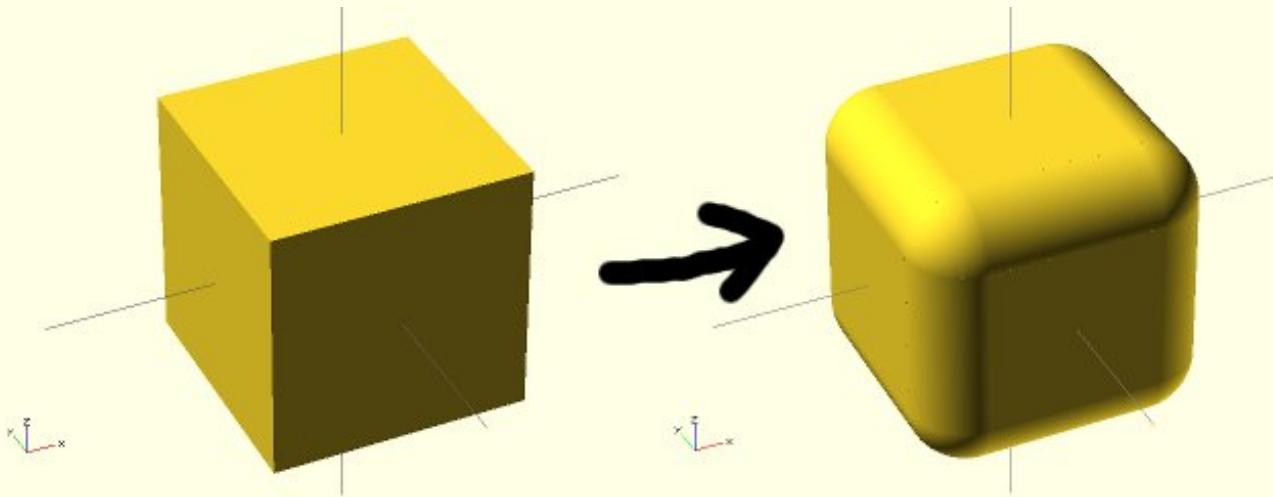
$$\sum_k (-1)^k \# \{\text{critical points of index } k\}$$

$$\int_M \text{Tr}(R^{m/2}) \text{Vol}_g$$

DIFFERENTIAL TOPOLOGY

Curvature, forms, Betti numbers,
Morse theory, integration,
Lipschitz-Killing curvatures

Minkowski functionals



- Weyl's Tube formula:

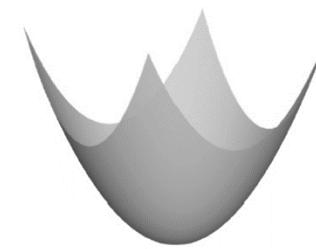
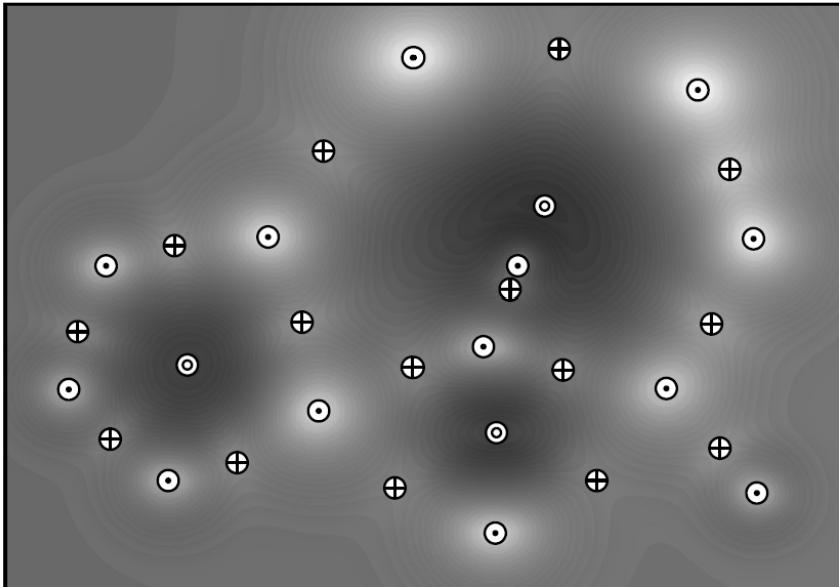
Minkowski functionals Q_k are the parameters specifying the contribution of volumes r^k to the volume of a cube M^r with rounded edges of radius r :

$$\text{Vol}(M^r) = Q_0 + Q_1 r + Q_2 r^2 + Q_3 r^3$$

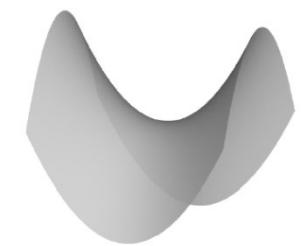
Euler & Morse

Number of singularities in field determines Euler characteristic:

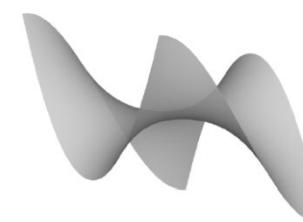
- ζ_0 : **minima**
- ζ_1 : **saddle 1**
- ζ_2 : **saddle 2**
- ζ_3 : **maxima**



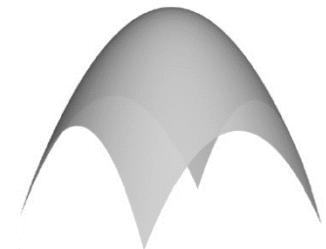
(a) Minimum, 0, \odot



(b) Saddle, 1, \oplus



(d) Monkey Saddle, \circledast



(c) Maximum, 2, \odot

$$\chi = \sum_{k=0}^d (-1)^k \zeta_k$$

Euler-Poincare

Euler Characteristic χ is alternating sum of Betti Numbers

3-manifold ??:

$$\begin{aligned}\chi(M) &= \beta_0 - \beta_1 + \beta_2 + \beta_3 \\ &\approx \beta_0 - \beta_1 + \beta_2\end{aligned}$$

boundary 2-manifold ??:

$$\chi(\partial M) = \beta_{0b} - \beta_{1b} + \beta_{2b}$$

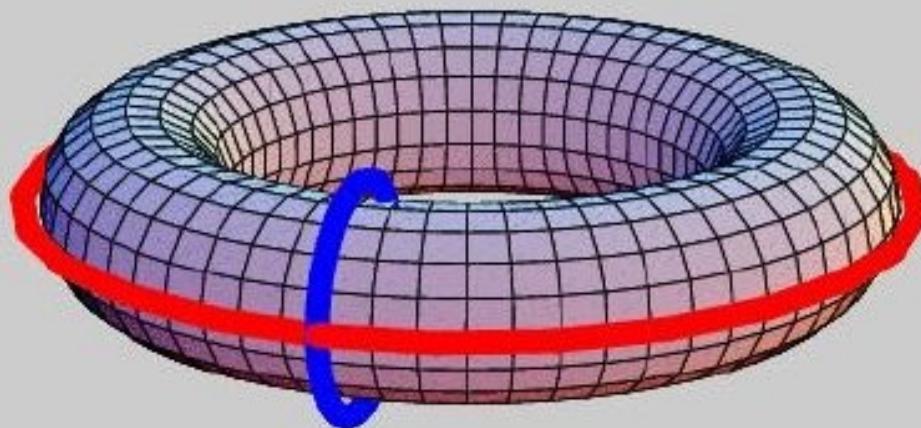
Betti Numbers - Homology Groups

Topology:

Study of connectivity and spatial relations that remain invariant under homeomorphisms (= continuous mapping between two topological objects)

Homology:

Description of topology of a space in terms of the relationship between cycles and boundaries.



Torus: one 0-cycle: rank group H_0 : 1 :
two 1-cycles: rank group H_1 : 2
one 2-cycle rank group H_2 : 1

p-chain: sum of p-simplices
p-cycle: boundary of (p+1) chain

0-cycle: closed component
1-cycle: closed loop of edges,
or finite union
2-cycle: closed surface,
or finite union

adding two p-cycles \longrightarrow p-cycle



Group of p-cycles:

Betti Numbers - Cosmological View

Homology:

focus on structural features & holes



relating to surrounding spaces/structures

Holes different Dimension

0-dim: connected components
“islands” connected mass clumps

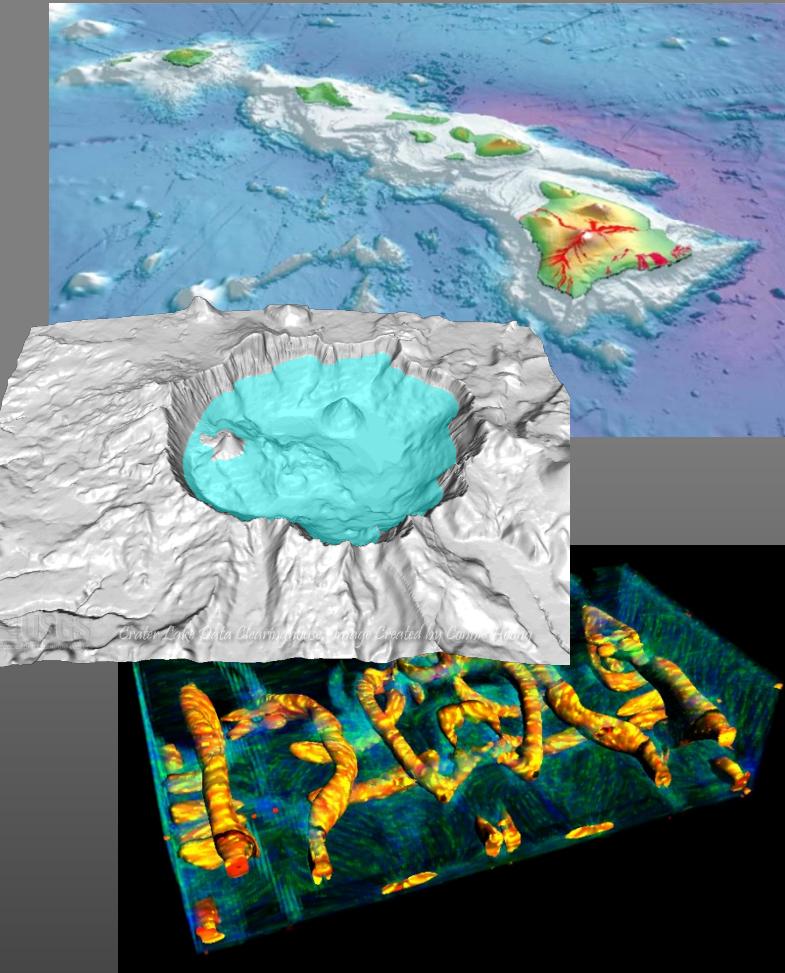
Superclusters

1-dim: Loops

Filaments & Tunnels

2-dim: Closed surface,

Voids



Betti Numbers:

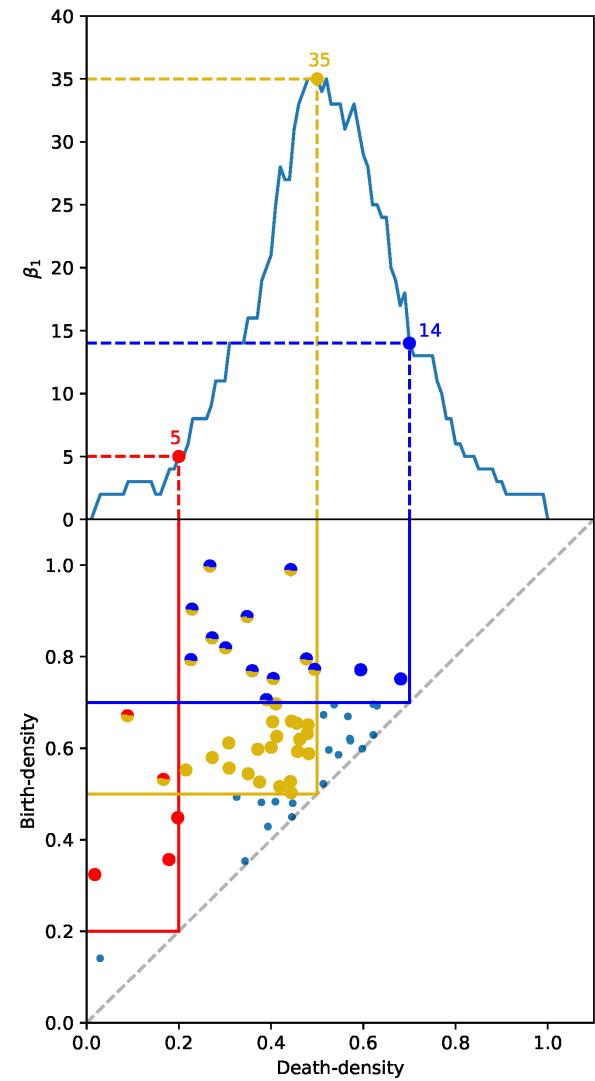
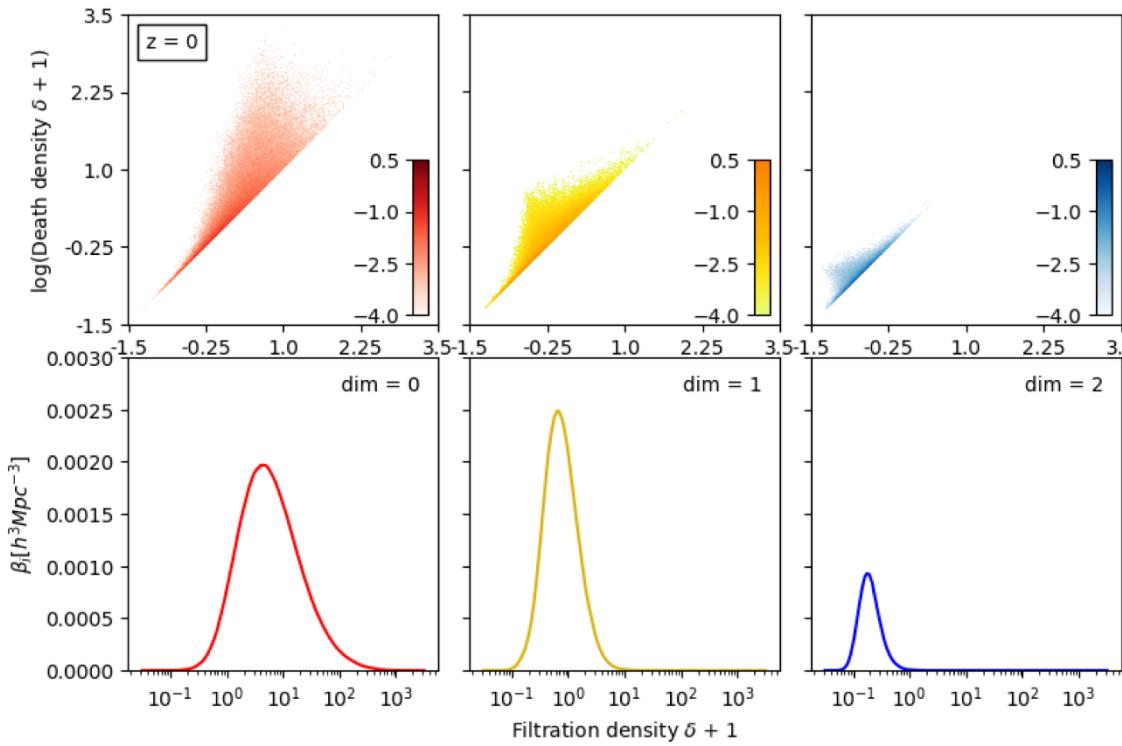
d-dimensional features

Betti Numbers - Persistence

Betti Number $\beta_i(v)$:

Integral over persistence pairs:

- **birth level** $v > v_{\text{birth}}$
- **death level** $v < v_{\text{death}}$



Reionization Bubble Network:

Multiscale Topology

100 seconds

380 000 years

300–500 million years



Light and matter are coupled

Dark matter evolves independently: it starts clumping and forming a web of structures

Light and matter separate

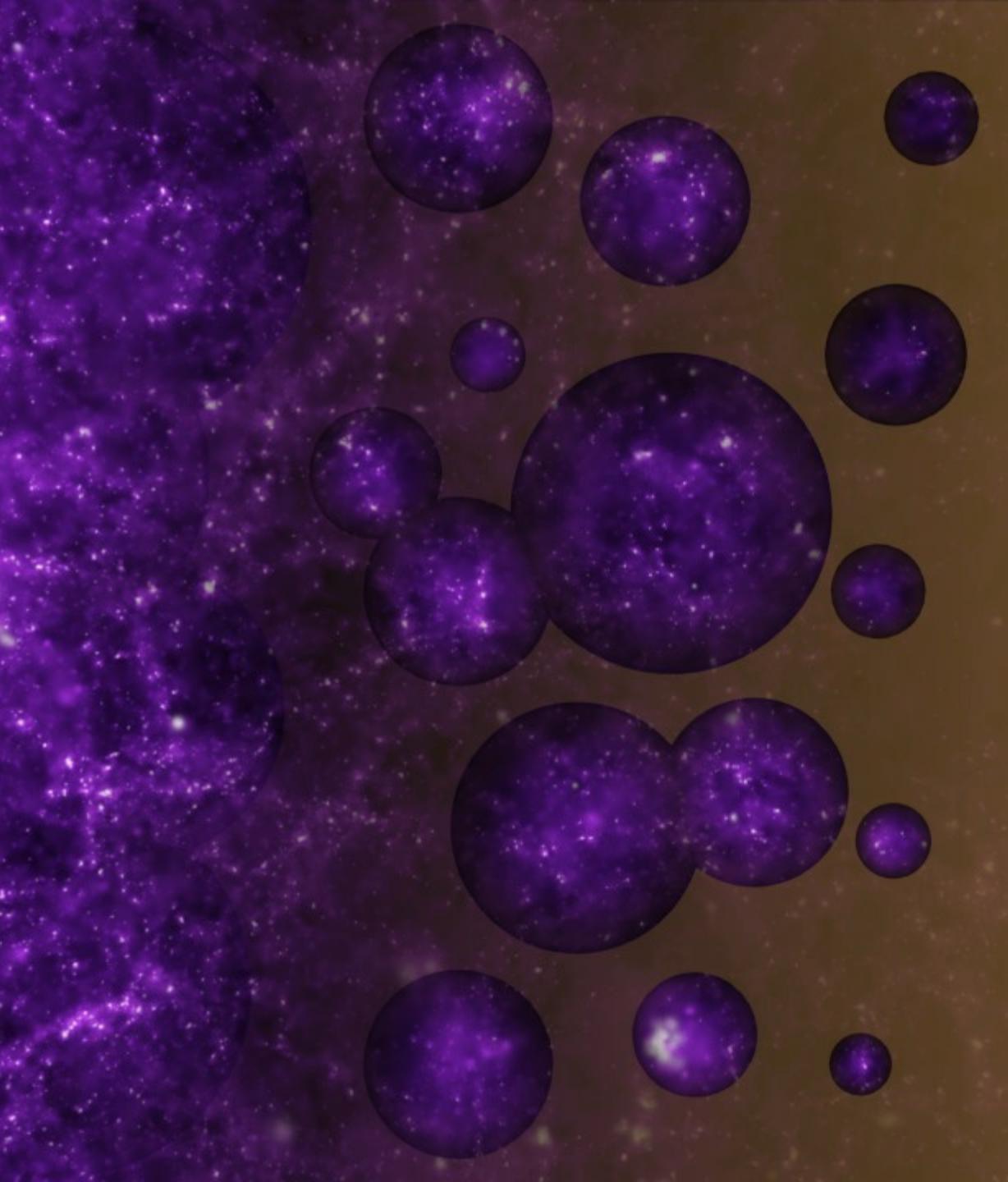
- Protons and electrons form atoms
- Light starts travelling freely: it will become the Cosmic Microwave Background (CMB)

Dark ages

Atoms start feeling the gravity of the cosmic web of dark matter

First stars

The first stars and galaxies form in the densest knots of the cosmic web



the intense UV radiation
of the first stars and galaxies

and possibly
the hard Xray radiation
of the first
Supermassive Black Holes

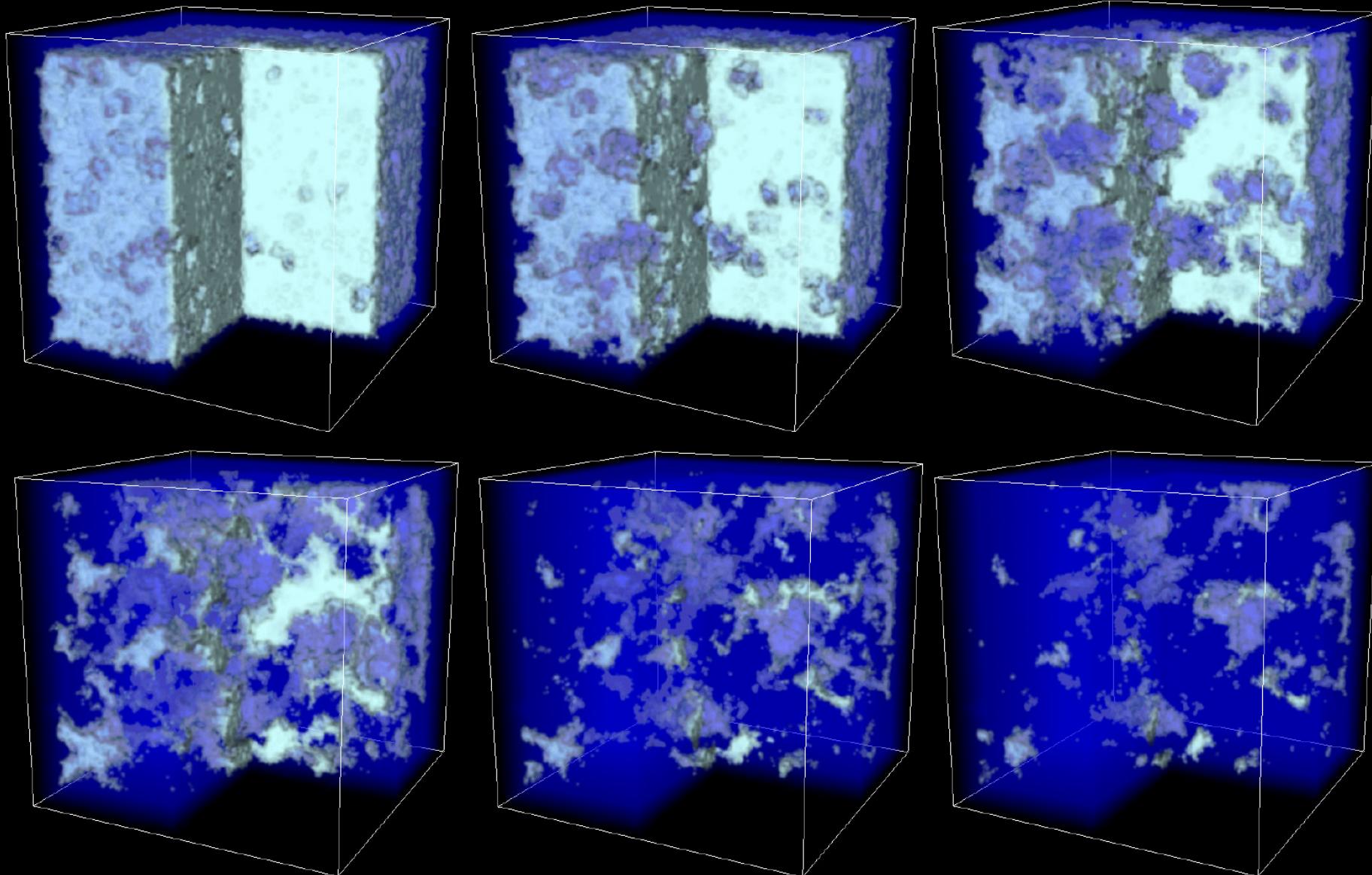
start to ionize the
neutral hydrogen in the
intergalactic medium

the IGM is mostly concentrated
in the filaments of the early Cosmic Web,
with still substantial amounts in the voids

the ionization bubbles around
these early objects expand and
gradually merge and occupy
major fractions of the Universe:

EPOCH OF REIONIZATION (EOR)

Reionization Process

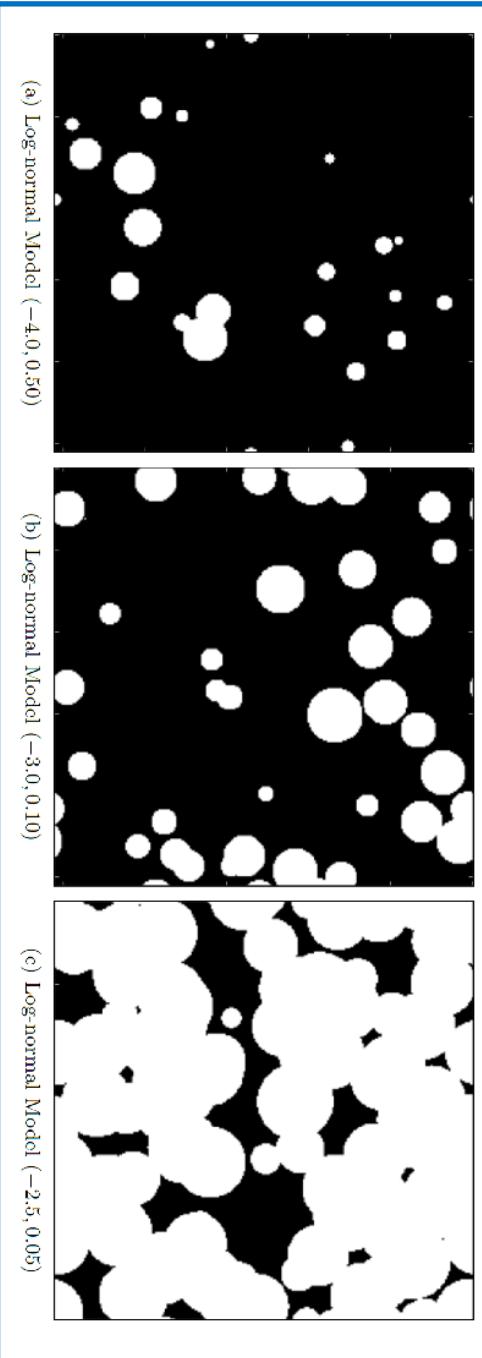
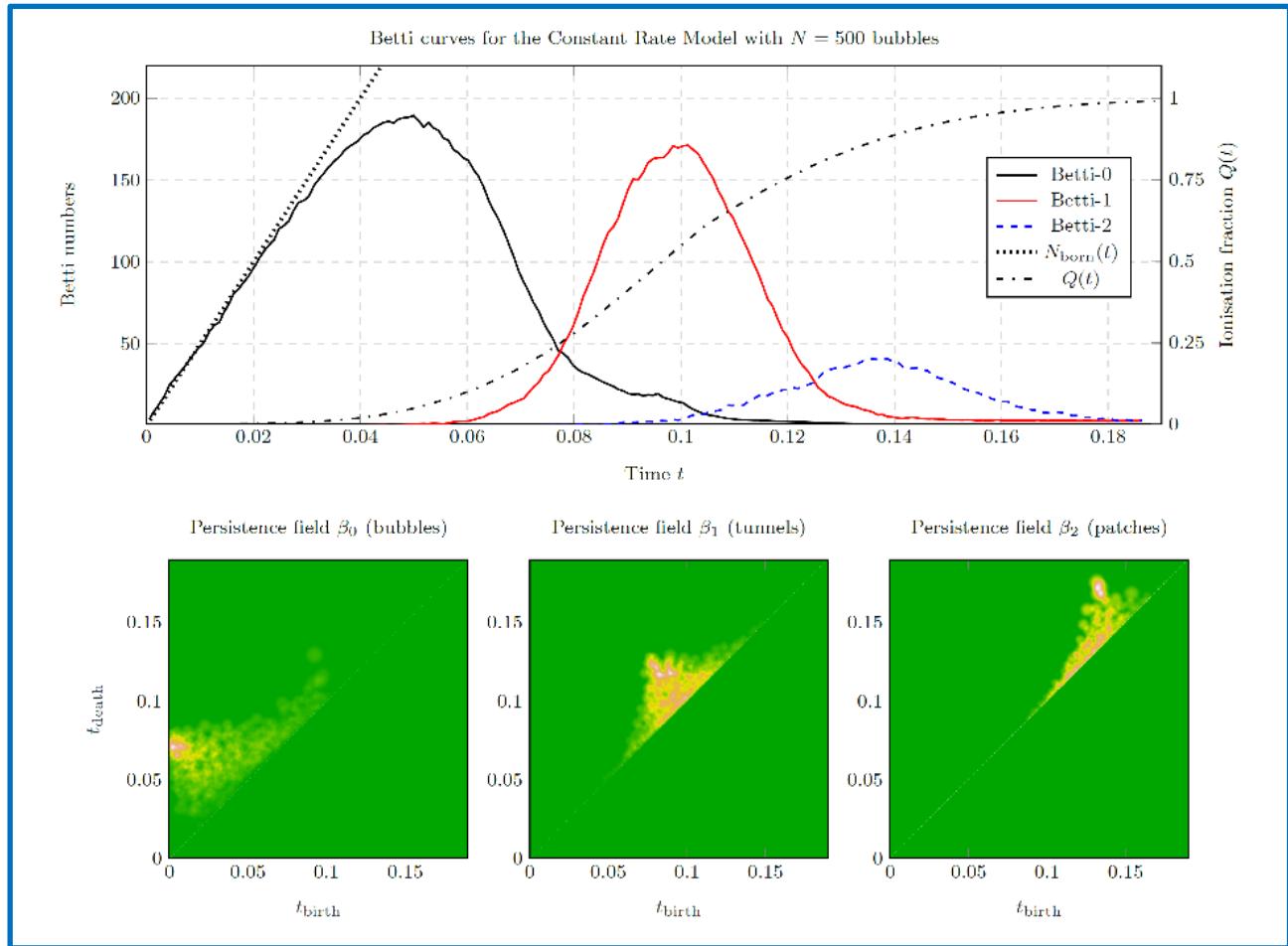


Cosmic Bubble Bath

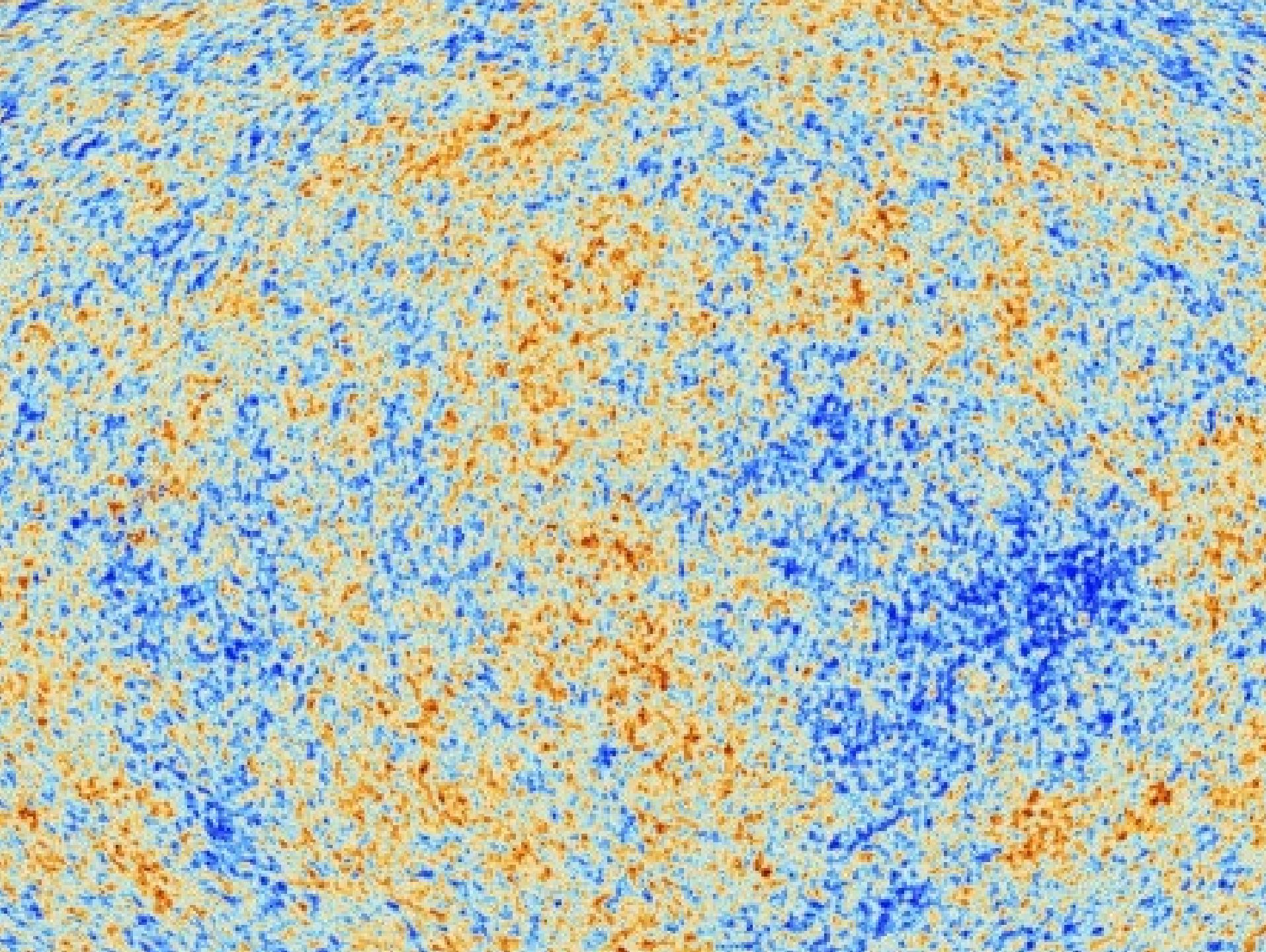
M. Alvarez, T Abel.

Reionization Bubble Network Persistent Topology

Elbers & vdW 2019, 2021

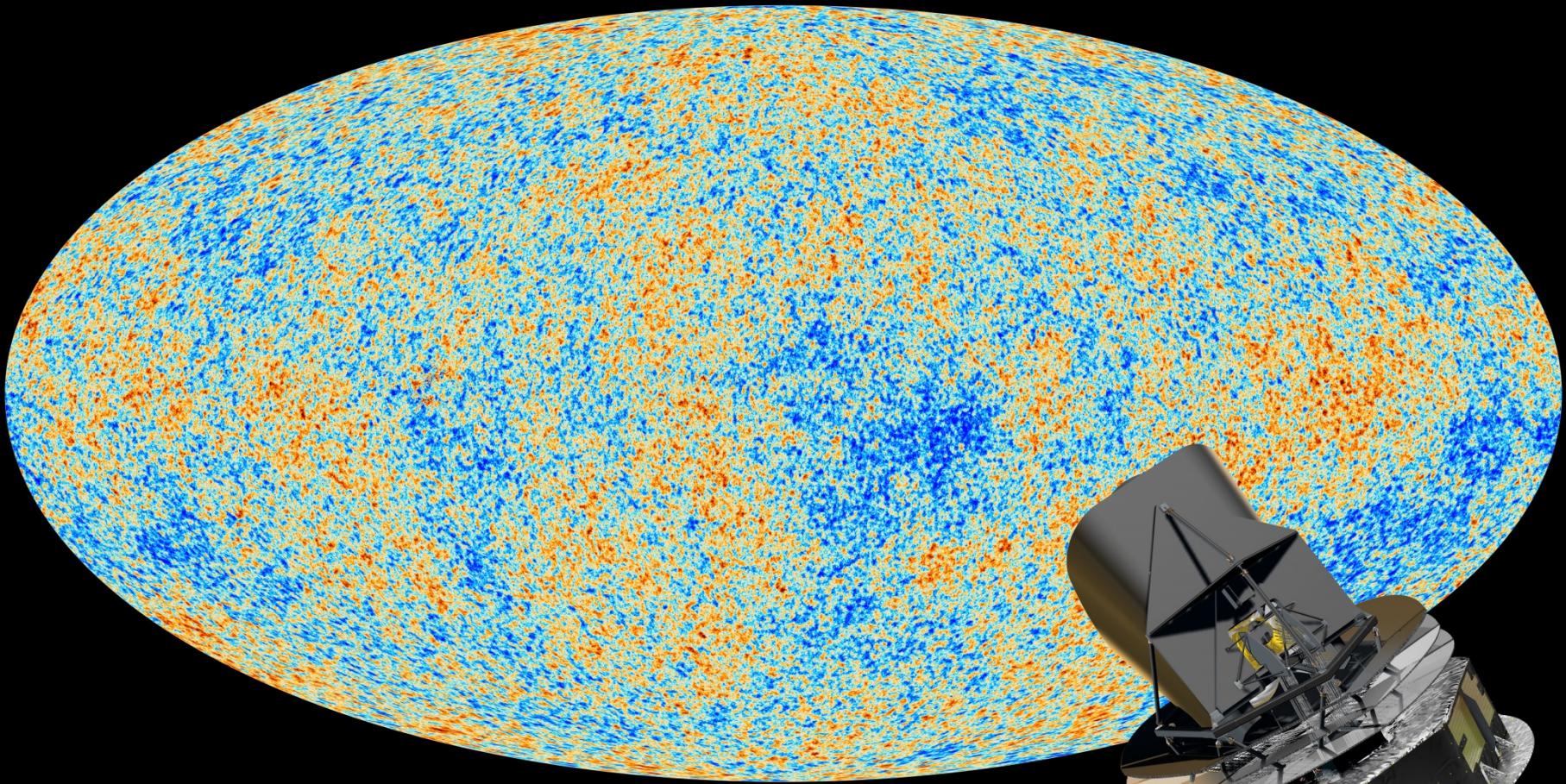


Topology of the Primordial Gaussian Field



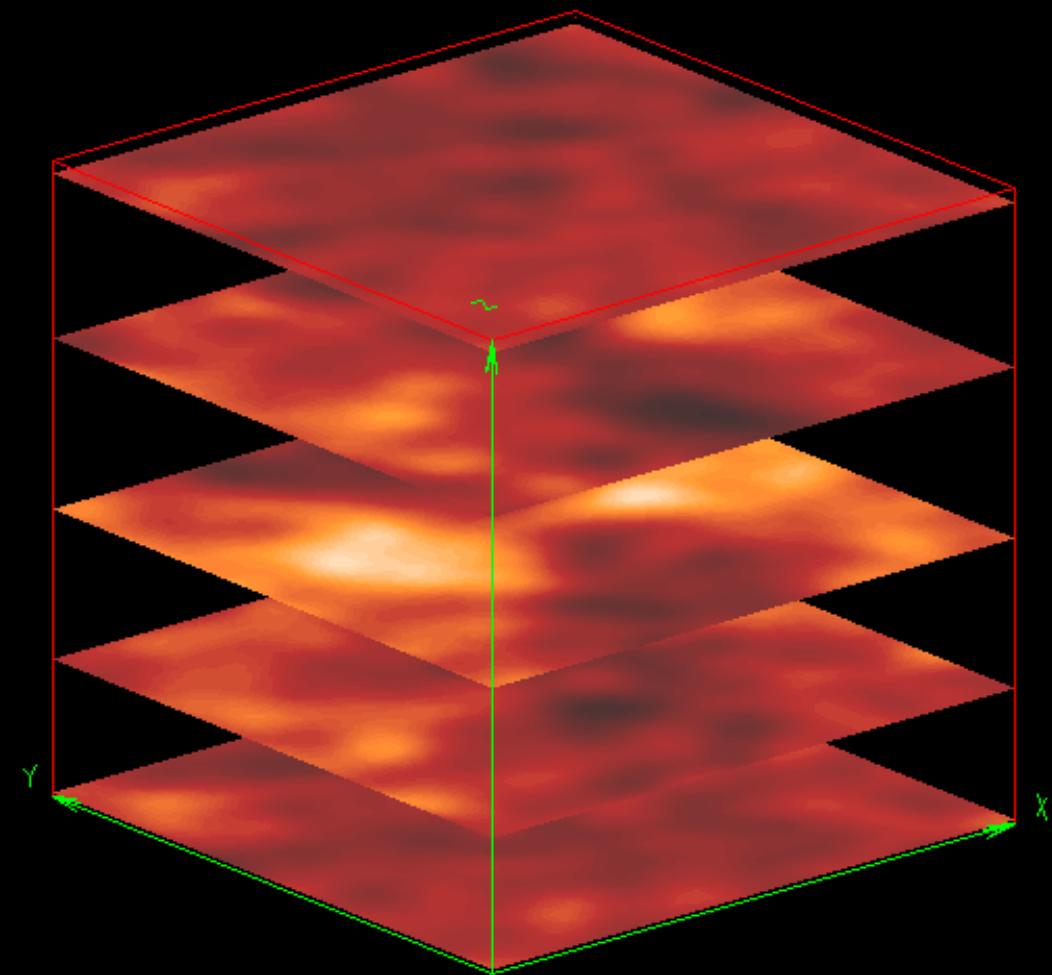
Cosmic Origins

- Universe 380.000 yrs after Big Bang
- 13.8 Gyrs ago (13.798±0.037 Gyrs)
- Temperature T = 2.72548 ± 0.00057 K
- temperature/density fluctuations ($\Delta T/T < 10^{-5}$)



Planck Baby Photo
of our Universe

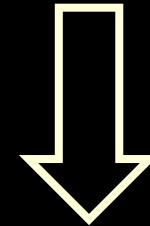
Primordial Gaussian Field



$$f(\vec{x}) = \int \frac{d\vec{k}}{(2\pi)^3} \hat{f}(\vec{k}) e^{-i\vec{k} \cdot \vec{x}}$$

$$\hat{f}(\vec{k}) = \hat{f}_r(\vec{k}) + i \hat{f}_i(\vec{k}) = |\hat{f}(\vec{k})| e^{i\theta(\vec{k})}$$

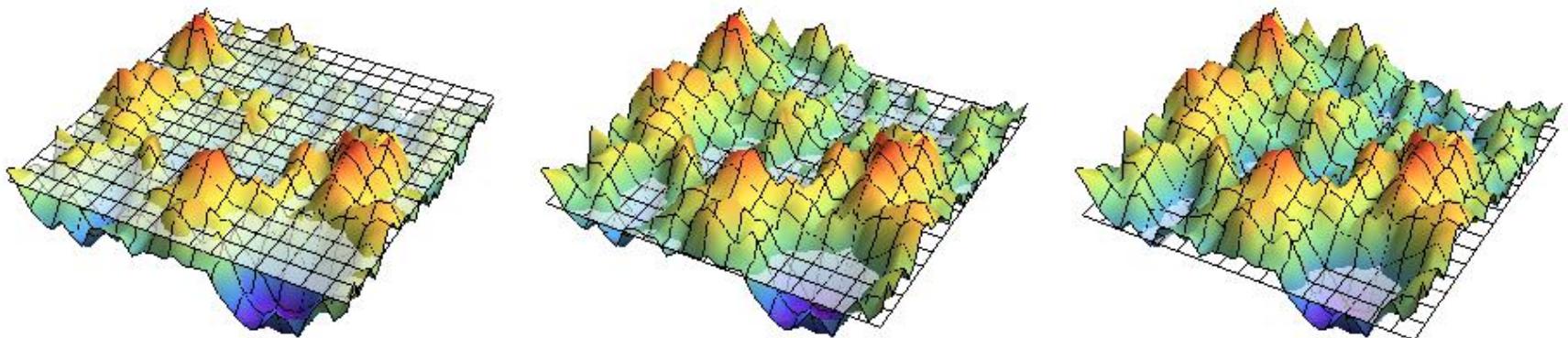
$$P_N = \frac{\exp\left[-\frac{1}{2}\sum_{i=1}^N \sum_{j=1}^N f_i(M^{-1})_{ij} f_j\right]}{\left[(2\pi)^N (\det M)\right]^{1/2}} \prod_{k=1}^N df_k$$



$$P_N \propto \exp\left(-\sum_i \frac{|\hat{f}(\vec{k}_i)|^2}{2P(k_i)}\right) \propto \prod_i \exp\left(-\frac{|\hat{f}(\vec{k}_i)|^2}{2P(k_i)}\right)$$

$$P_1(|\hat{f}(\vec{k})|) d|\hat{f}(\vec{k})| = \exp\left(-\frac{|\hat{f}(\vec{k})|^2}{2P(k)}\right) \frac{|\hat{f}(\vec{k})| d|\hat{f}(\vec{k})|}{P(k)}$$

Gaussian Random Fields: Excursion Sets



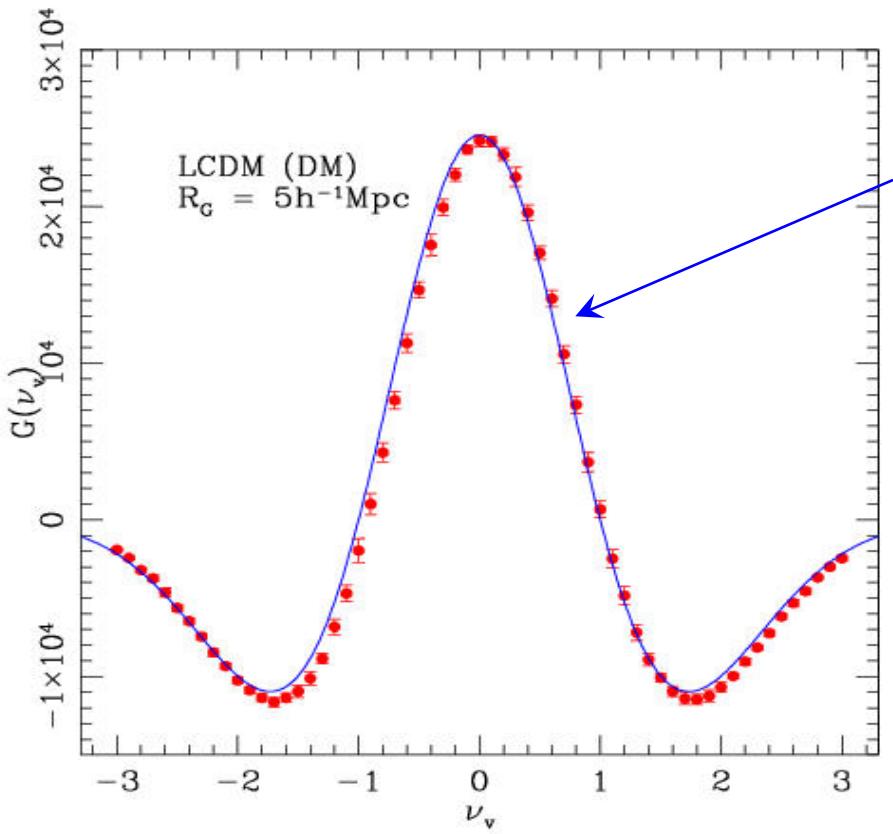
Superlevel Sets

$$\begin{aligned}\mathfrak{M}_\nu &= \left\{ \vec{x} \in \mathfrak{M} \mid f_s(\vec{x}) \in [f_\nu, \infty) \right\} \\ &= f_s^{-1}[f_\nu, \infty)\end{aligned}$$

Gaussian Random Fields: Genus

Genus Gaussian Field, the “cosmological” way :

$$g = G - c$$



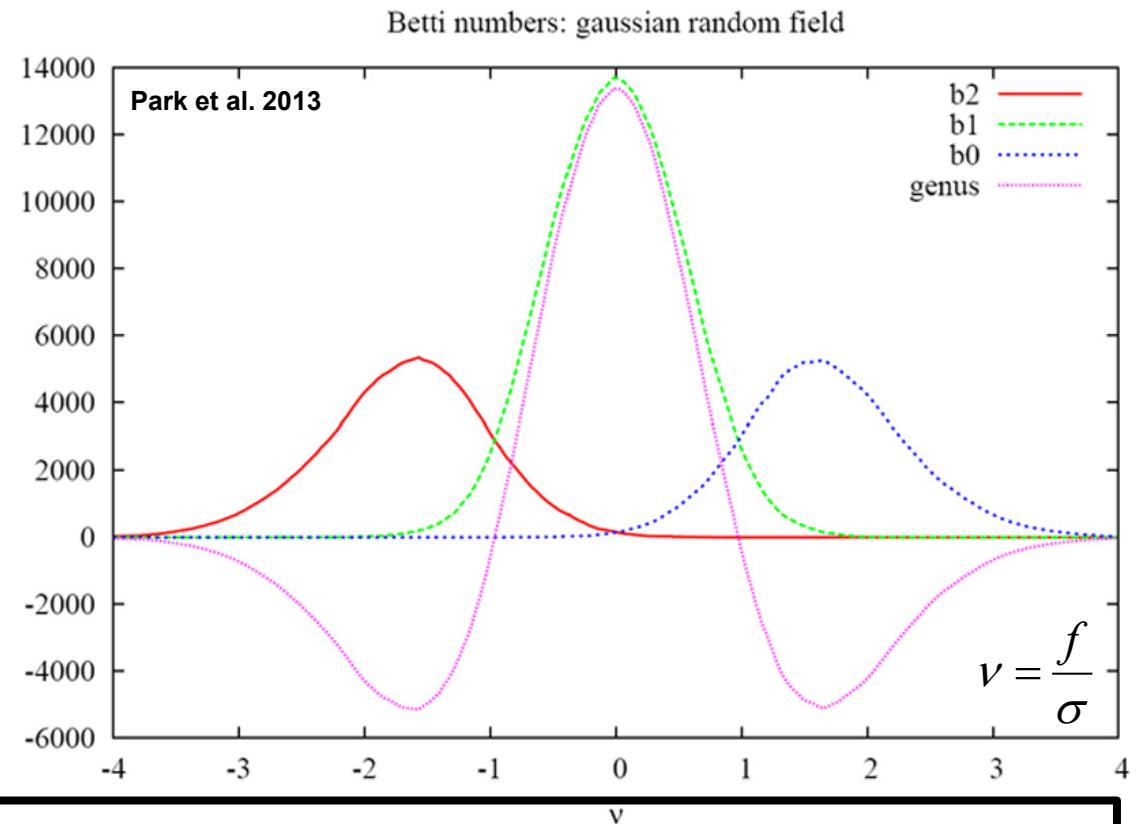
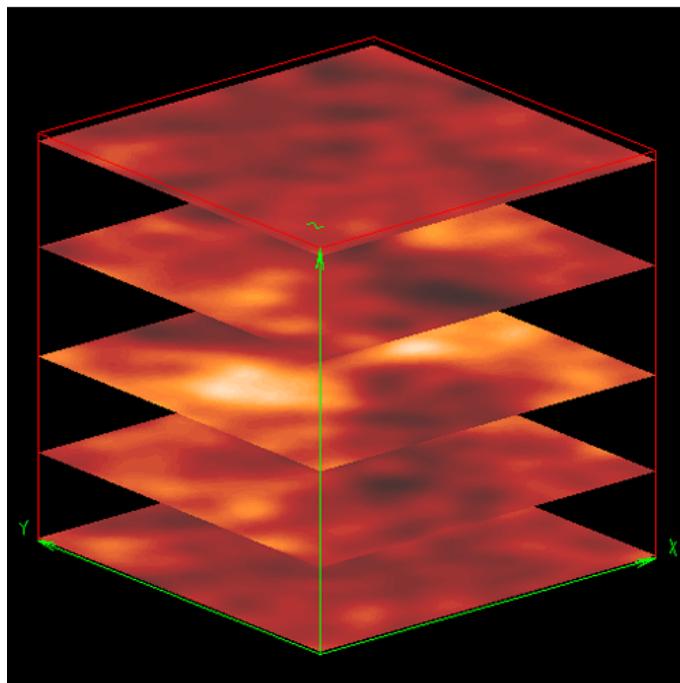
$$g(\nu) = -\frac{1}{8\pi^2} \left(\frac{\langle k^2 \rangle}{3} \right)^{3/2} (1 - \nu^2) e^{-\nu^2/2}$$

$$g(\nu) = -\beta_0(\nu) + \beta_1(\nu) - \beta_2(\nu)$$

Park et al. 2013

Gaussian Random Fields: Betti Numbers

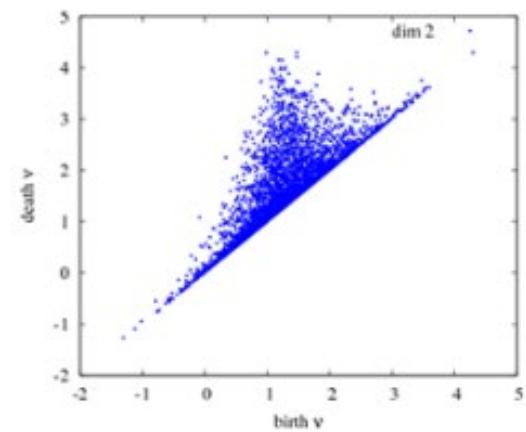
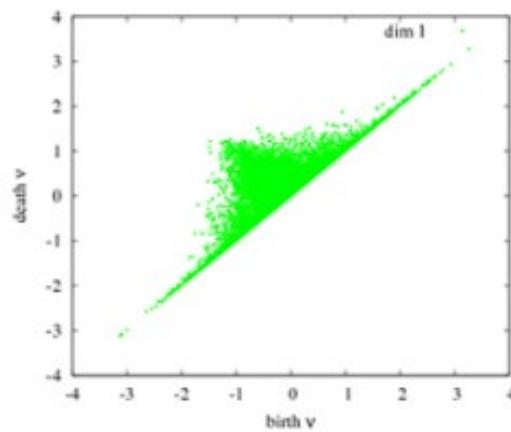
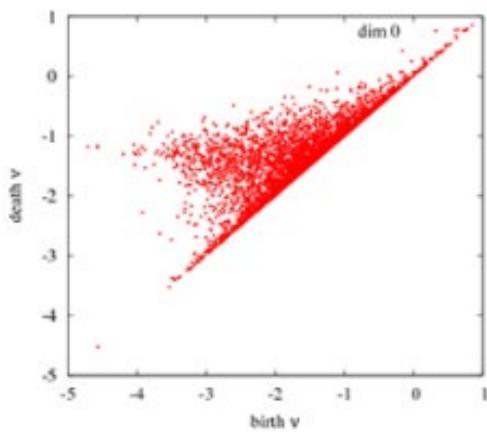
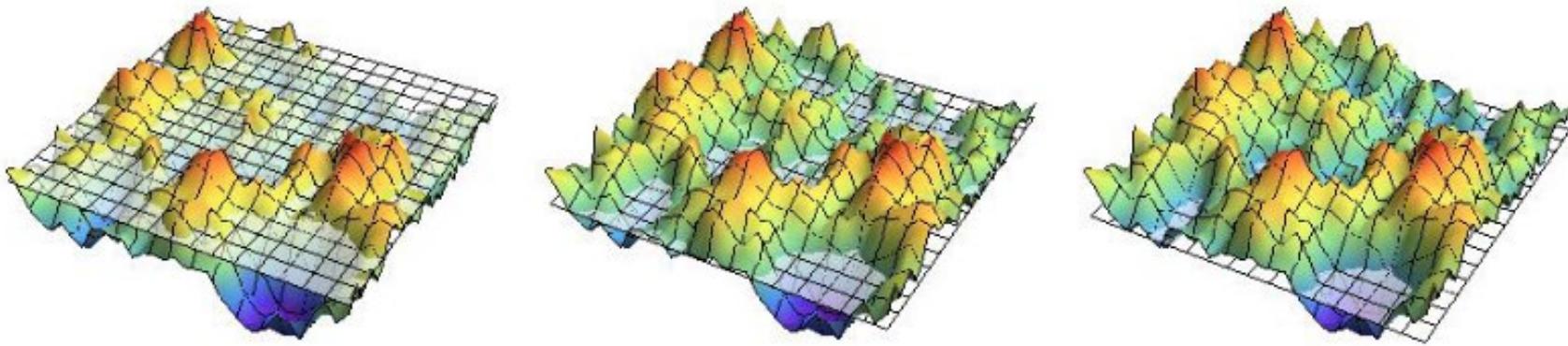
$$\chi = \beta_0 - \beta_1 + \beta_2$$



In a Gaussian field:

- # tunnels dominant at intermediate density levels, when superlevel domain spongelike
- overlap between β_0 and β_2 at $\nu=0$, domain punctured by clumps with cavities
- # clumps/islands reaches maximum at $\nu = \sqrt{3}$, # cavities/voids at $\nu = -\sqrt{3}$

Gaussian Random Fields: Persistence



2D Gaussian Field

Graph towards Betti Numbers

Gaussian Field Singularities

Number density of singularities in Gaussian random field

(BBKS 1986, Pogosyan et al. 2011, Wintraecken)

$$N_i(v) = \iint |\det H| P(f, f_i = 0, J_i) (J_1^2 - J_2) dJ_1 dJ_2$$

$N_0(v)$: number density minima

$N_1(v)$: number density saddles

$N_2(v)$: number density maxima

$P(f, f_i, J_i)$: Gaussian probability

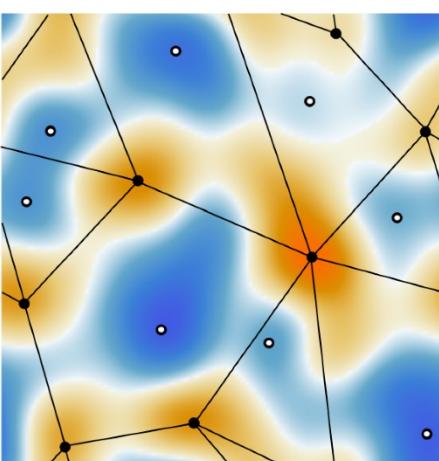
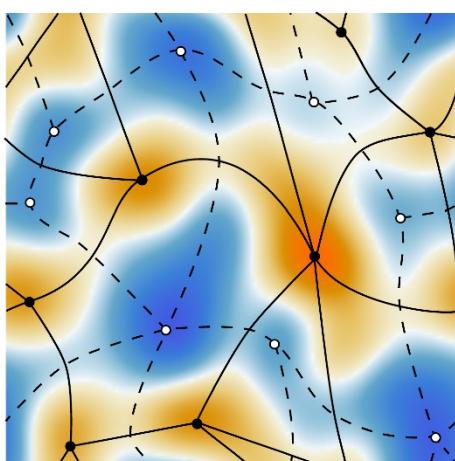
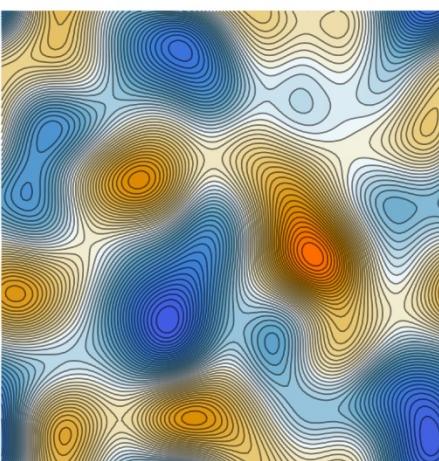
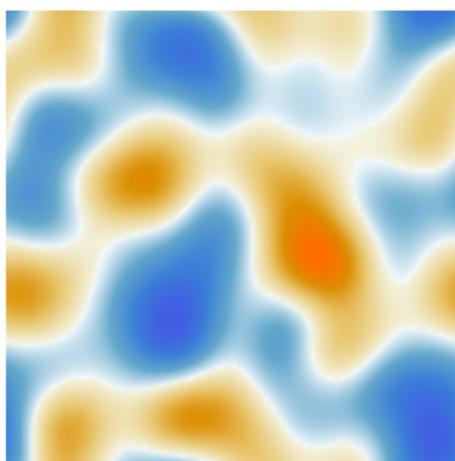
$$f_i = \frac{\partial f}{\partial x_i}; \quad f_1 = f_2 = 0;$$

$$H: \quad f_{ij} = \frac{\partial^2 f}{\partial x_i \partial x_j};$$

$$J_1 = \lambda_1 + \lambda_2 = f_{11} + f_{22}$$

$$J_2 = (\lambda_1 - \lambda_2)^2 = (f_{11} - f_{22})^2 + 4f_{12}^2$$

Gaussian Random Fields: Morse-Smale complex



density field: isolevel set contours

$N_0(v)$, $N_1(v)$, $N_2(v)$:
maxima, saddles, minima in field

Morse-Smale complex



Tessellation of Space

Simplified Simplicial Representation:

Morse Graph

Incremental Algorithm

Computing Betti Numbers:
using ***adapted Morse-Smale complex***, in combination with ***incremental algorithm***.

Critical point	index	β_0	β_1	β_2	addition of
Minimum	0		↓	↑	face
Saddle point	1	↓	↑		edge
Maximum	2	↑			vertex

Influence of Critical Points on the Betti Numbers of a simplicial complex

Betti Number Formulae

Using the incremental algorithm,

$$\beta_0(\nu_0) = \int_{\nu_0}^{\infty} [N_0(\nu) - (1 - g(\nu))N_1(\nu)] d\nu$$

$$\beta_1(\nu_0) = \int_{\nu_0}^{\infty} [g(\nu)N_1(\nu) - N_2(\nu)] d\nu$$

where $g(\nu)$ is the probability that an added saddle point at height ν increases β_1 by 1

Note:

because of symmetry maxima and minima in Gaussian field,

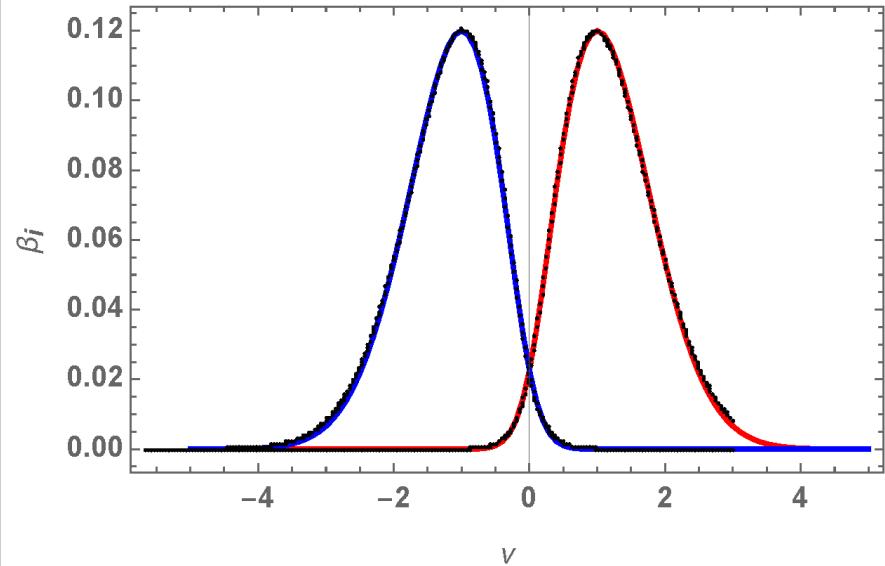
$$\beta_0(\nu) = \beta_1(-\nu)$$

Betti Number Prediction

$g(v)$ can be accurately fitted by the function,

$$g(v) \approx \frac{e^{-\alpha v} \left(1 - \frac{N_0(v)}{N_1(v)} \right) + e^{\alpha v} \frac{N_2(v)}{N_1(v)}}{e^{-\alpha v} + e^{\alpha v}}$$

$$\alpha \sim 2.7$$



Primordial non-Gaussianities

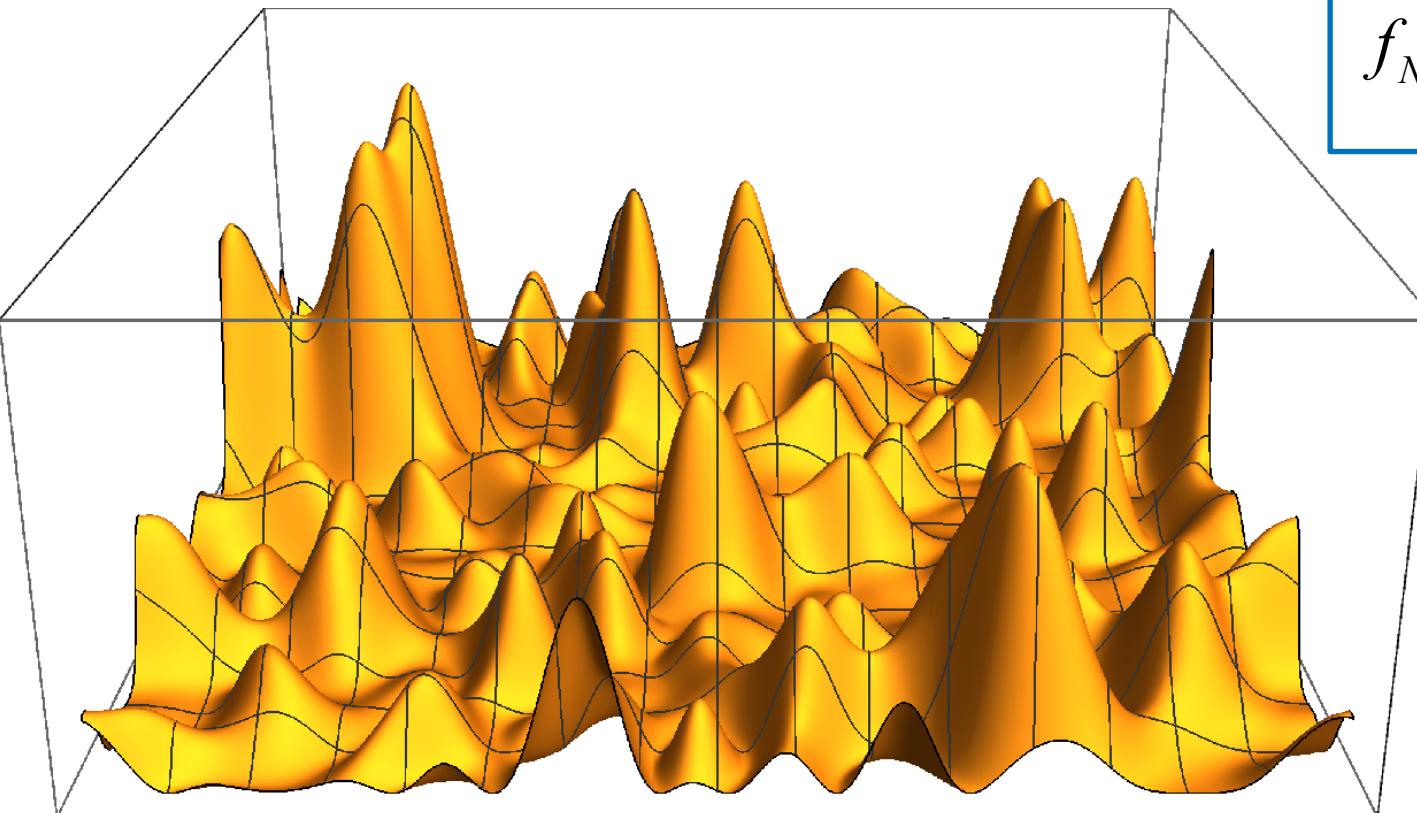
Topological Signatures

Primordial non-Gaussianities: f_{NL} parameter

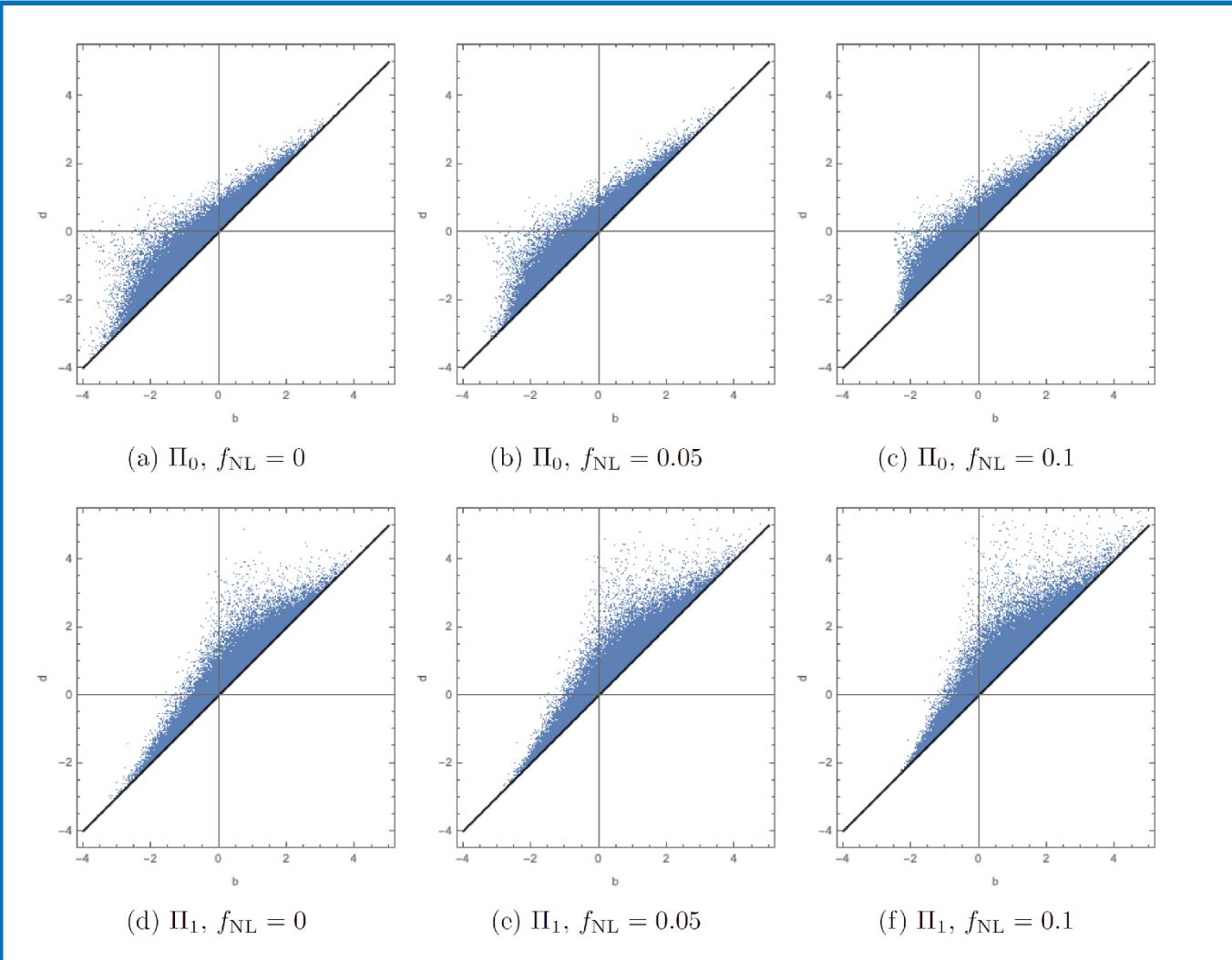
Feldbrugge, van Engelen, vdW, Vegter et al. 2014

also see Cole, Chiu et al. 2019, 2020

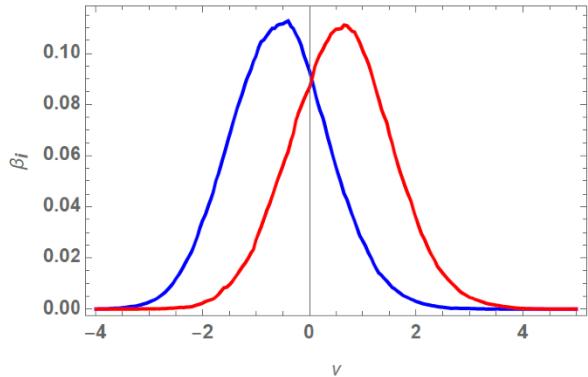
$$f_{NG} = f_G + f_{NL}f_G^2$$



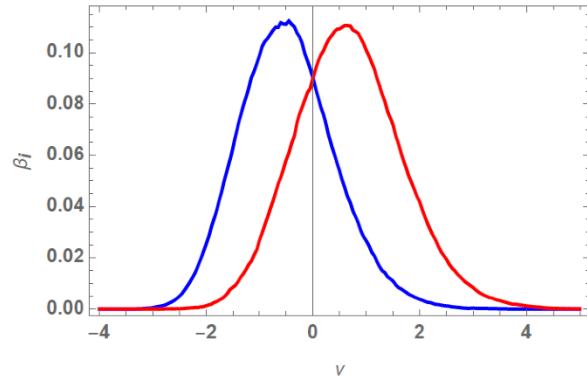
Primordial non-Gaussianities: Persistence



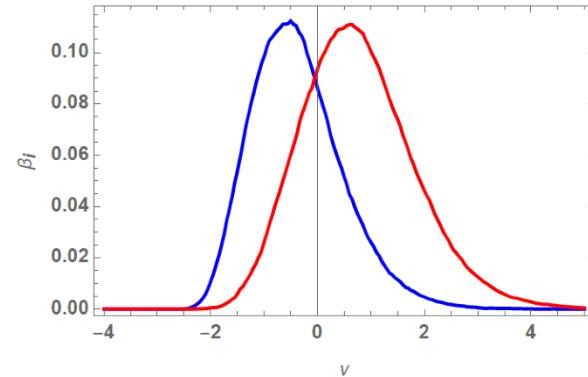
Primordial non-Gaussianities: Betti Numbers



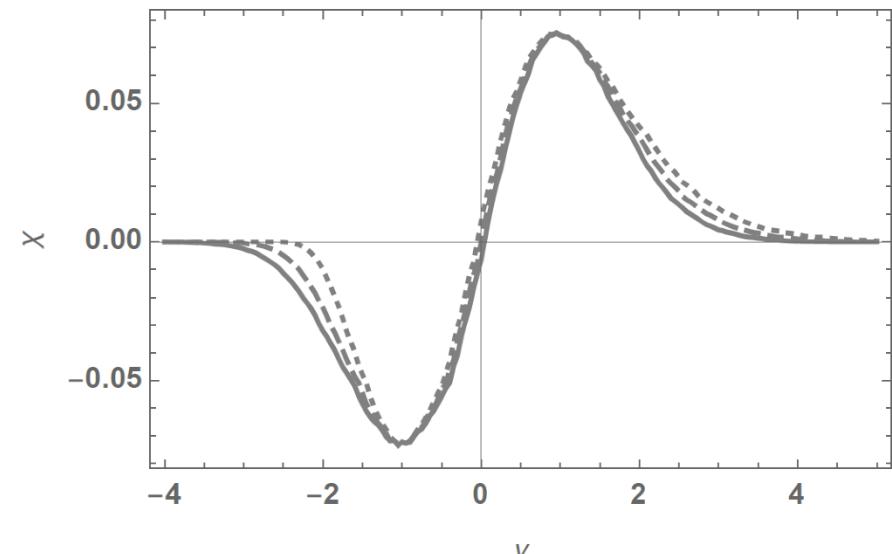
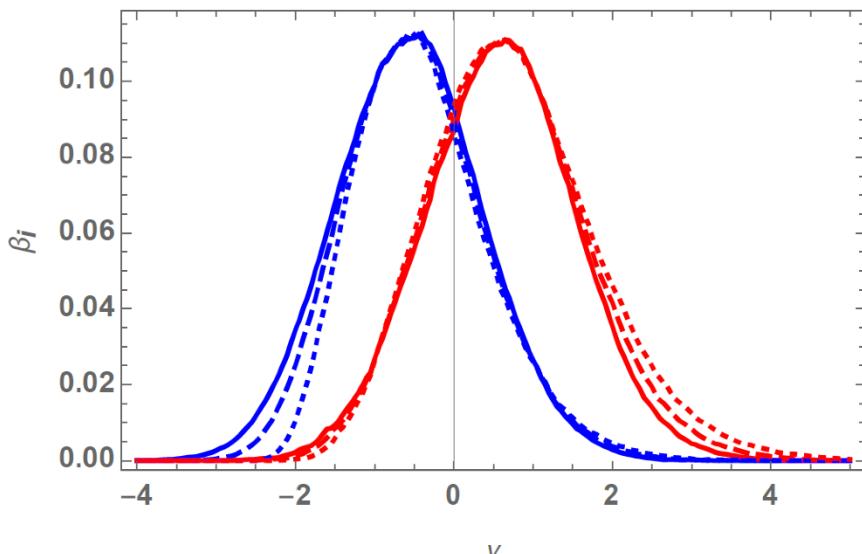
(a) $f_{\text{NL}} = 0$



(b) $f_{\text{NL}} = 0.05$

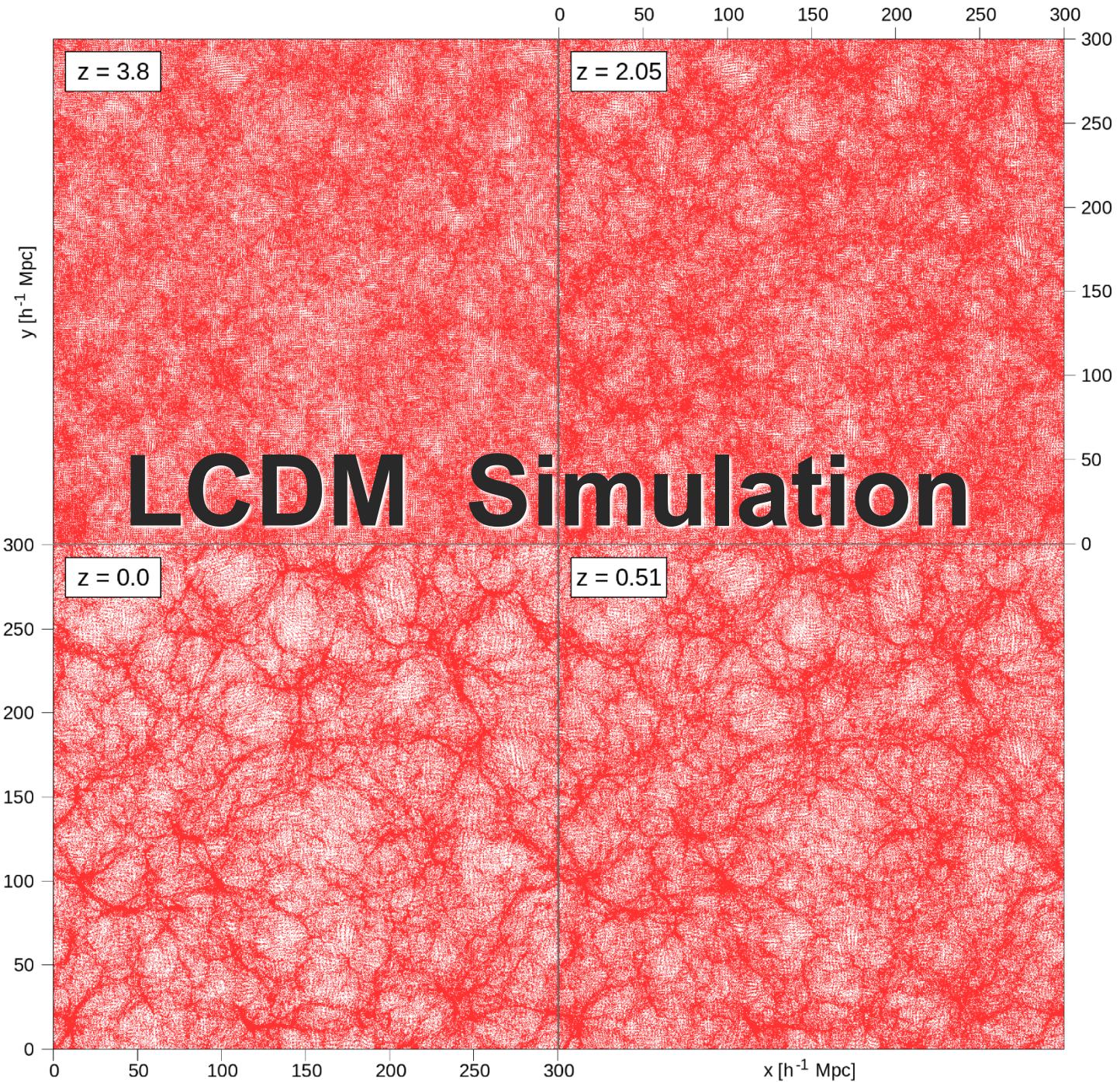


(c) $f_{\text{NL}} = 0.1$



Cosmic Web Topology:

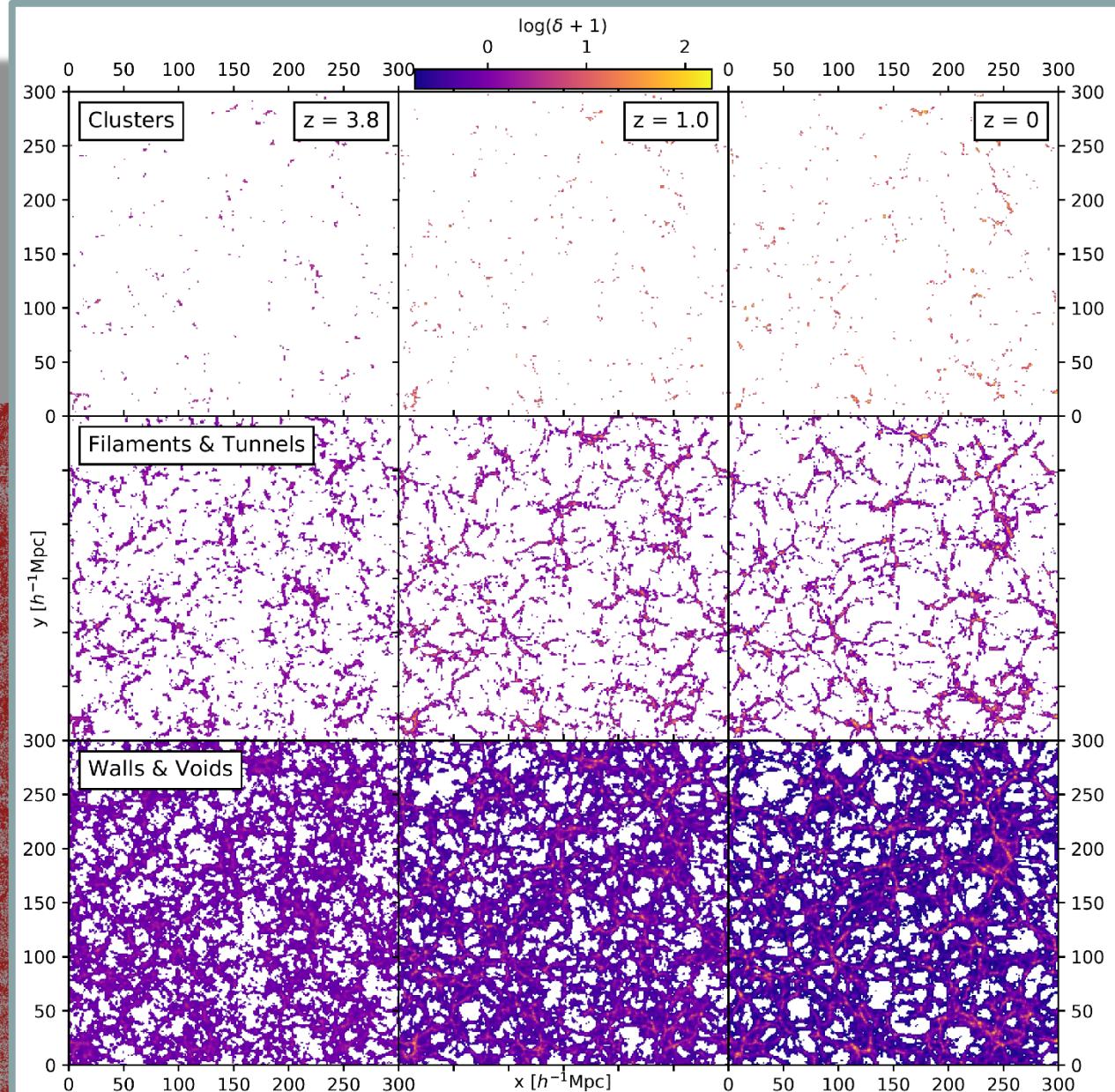
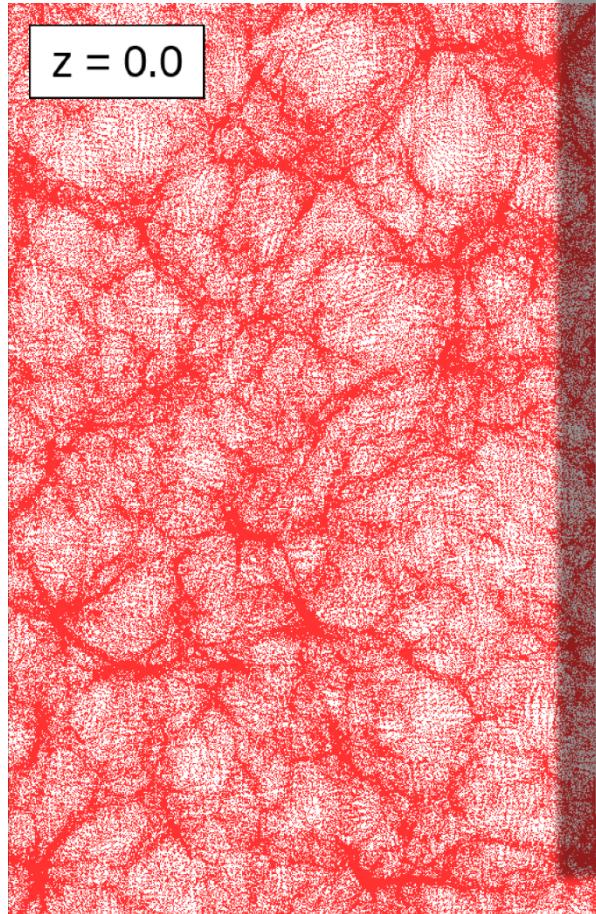
LCDM cosmology



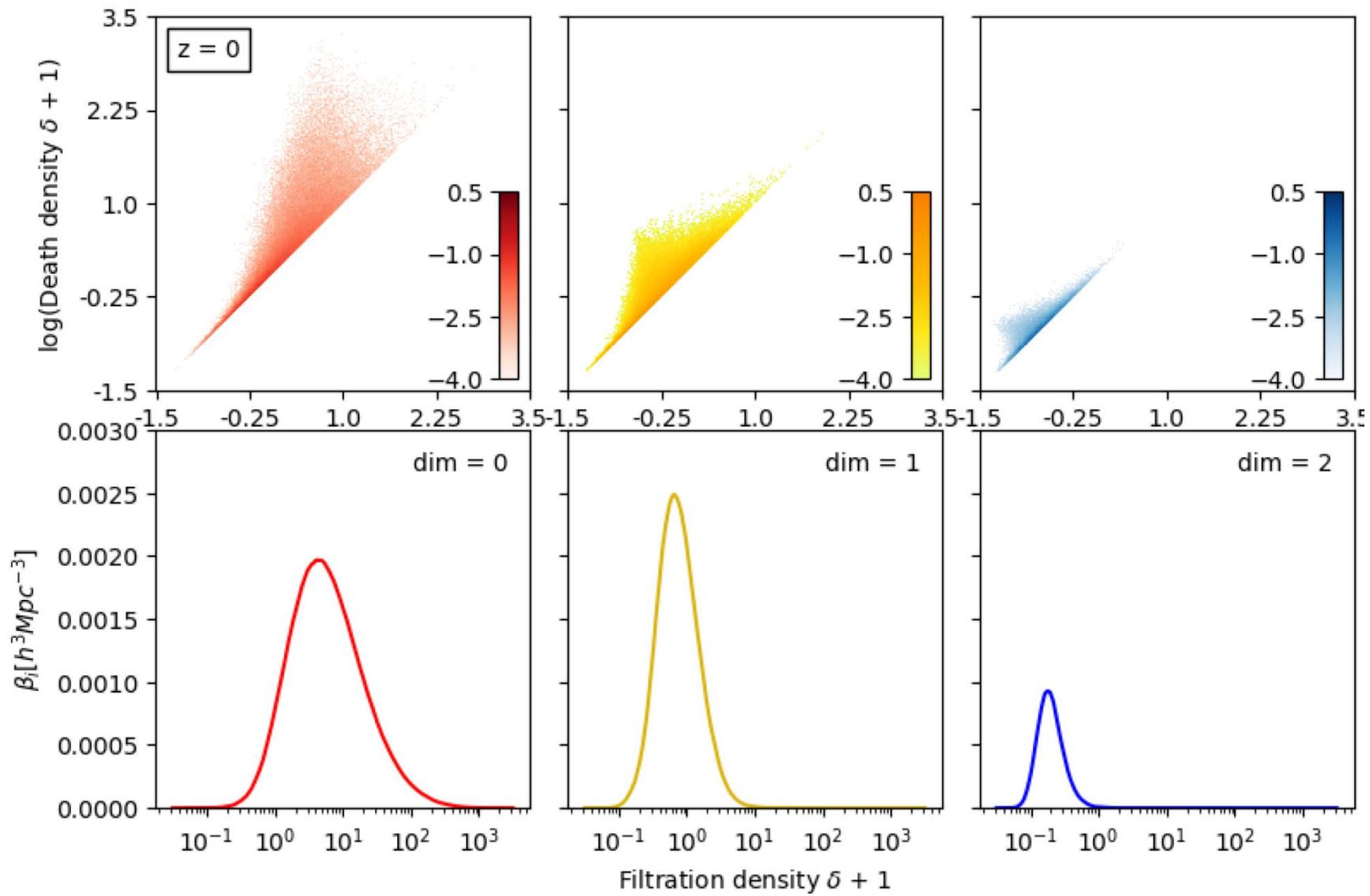
LCDM simulation

Density field

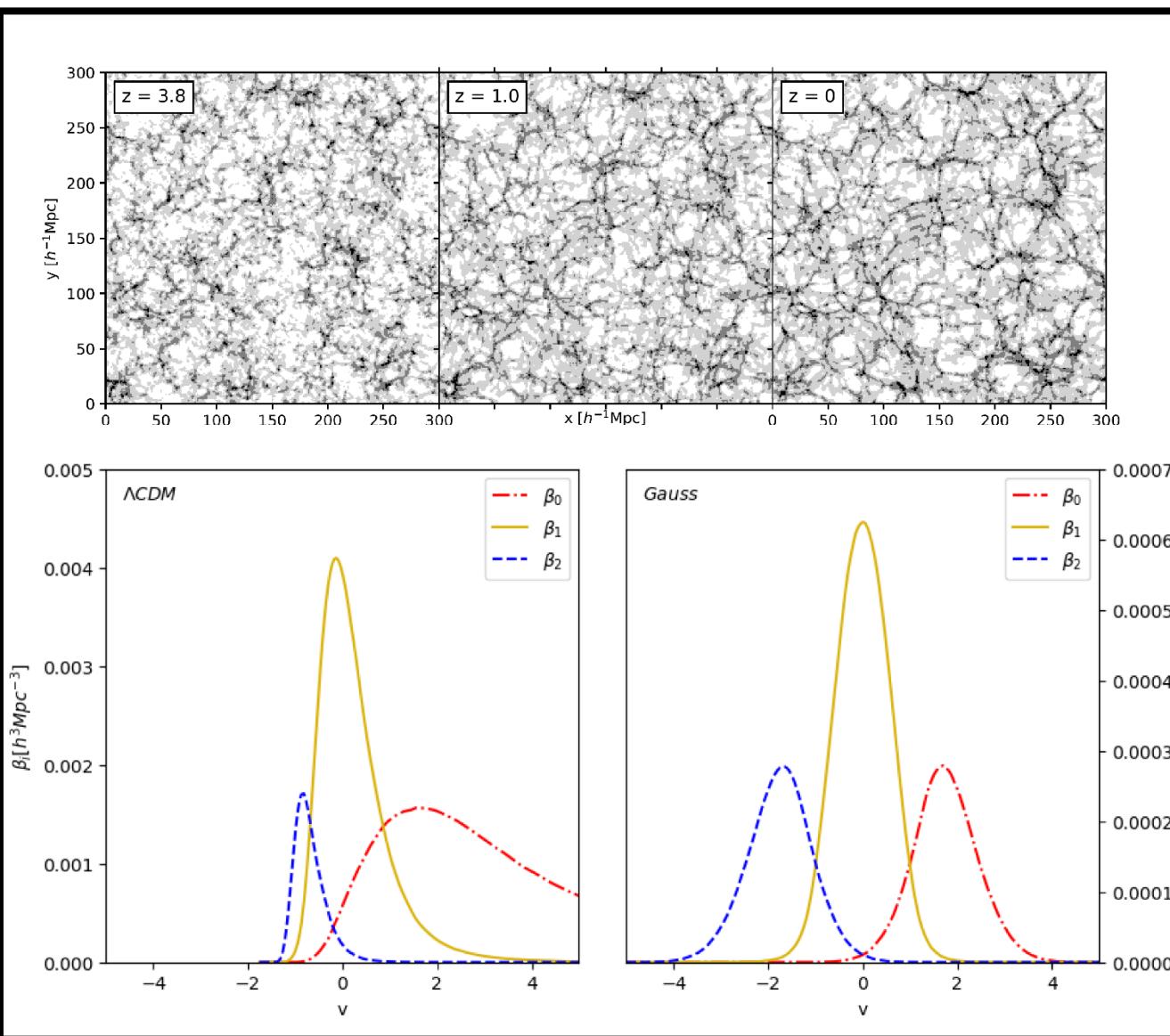
Superlevel sets



LCDM Betti- Persistence



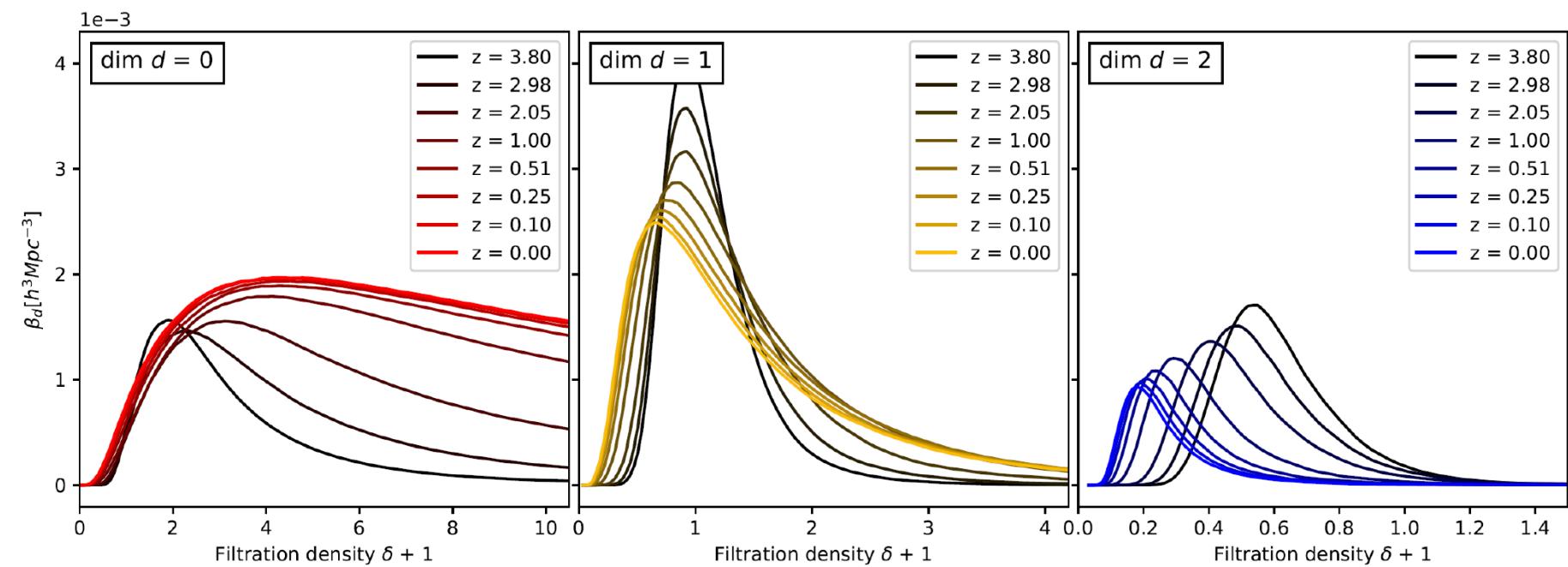
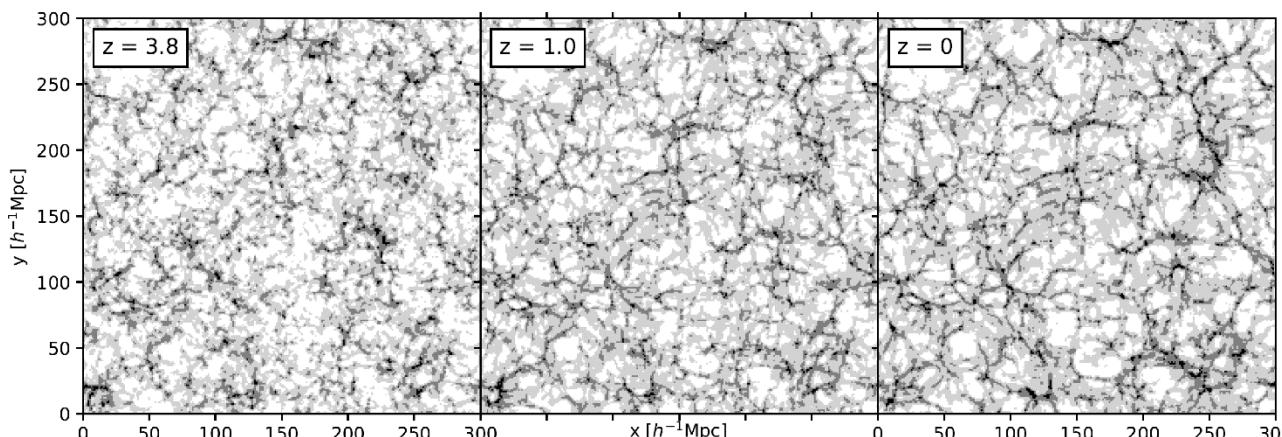
LCDM: Betti Number Evolution



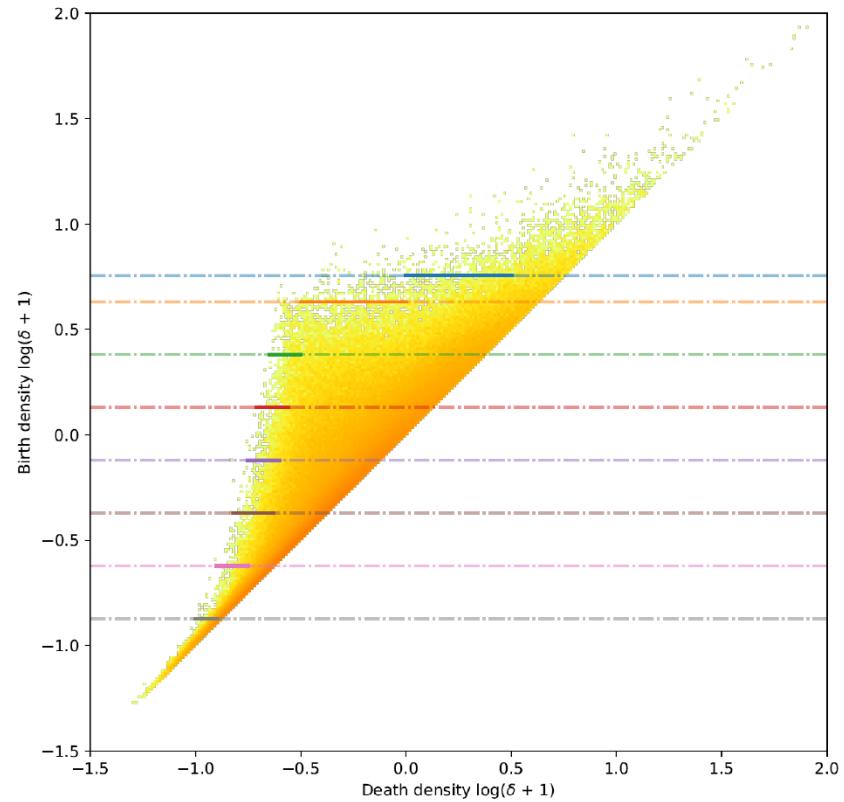
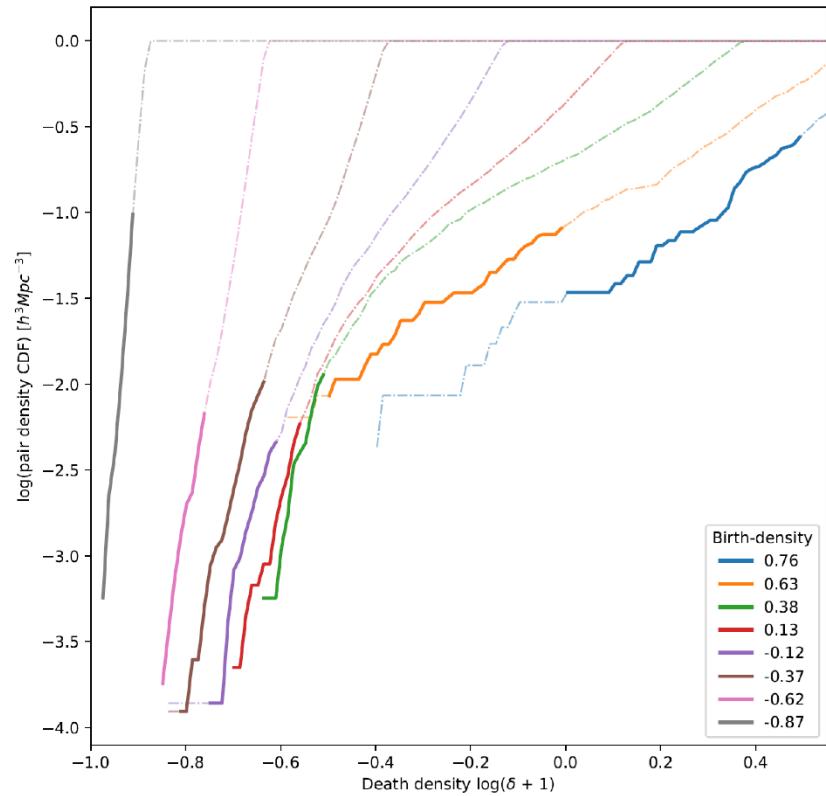
Betti Curve Evolution:

- Shift from symmetric Gaussian curves
- Asymmetry Overdense-Underdense
- High number of overdense clumps/islands
- Lower number of large deep voids

LCDM: Betti Number Evolution



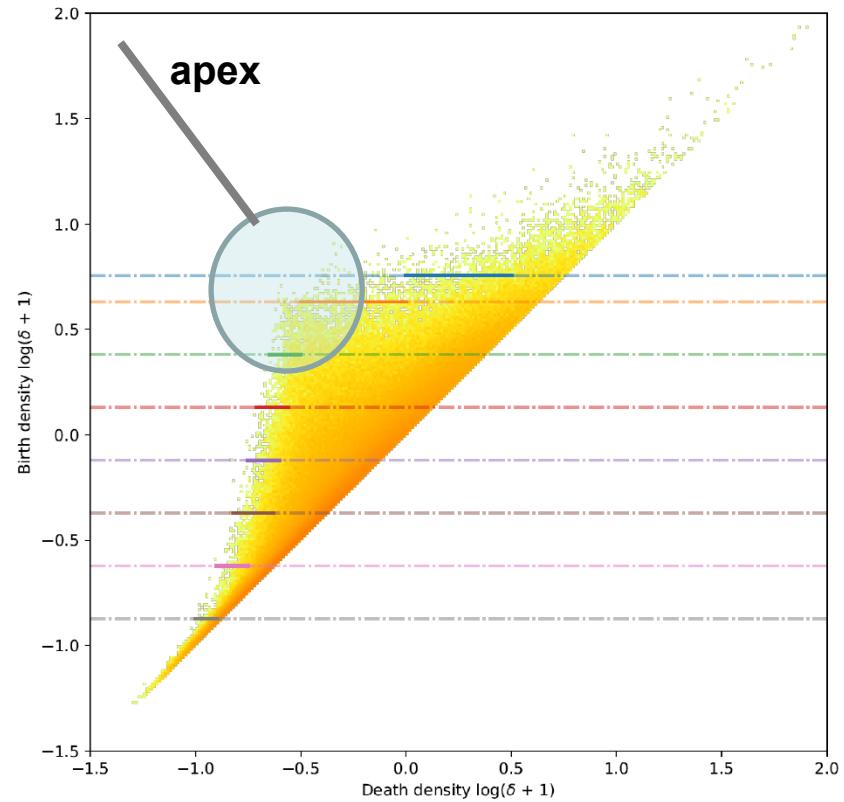
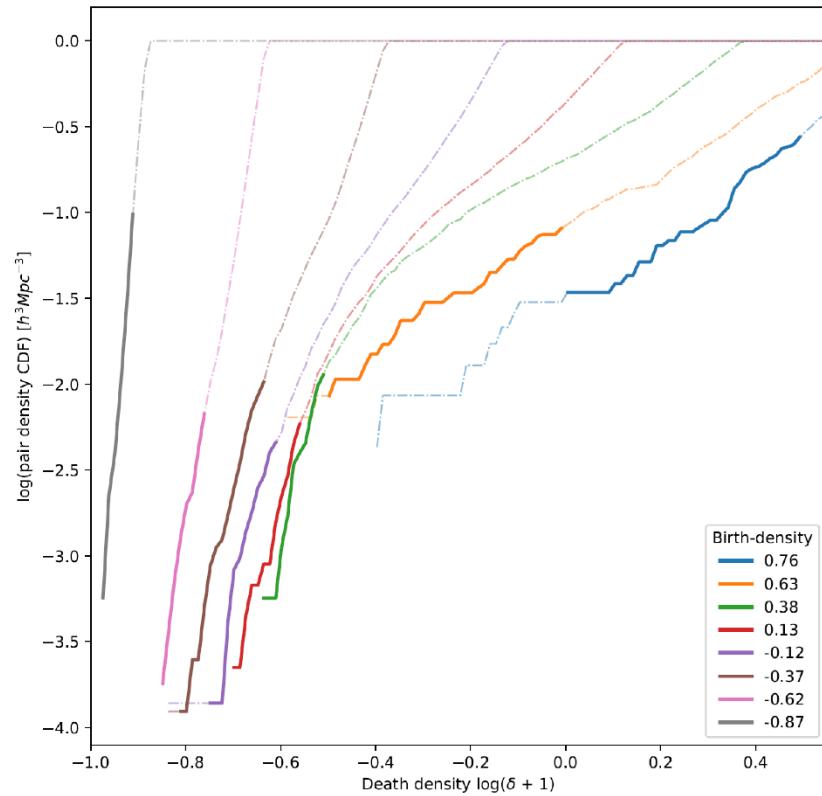
Singularity Persistence



LCDM

Internal structure - dim=1 persistence diagram

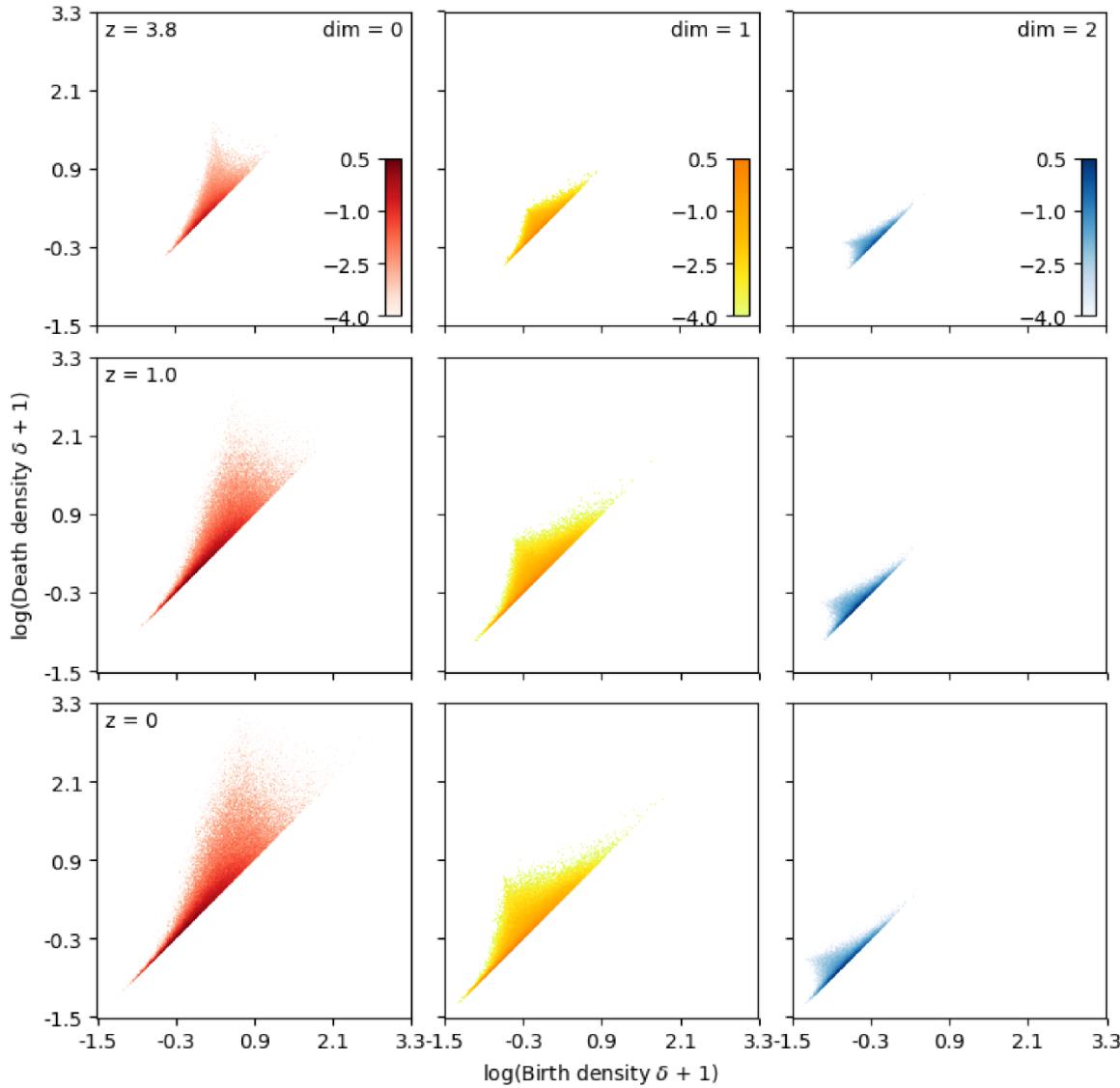
Singularity Persistence



Characteristic features correspond with dynamics structure formation process:

- Apex corresponds to characteristic densities major transition in population objects
- Dim=0 - Dim=1: disappearance supercluster islands - appearance weblike network by filaments
- Dim=1 - Dim=2: filaments disappear as walls close - birth of void population

LCDM: Persistent Evolution

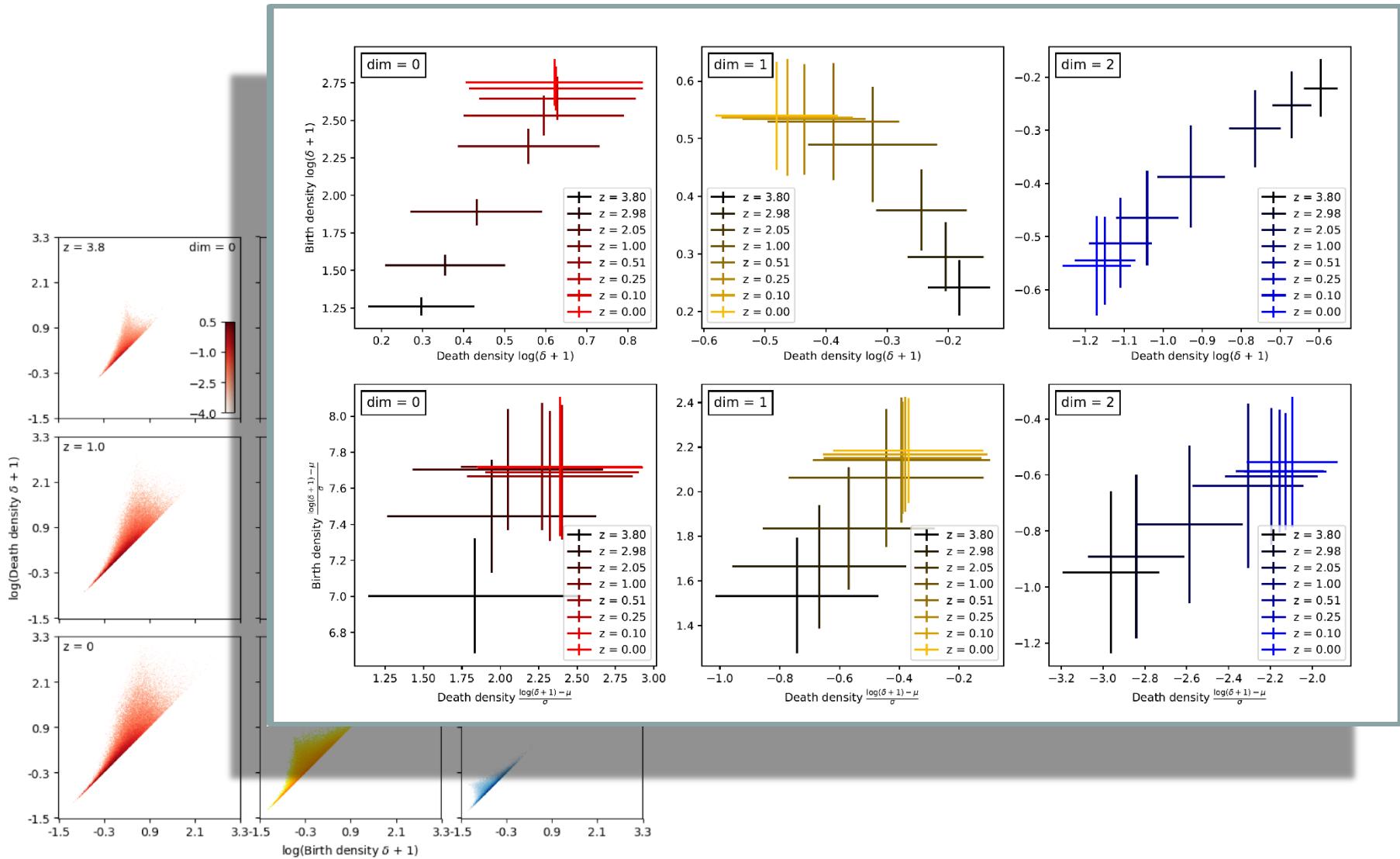


Shifting Persistence Diagrams:

- **Evolving Cosmic Web**
- **Traces characteristic density levels of island complexes filaments/tunnels voids**
- **Dynamical transitions: apex shift**

Persistence Apex Shift:

Self-similar (hierarchical) evolution



Topological Bias:

Halo Population

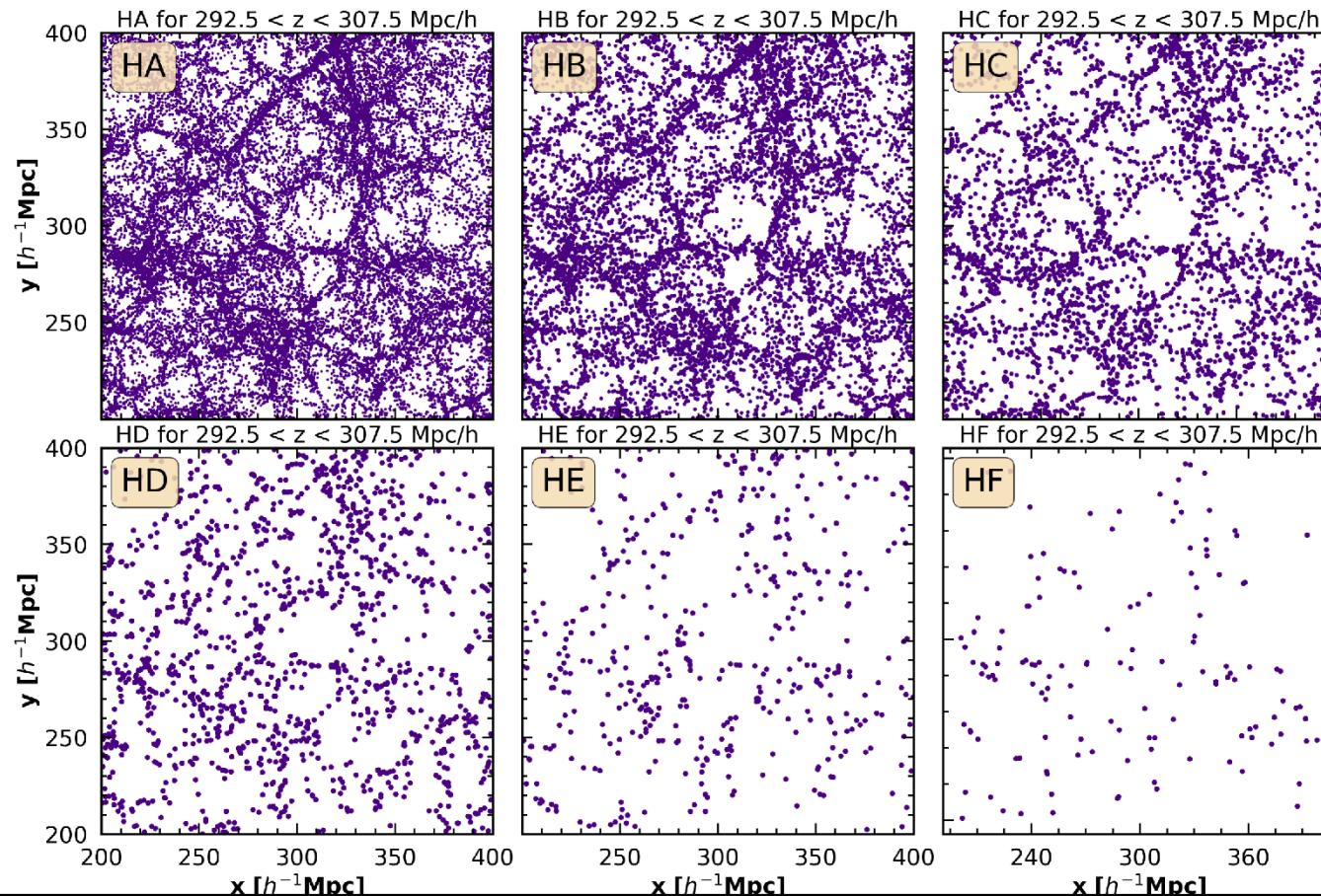


Millennium

Halo Tracing – Cosmic Web

Cosmic Web Observed:

- traced by galaxies – galaxy halos
- Dependence observed structure cosmic web on halo/galaxy population
- Dependence topological characteristics on halo/galaxy population

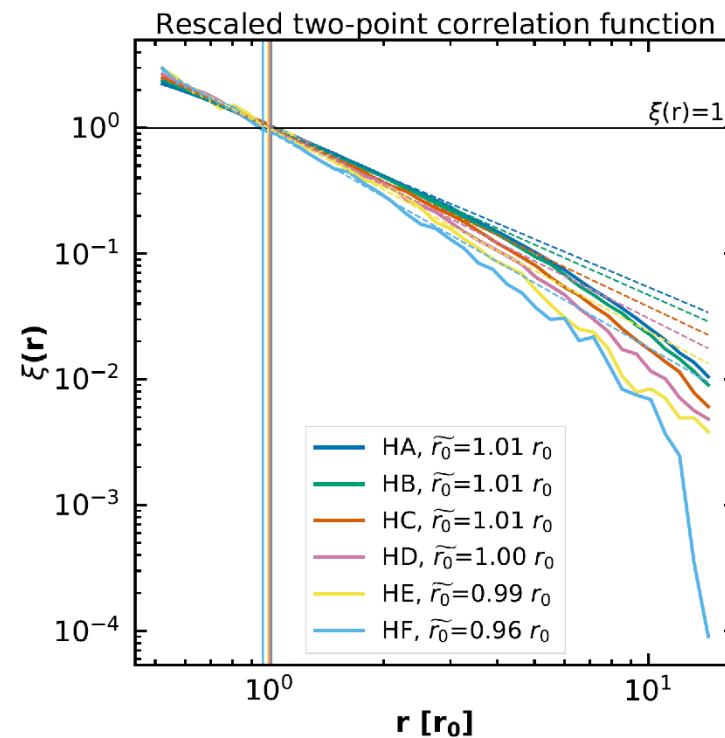
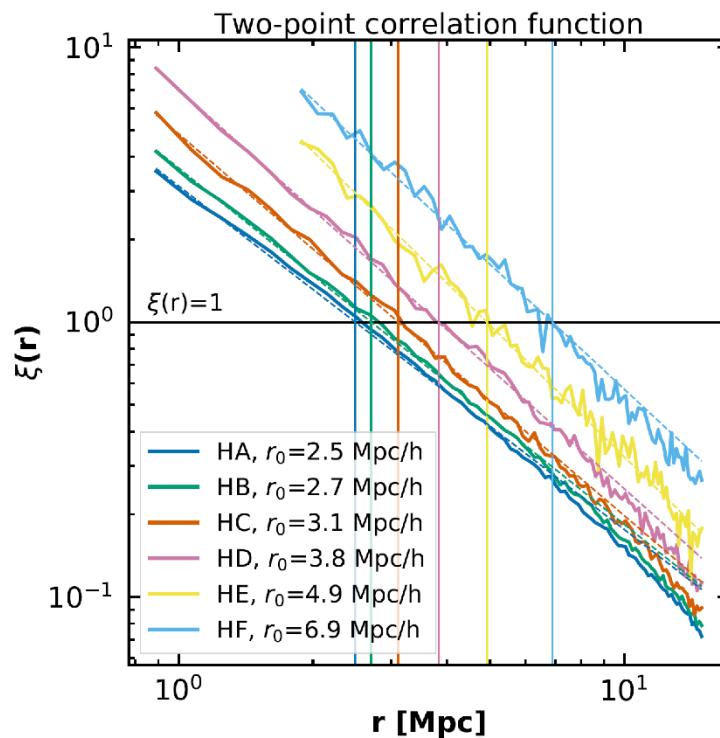


Halo Tracing – Cosmic Web

Cosmic Web Observed:

- P-Millennium simulation (Cole et al.)
- Halo population in 6 mass ranges
- Clustering:
- 2nd order correlation function $\xi(r)$
- Rescaling by correlation length

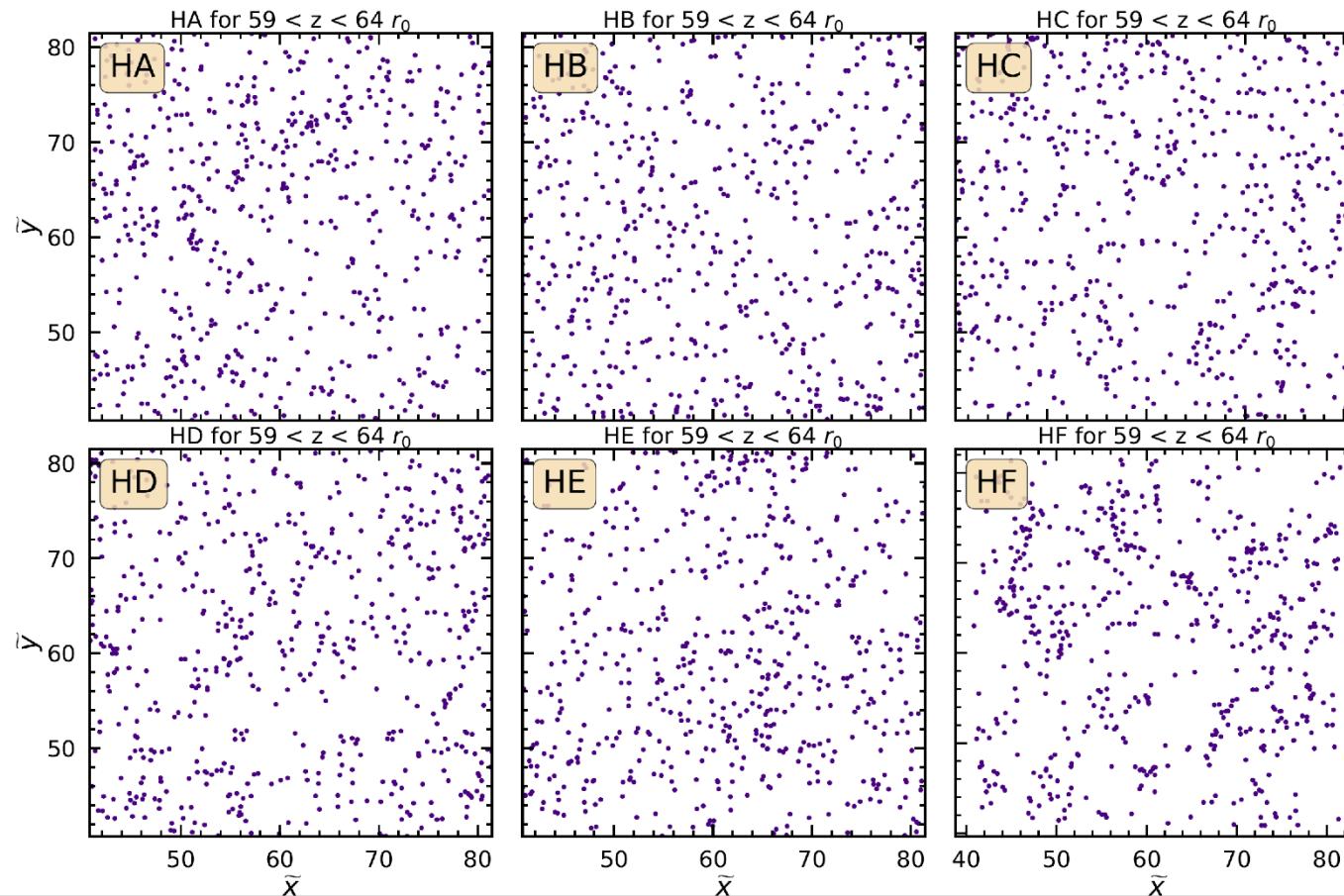
Halo Population	$\log_{10}(M) [h^{-1}M_\odot]$	$r_0 [h^{-1}Mpc]$	γ	$\tilde{r}_0 [r_0]$	$\tilde{\gamma}$
HA	(10.5,11]	2.49 ± 0.01	-1.251 ± 0.008	1.011 ± 0.005	-1.27 ± 0.01
HB	(11,11.5]	2.70 ± 0.01	-1.297 ± 0.006	1.009 ± 0.004	-1.34 ± 0.01
HC	(11.5,12]	3.11 ± 0.02	-1.390 ± 0.008	1.008 ± 0.004	-1.43 ± 0.01
HD	(12,12.5]	3.83 ± 0.02	-1.454 ± 0.006	1.005 ± 0.005	-1.52 ± 0.01
HE	(12.5,13]	4.92 ± 0.04	-1.59 ± 0.02	0.99 ± 0.01	-1.61 ± 0.03
HF	(13,13.5]	6.89 ± 0.09	-1.51 ± 0.02	0.961 ± 0.007	-1.72 ± 0.03



Halo Tracing – Cosmic Web

Cosmic Web Observed:

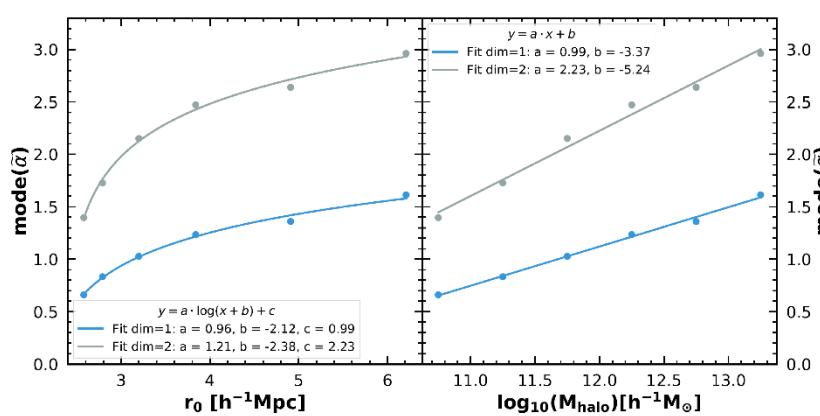
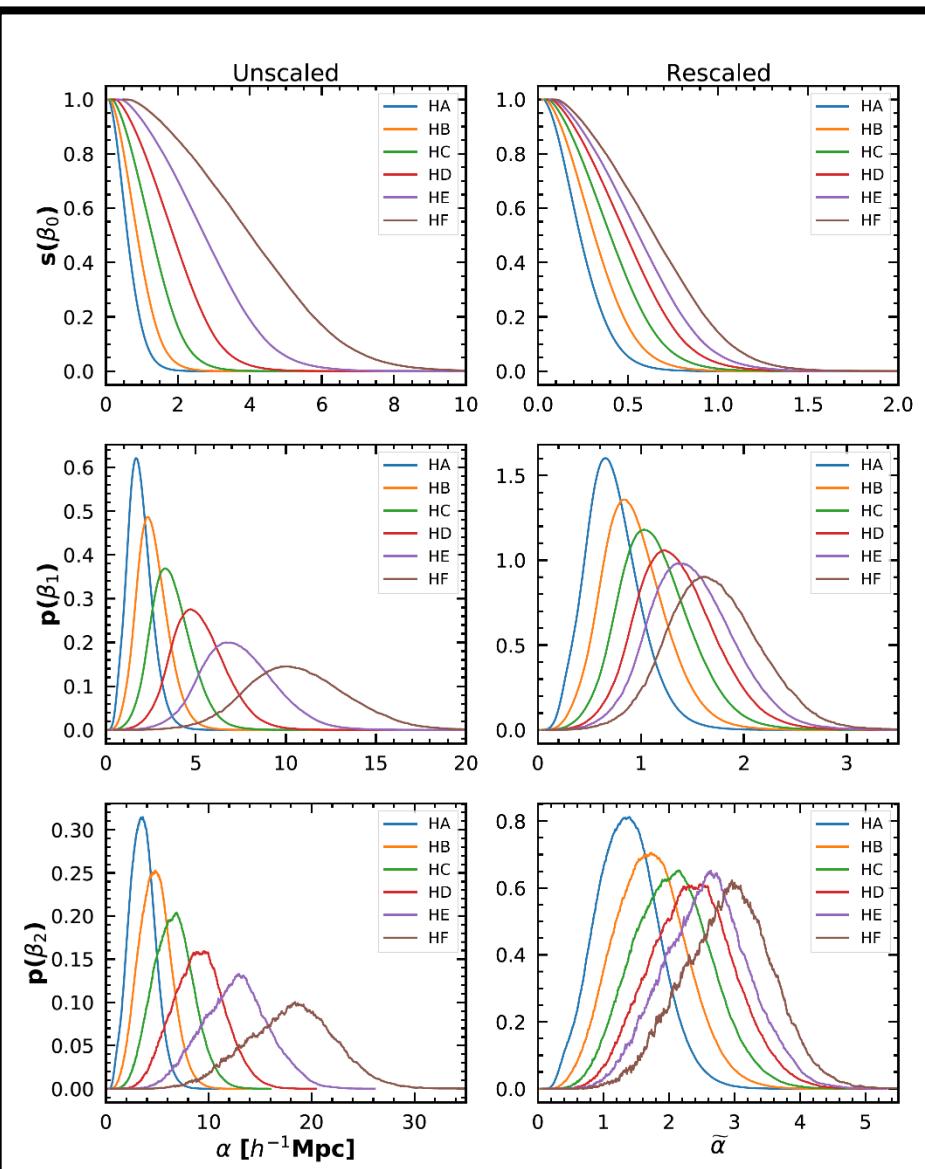
- traced by galaxies – galaxy halos
- Incorporating 2nd order clustering – rescaling by correlation length
- Still dependence observed structure cosmic web on halo/galaxy population: higher order structural pattern



Betti Number - Halo Population

Cosmic Web Observed:

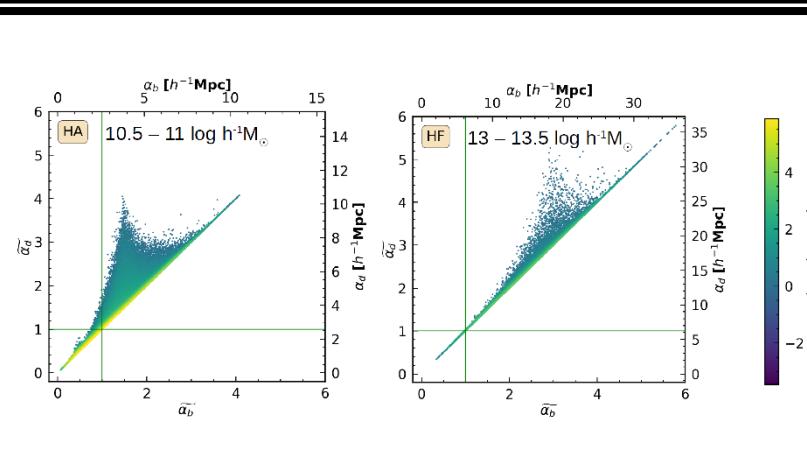
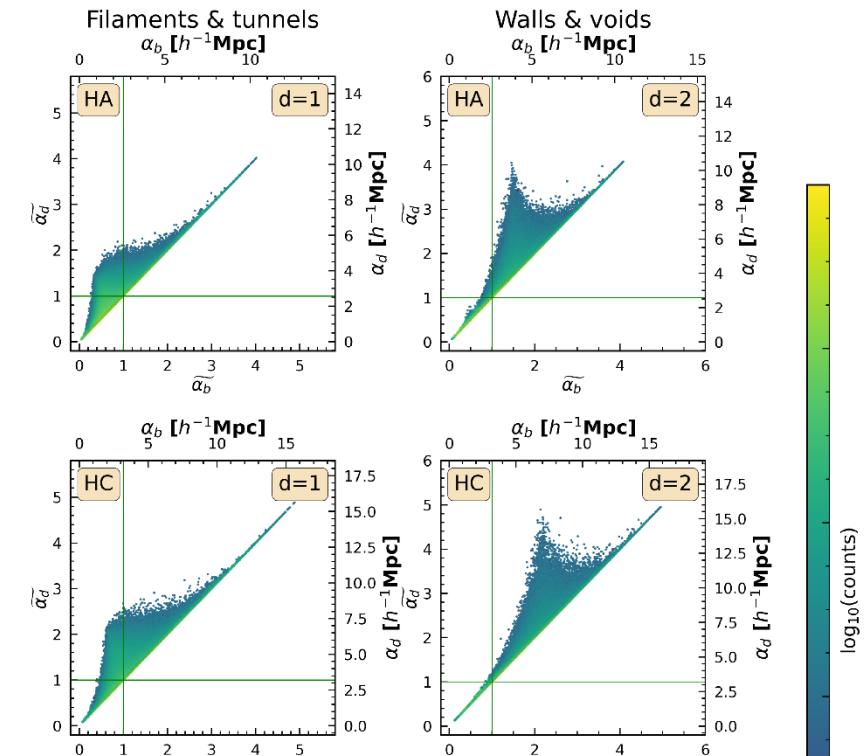
- Expected scaling with increasing mass range:
shift to larger scales
- After rescaling by correlation length:
 - still shift to larger (scaled) scales
 - beautiful self-similar increase
characteristic scale
 - different topology cosmic web traced
by different populations
 - eg. Voids different halo tracers significantly
different features !



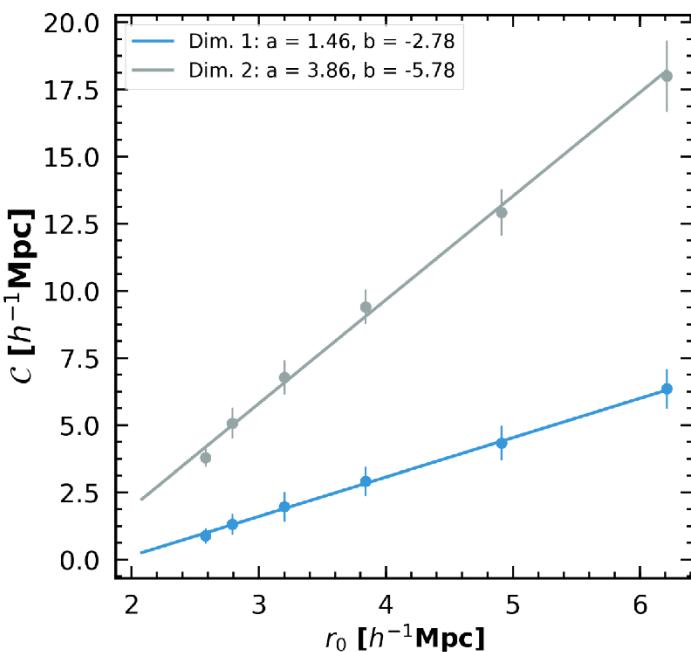
Persistence - Halo Population

Cosmic Web Observed:

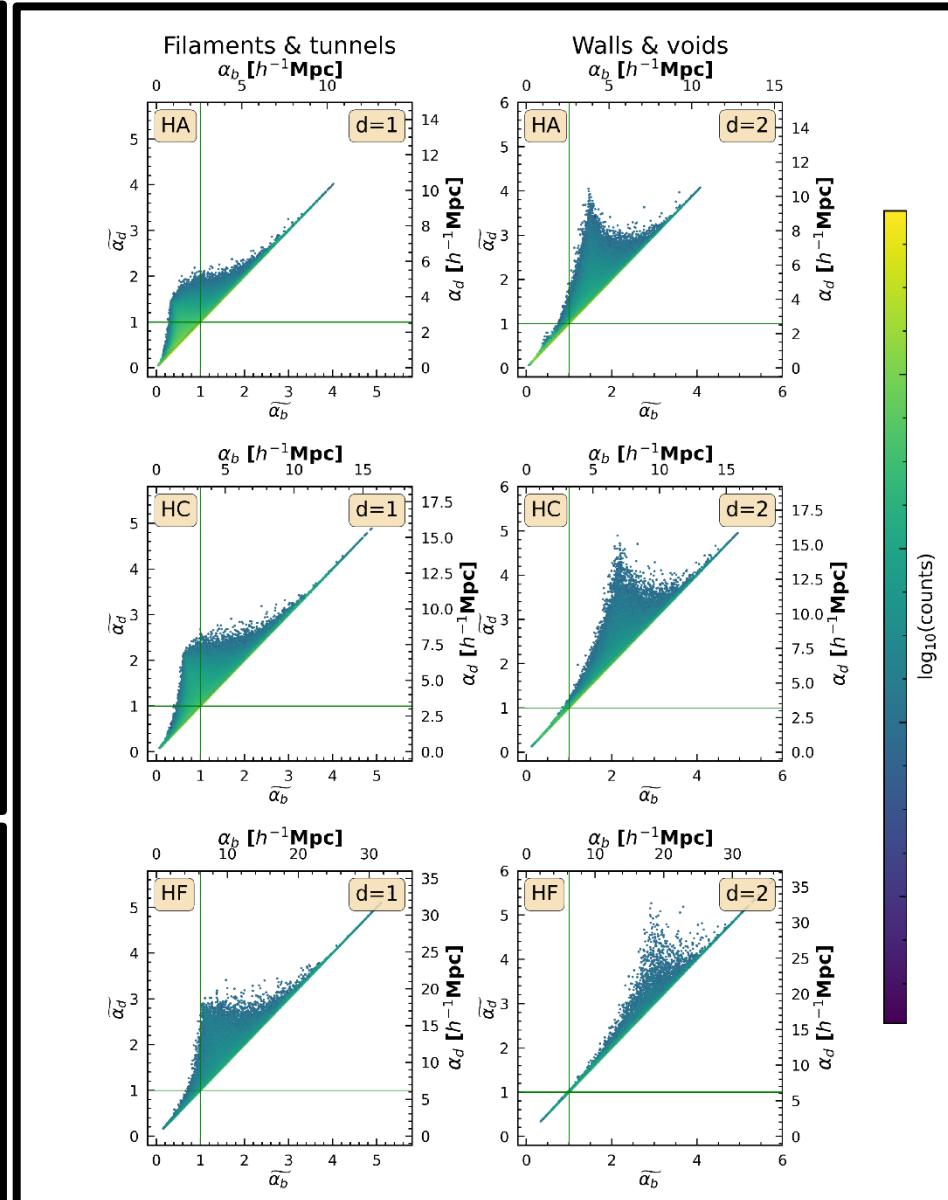
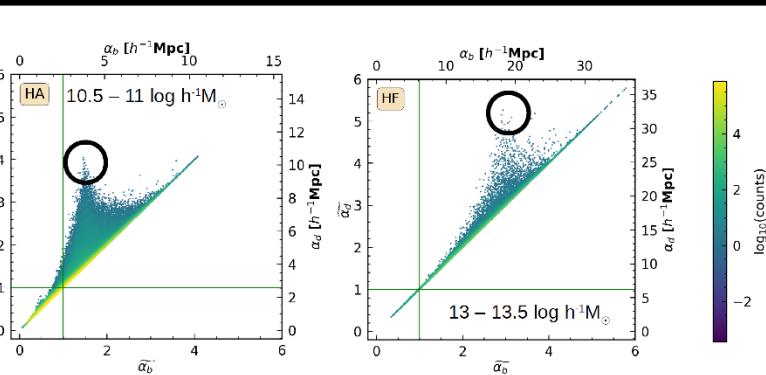
- Expected scaling with increasing mass range: shift to larger scales
- After rescaling by correlation length:
 - still shift to larger (scaled) scales
 - beautiful self-similar increase characteristic scale
 - different topology cosmic web traced by different populations
 - eg. Voids different halo tracers significantly different features !



Persistence - Halo Population



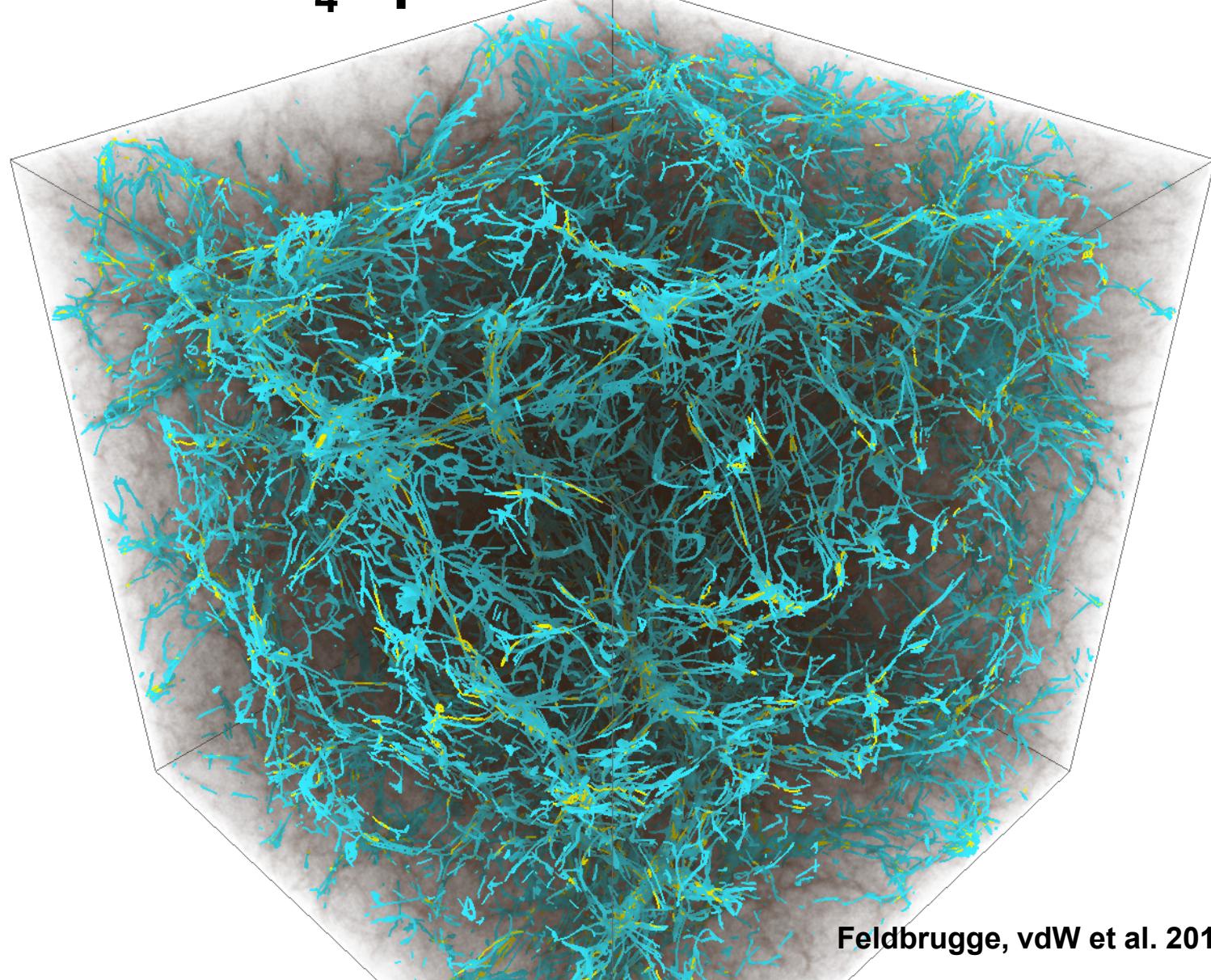
- beautiful self-similar increase
characteristic scale apex



Cosmic Web Connectivity: Caustic Skeleton

Feldbrugge, vdW et al. 2018, 2019
Hidding, Shandarin & vdW 2014
Wilding, Feldbrugge, vdW 2021

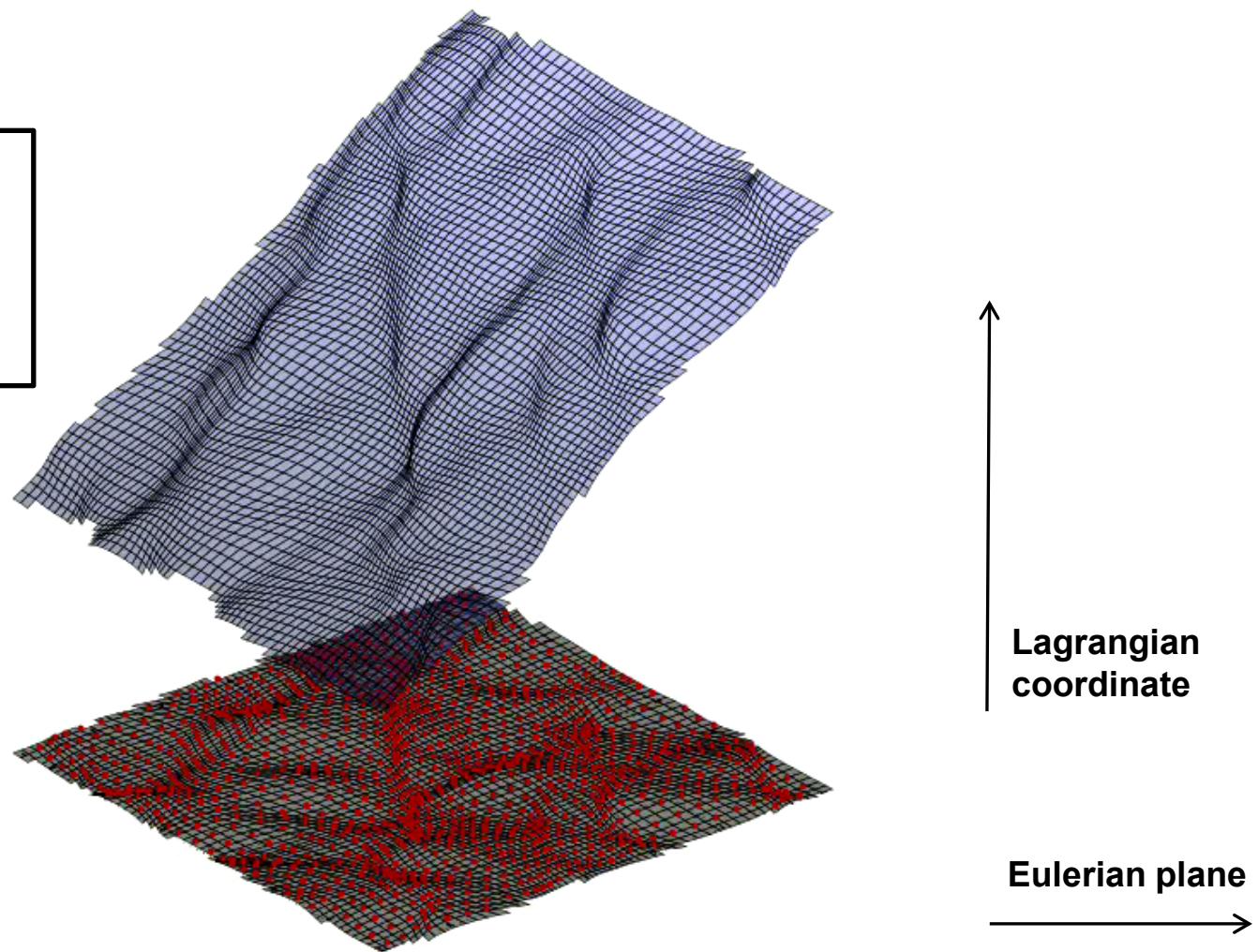
Skeleton (3D) Cosmic Web: A_4 spine - swallowtails



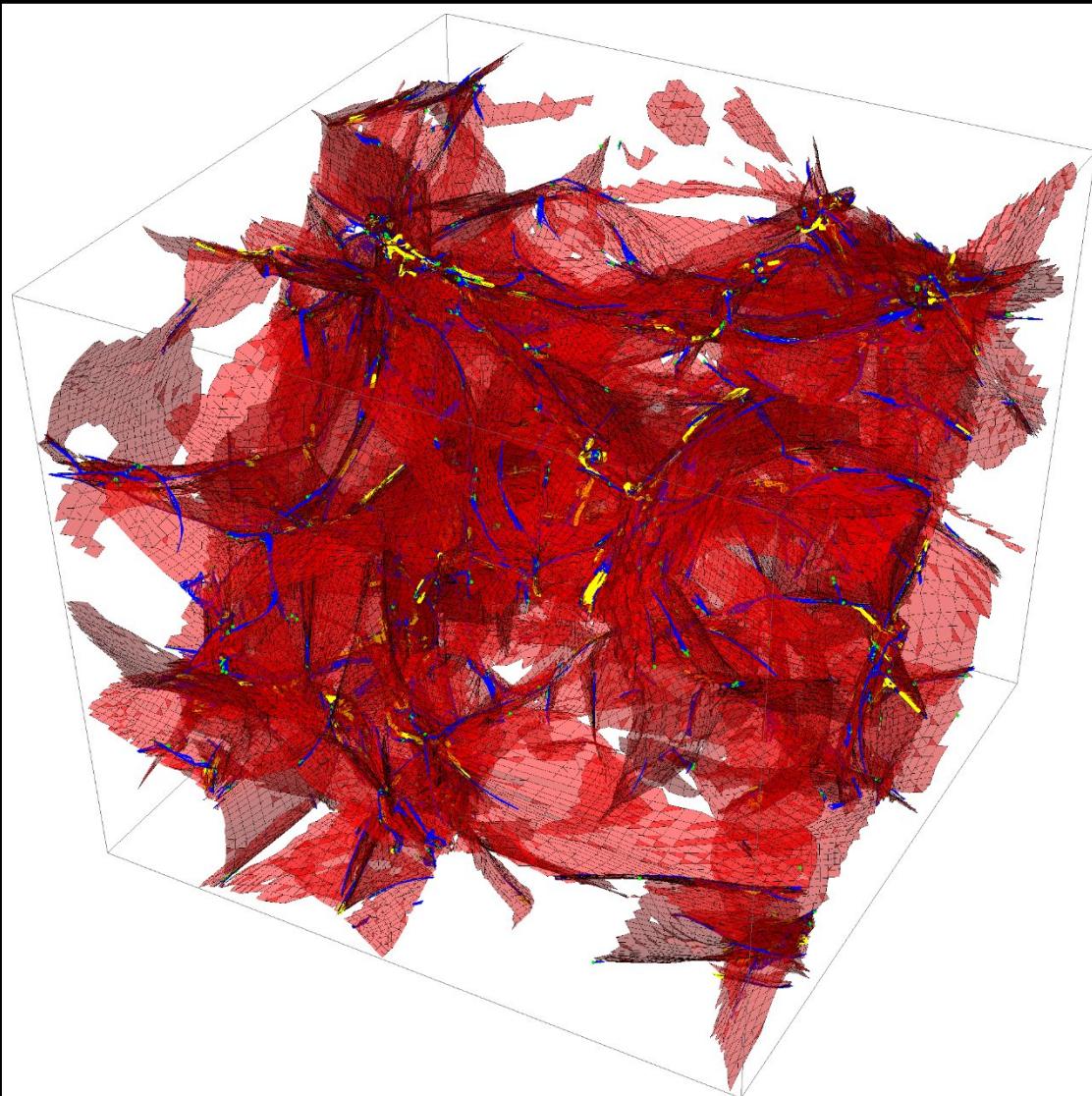
Phase-Space Dynamics

Dynamical Evolution:

folding the
phase-space sheet $\{q, x\}$



Cosmic Web - Caustic Skeleton



Cosmic Web: Caustic features



Deformation field (Gaussian initial density field)

Caustic structure

Red sheets - cusps A3 singularities - walls

Blue lines - swallowtail A4 sing. - filaments

Green dots - butterfly A5 sing. - nodes

Yellow lines - D4 umbilics - filaments

Caustic connectivity:

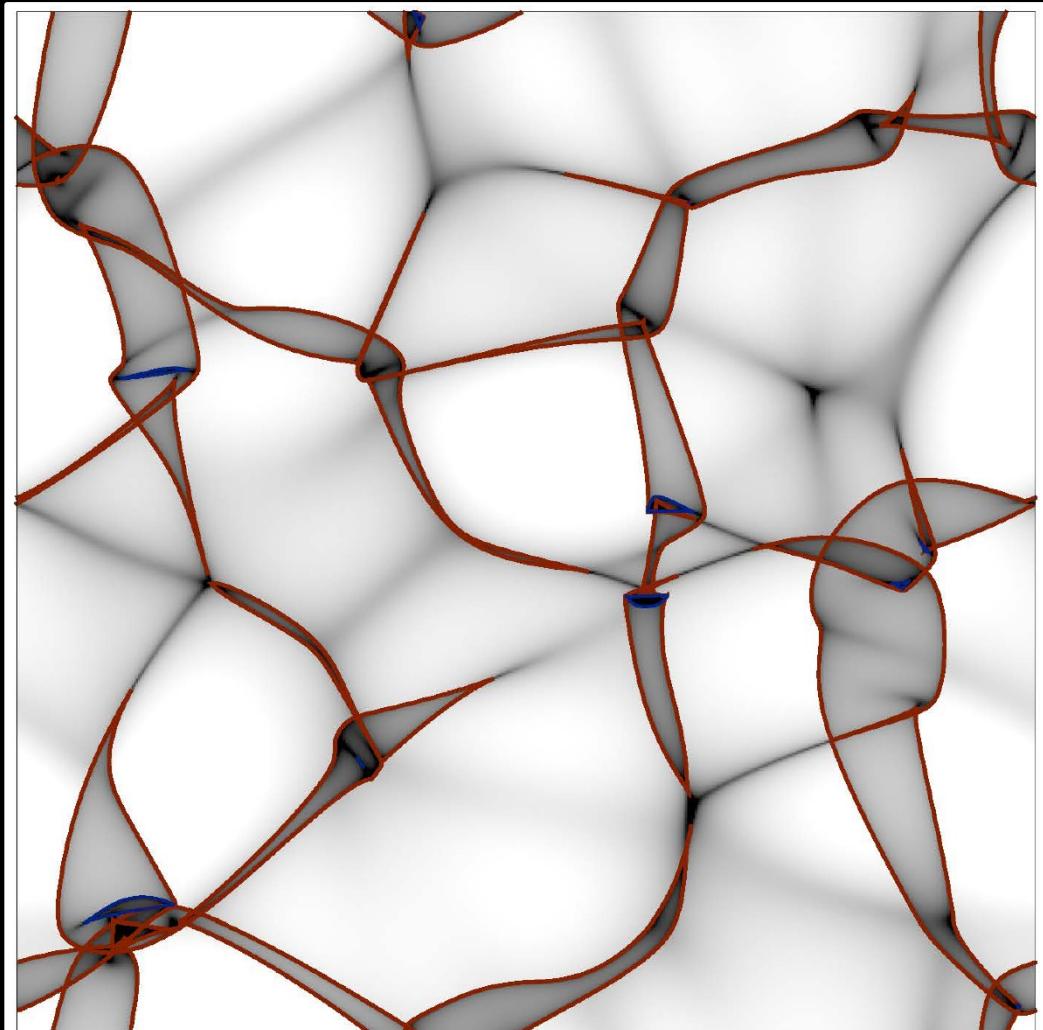
Topology of the deformation field

Zel'dovich Approximation

$$\vec{x} = \vec{q} + D(t) \vec{u}(\vec{q})$$

$$\vec{u}(\vec{q}) = -\vec{\nabla} \Phi(\vec{q})$$

$$\Phi(\vec{q}) = \frac{2}{3Da^2H^2\Omega} \phi_{lin}(\vec{q})$$

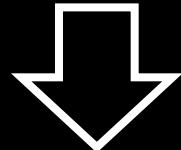


Zel'dovich Approximation

$$\vec{x} = \vec{q} + D(t) \vec{u}(\vec{q})$$

$$\vec{u}(\vec{q}) = -\vec{\nabla} \Phi(\vec{q})$$

$$d_{ij} = -\frac{\partial u_i}{\partial q_j}$$

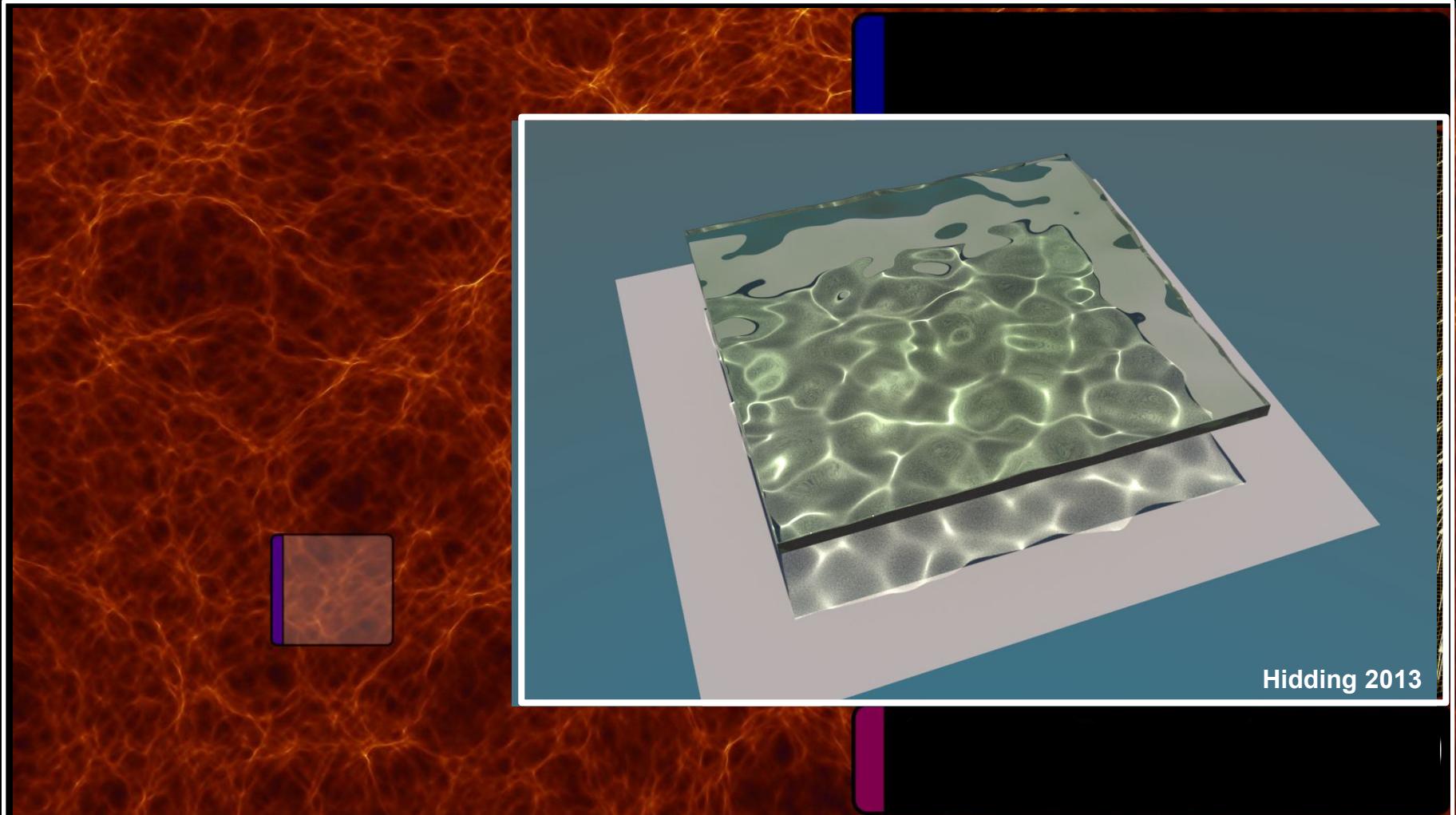


$$\rho(\vec{q}, t) = \frac{\rho_u(t)}{(1 - D(t)\lambda_1(\vec{q}))(1 - D(t)\lambda_2(\vec{q}))(1 - D(t)\lambda_3(\vec{q}))}$$

structure of the cosmic web determined by the spatial field of eigenvalues

$\lambda_1, \lambda_2, \lambda_3$

Deformation, Streaming & Caustics



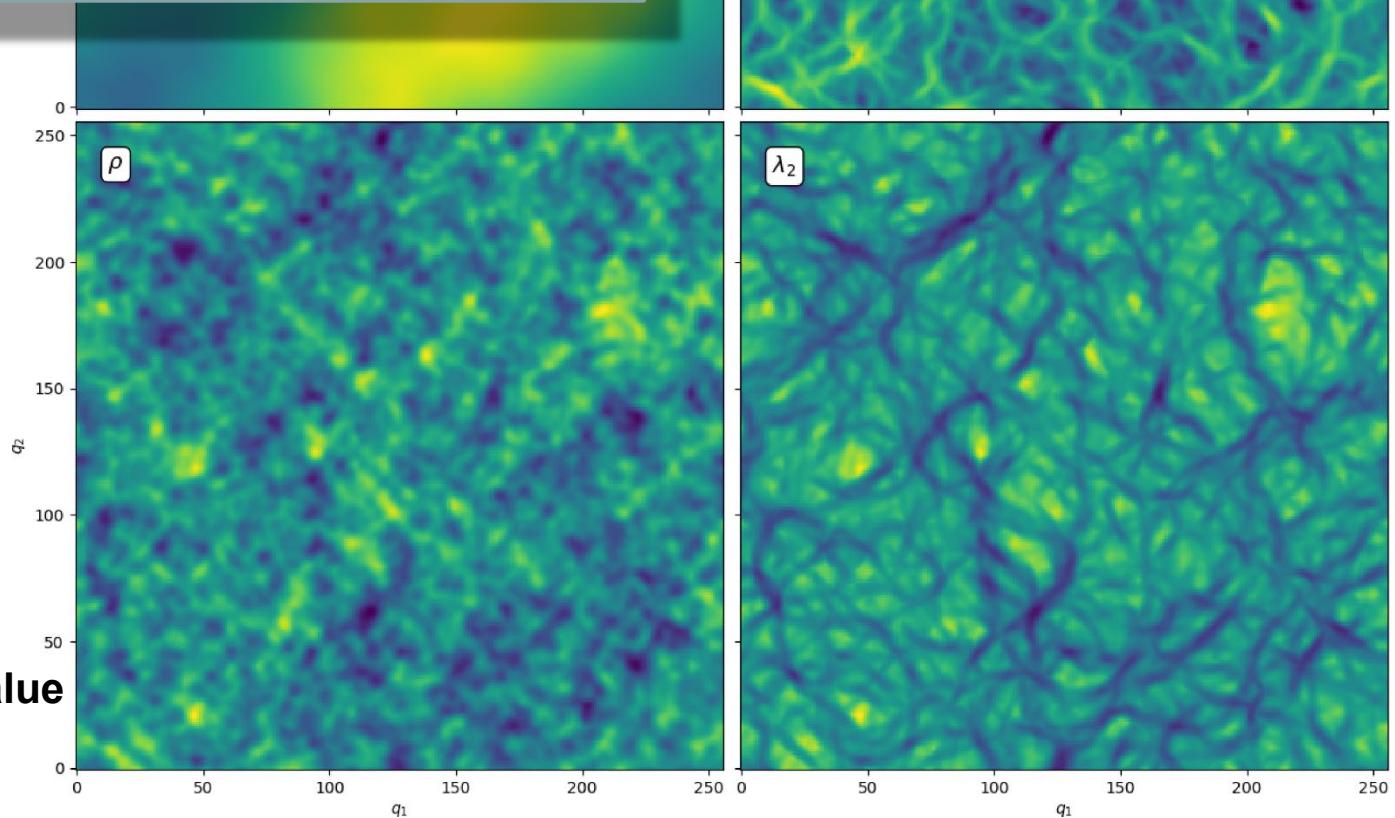
Hidding 2013

Illustration of the formation of caustics due to
streaming paths of light through deforming medium

Deformation Field Topology

$$\vec{x}(\vec{q}, t) = \vec{q} - D(t) \vec{\nabla} \Phi(\vec{q}) \quad \Rightarrow \quad d_{ij} = \frac{\partial^2 \Phi}{\partial q_i \partial q_j}: \lambda_1, \lambda_2, \lambda_3$$

$$\rho(\vec{q}, t) = \frac{\rho_u(t)}{(1 - D(t)\lambda_1(\vec{q}))(1 - D(t)\lambda_2(\vec{q}))(1 - D(t)\lambda_3(\vec{q}))}$$



Zeldovich
deformation eigenvalue
landscape

Caustic Conditions & Structure Formation: Eigenvalues & Eigenvectors Deformation Field

In Lagrangian space L (coordinates q):

A singularity forms in a manifold $M \subset L$ at location q_s when at q_s ,

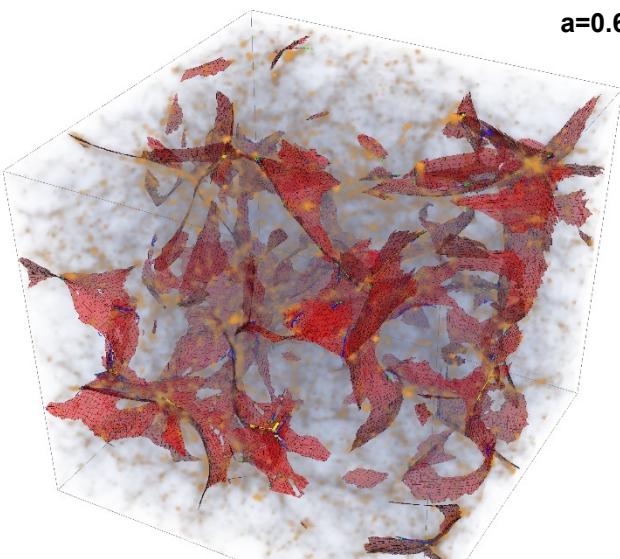
- the deformation tensor eigenvalue $\mu_i(q_s)$
- the corresponding eigenvector $\vec{v}_i(q_s)$

when at least one nonzero tangent vector \vec{T} such that

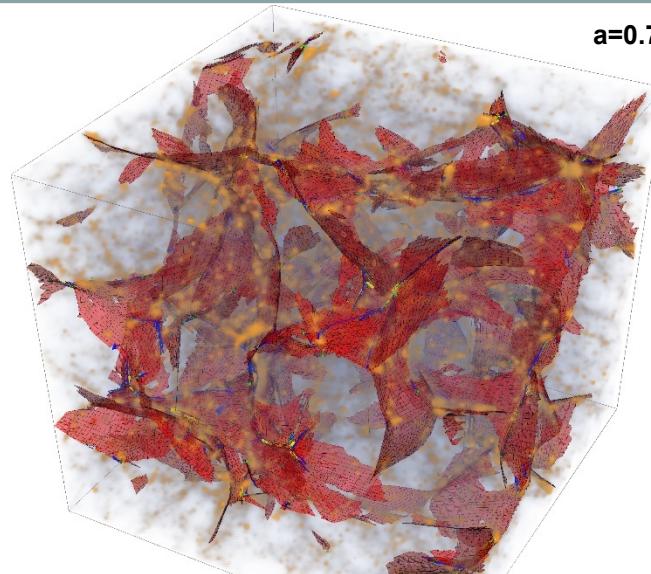
$$\{1 + \mu_i(q_s)\} \vec{v}_i^*(q_s) \cdot \vec{T} = 0$$

NOTE: Nature of singularity not only dependent on EIGENVALUES $\mu_i(q_s)$, but also EIGENVECTORS $\vec{v}_i(q_s)$

Cosmic Web - Evolving Skeleton

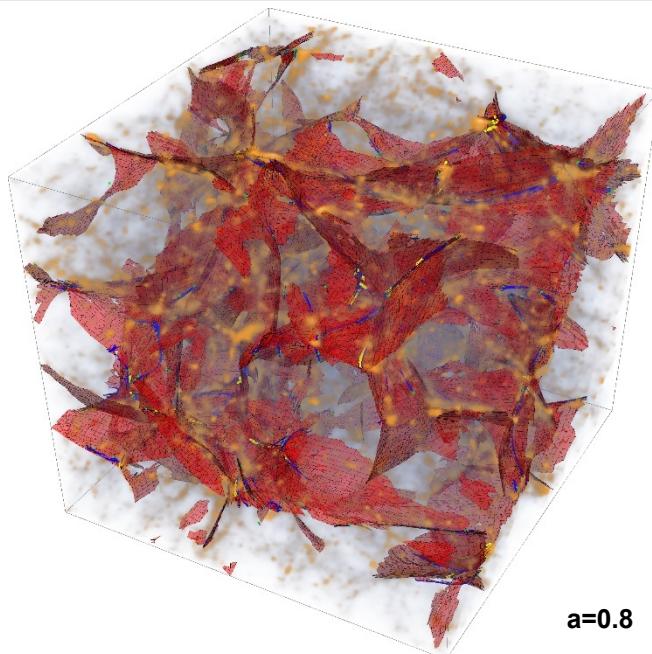


$a=0.6$

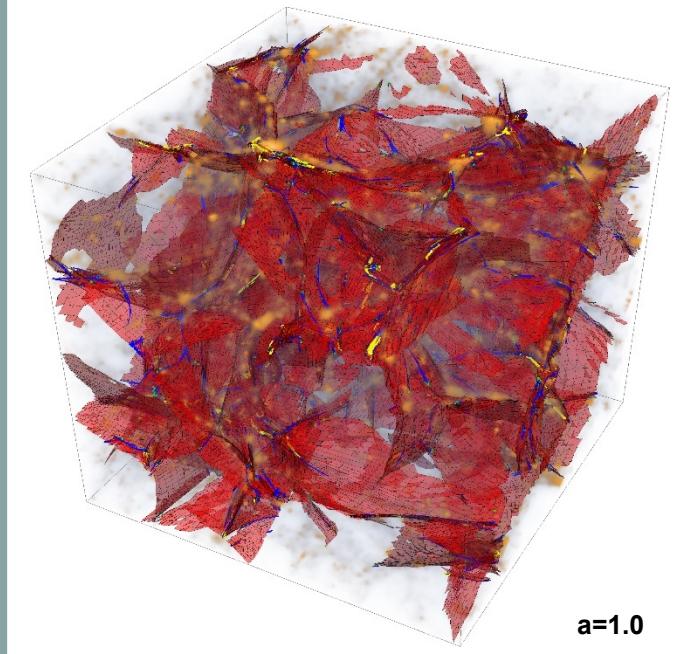


$a=0.7$

Feldbrugge, vdW et al. 2017b



$a=0.8$

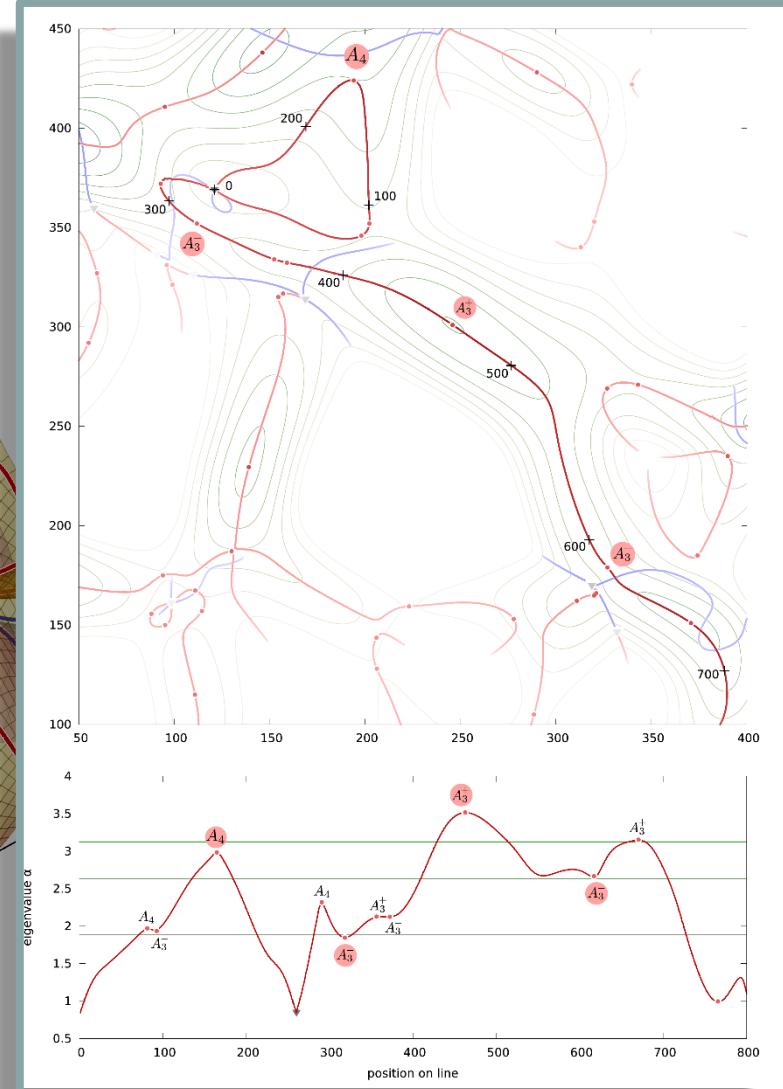
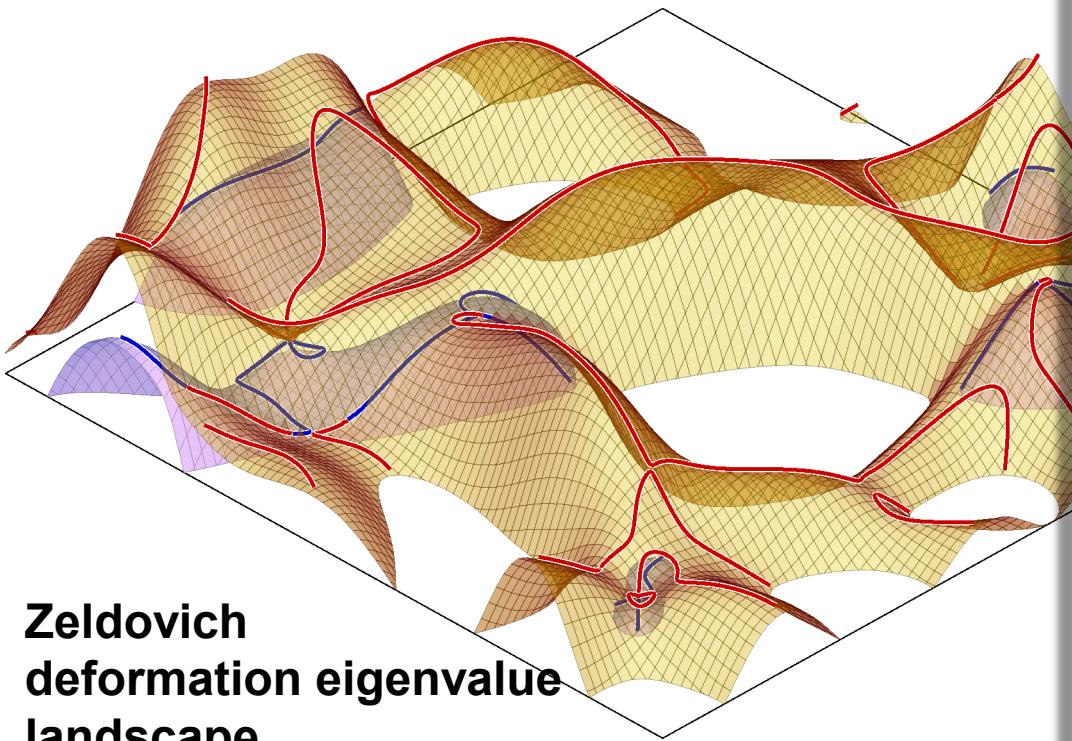


$a=1.0$

Singularities & Catastrophes: Deformation Field

$$\vec{x}(\vec{q}, t) = \vec{q} - D(t) \vec{\nabla} \Phi(\vec{q}) \quad \Rightarrow \quad d_{ij} = \frac{\partial^2 \Phi}{\partial q_i \partial q_j}: \lambda_1, \lambda_2, \lambda_3$$

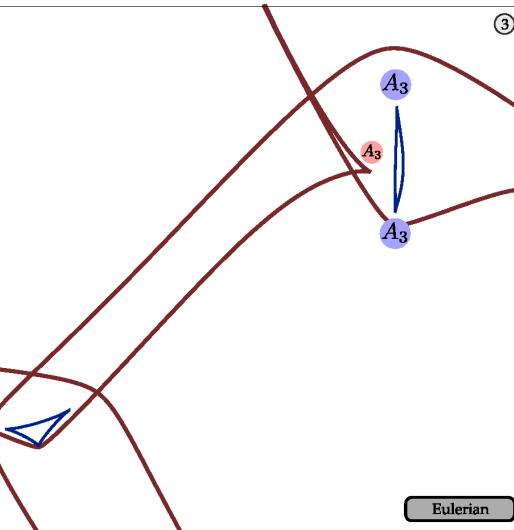
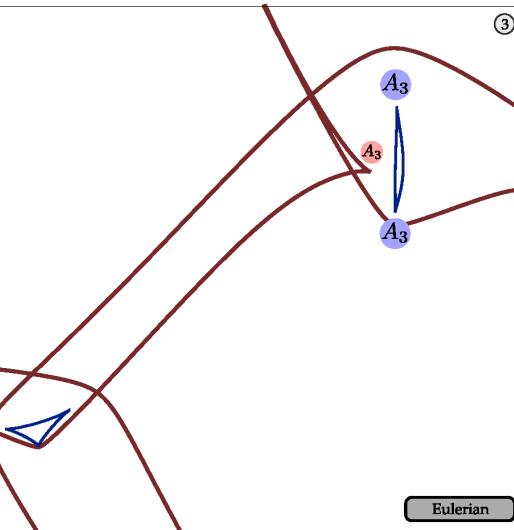
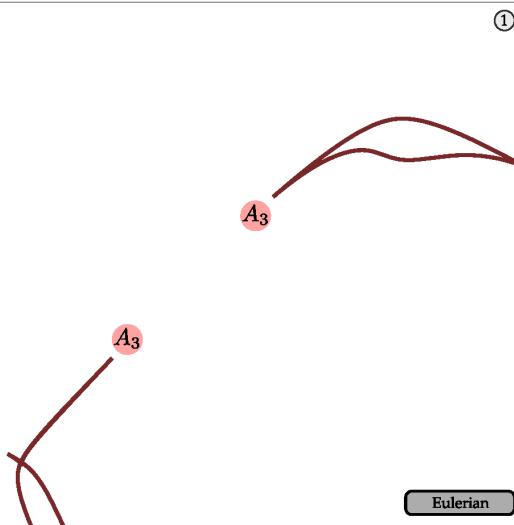
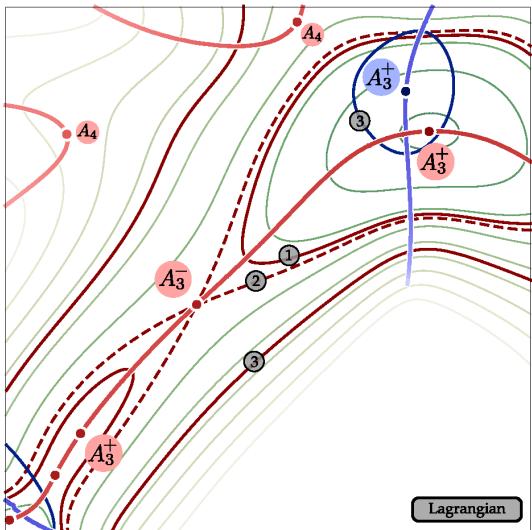
$$\rho(\vec{q}, t) = \frac{\rho_u(t)}{(1 - D(t)\lambda_1(\vec{q}))(1 - D(t)\lambda_2(\vec{q}))(1 - D(t)\lambda_3(\vec{q}))}$$



Caustic Merging & Annihilation

Caustic Hierarchy

Merging & Annihilation of Structures



Evolution & Merging Caustic Features:

Topological transformations in the deformation field:

- Growth of islands ("pancakes") starts at **maxima (A_3^+)** in deformation field λ_i
- Merging of islands at **saddle points (A_3^-)** in deformation field λ_i

The merger of two pancakes.

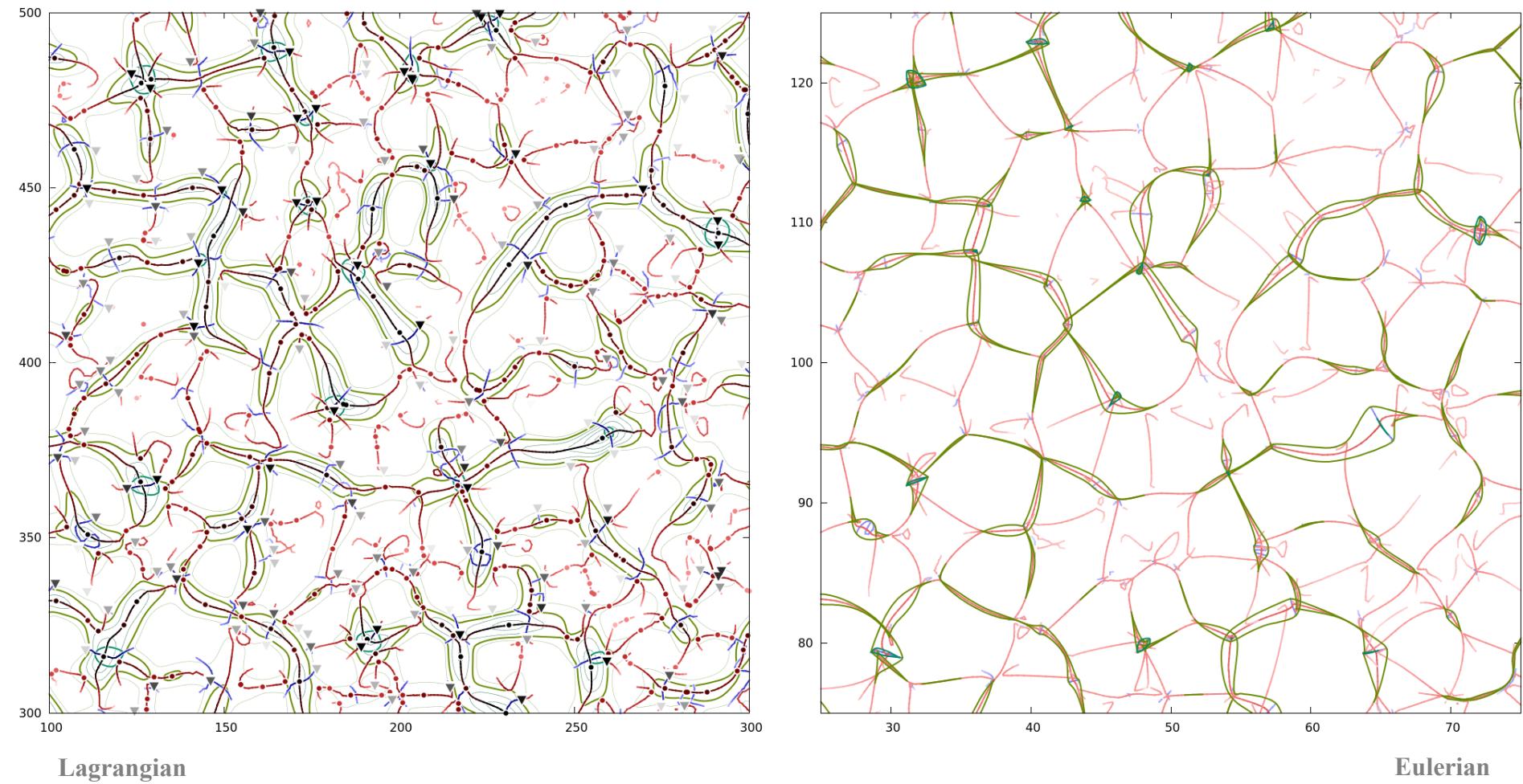
- Top-left: Lagrangian space.
- Other panels: the evolution of caustics in Eulerian space.

Hierarchical Buildup of Cosmic Skeleton:

- Topological transformations at maxima and saddles
- Persistence diagrams (birth-death)

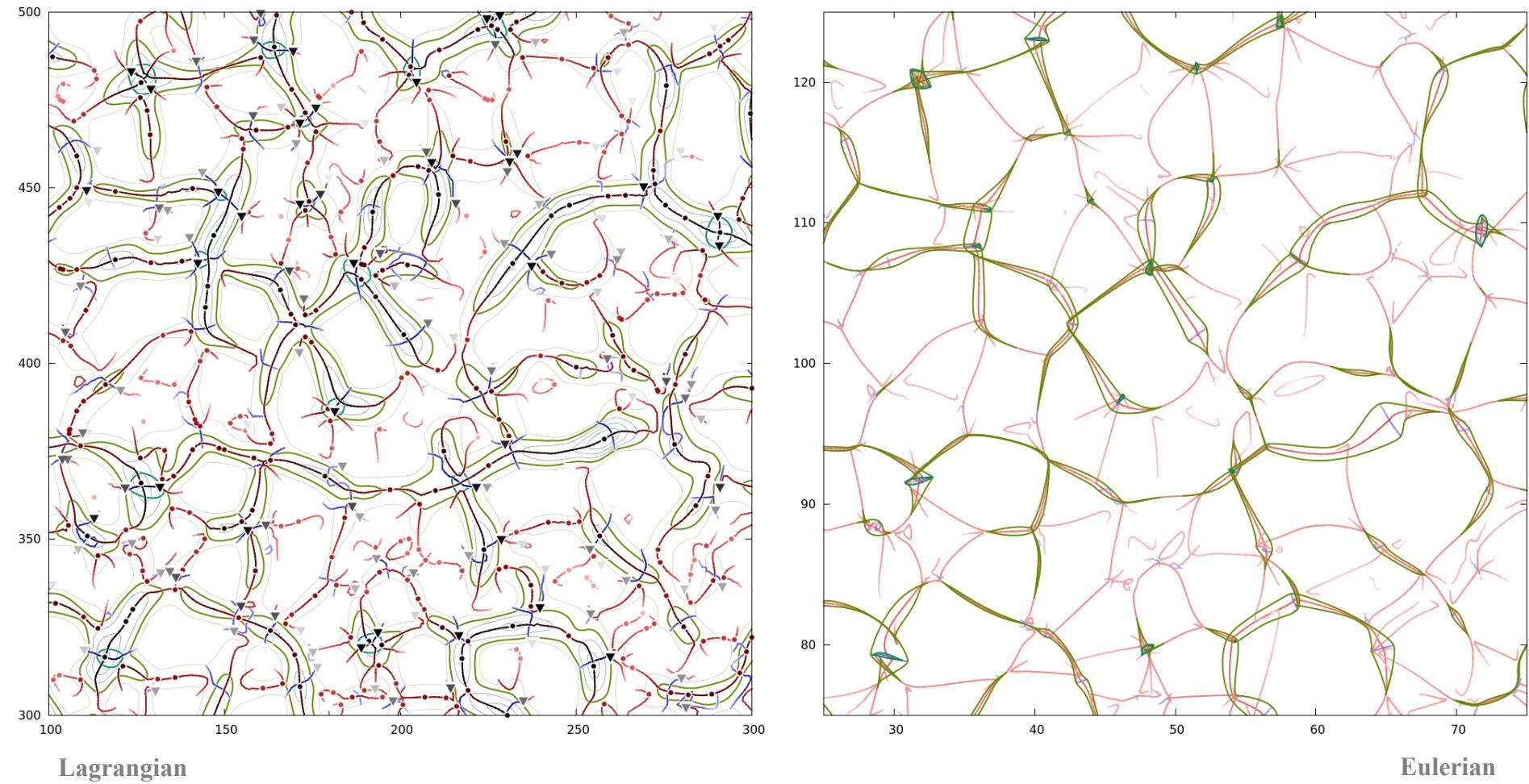
Caustic Hierarchy

Merging & Annihilation of Structures



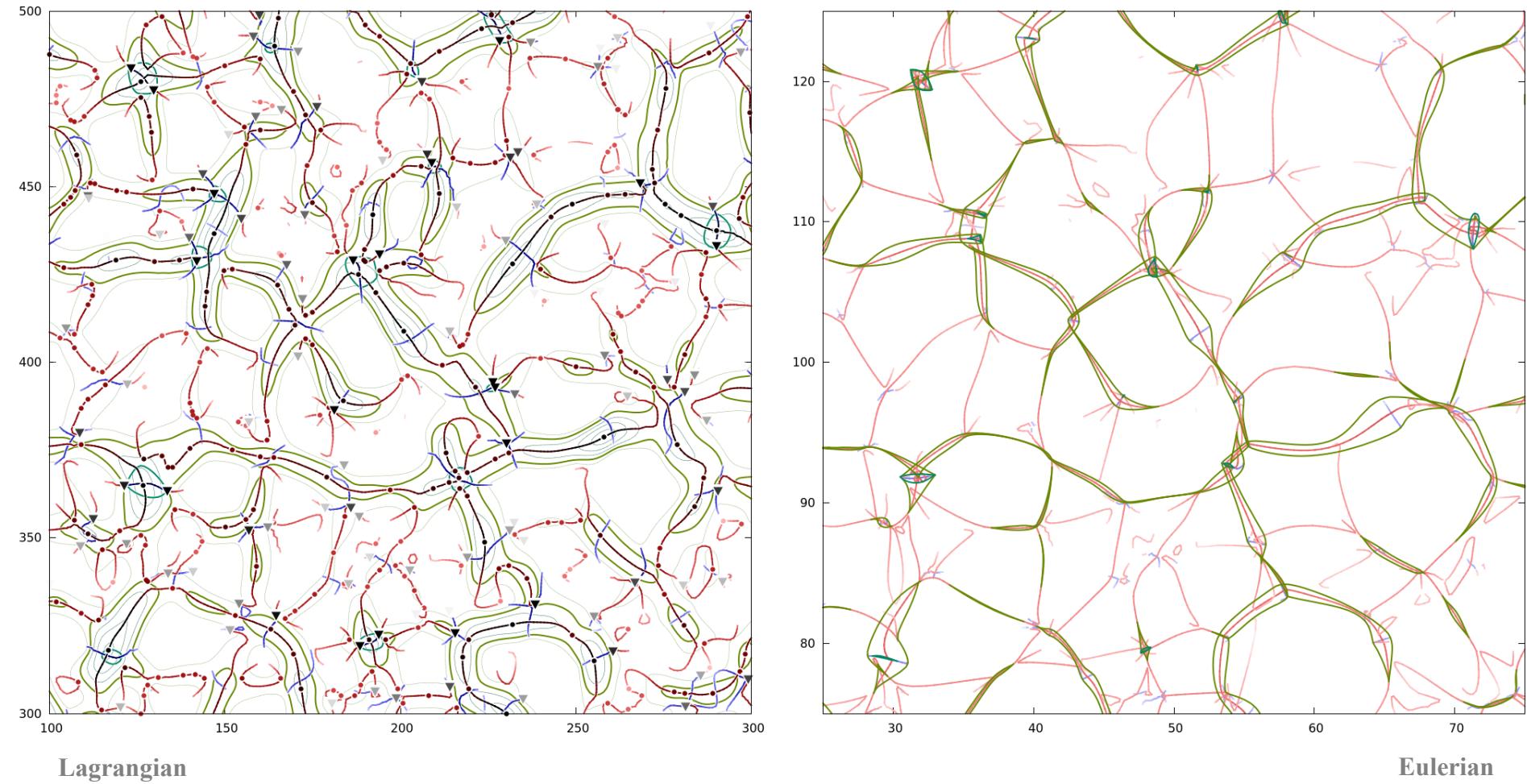
Caustic Hierarchy

Merging & Annihilation of Structures



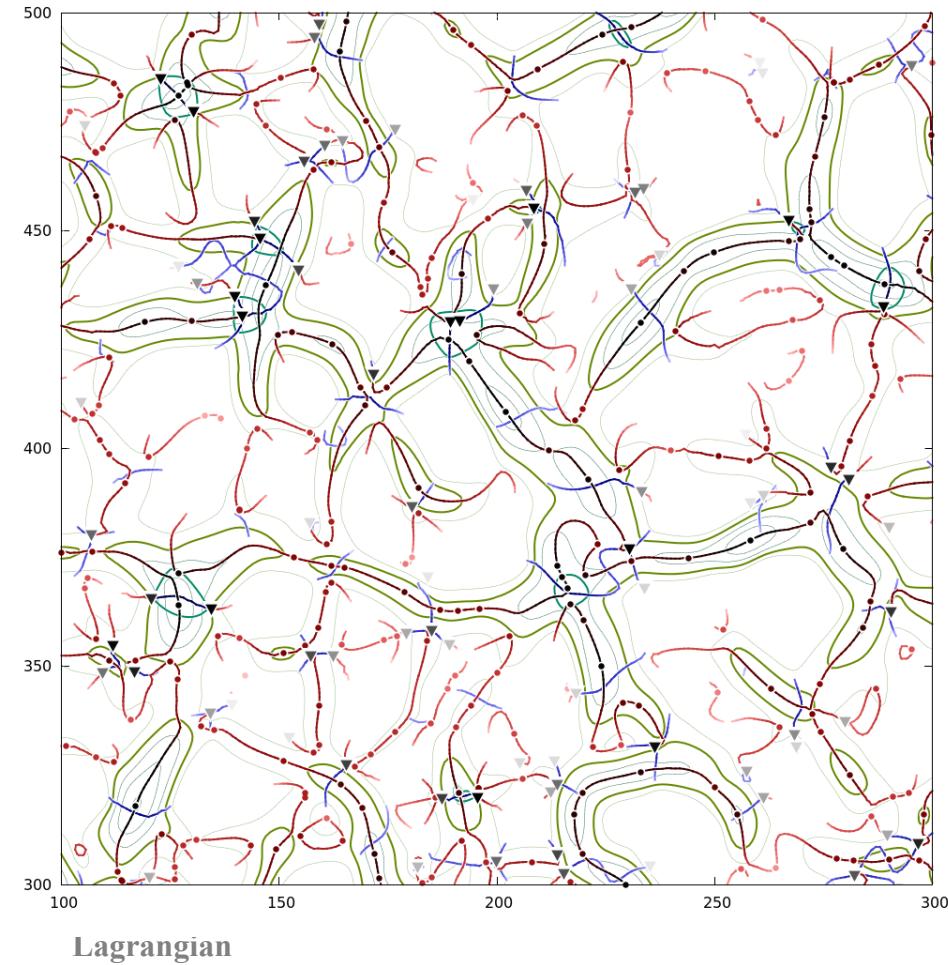
Caustic Hierarchy

Merging & Annihilation of Structures

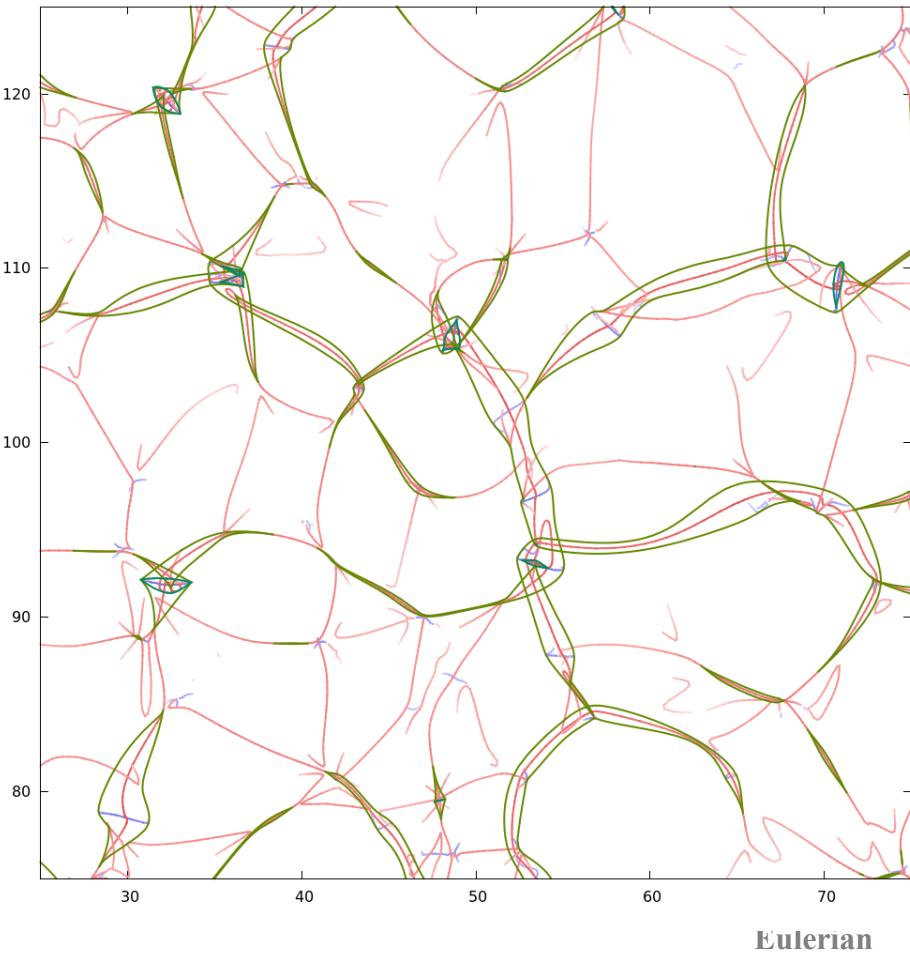


Caustic Hierarchy

Merging & Annihilation of Structures



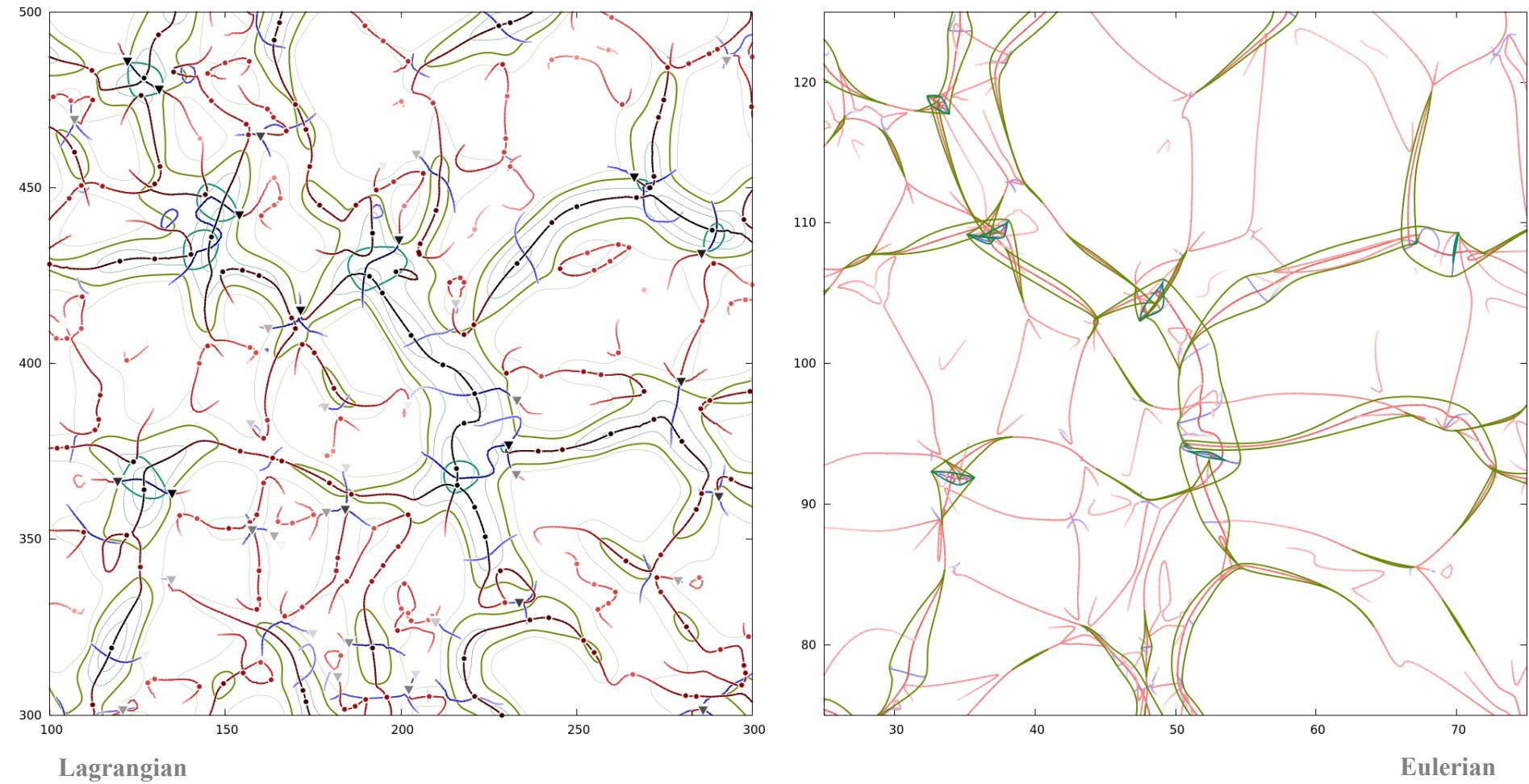
Lagrangian



Eulerian

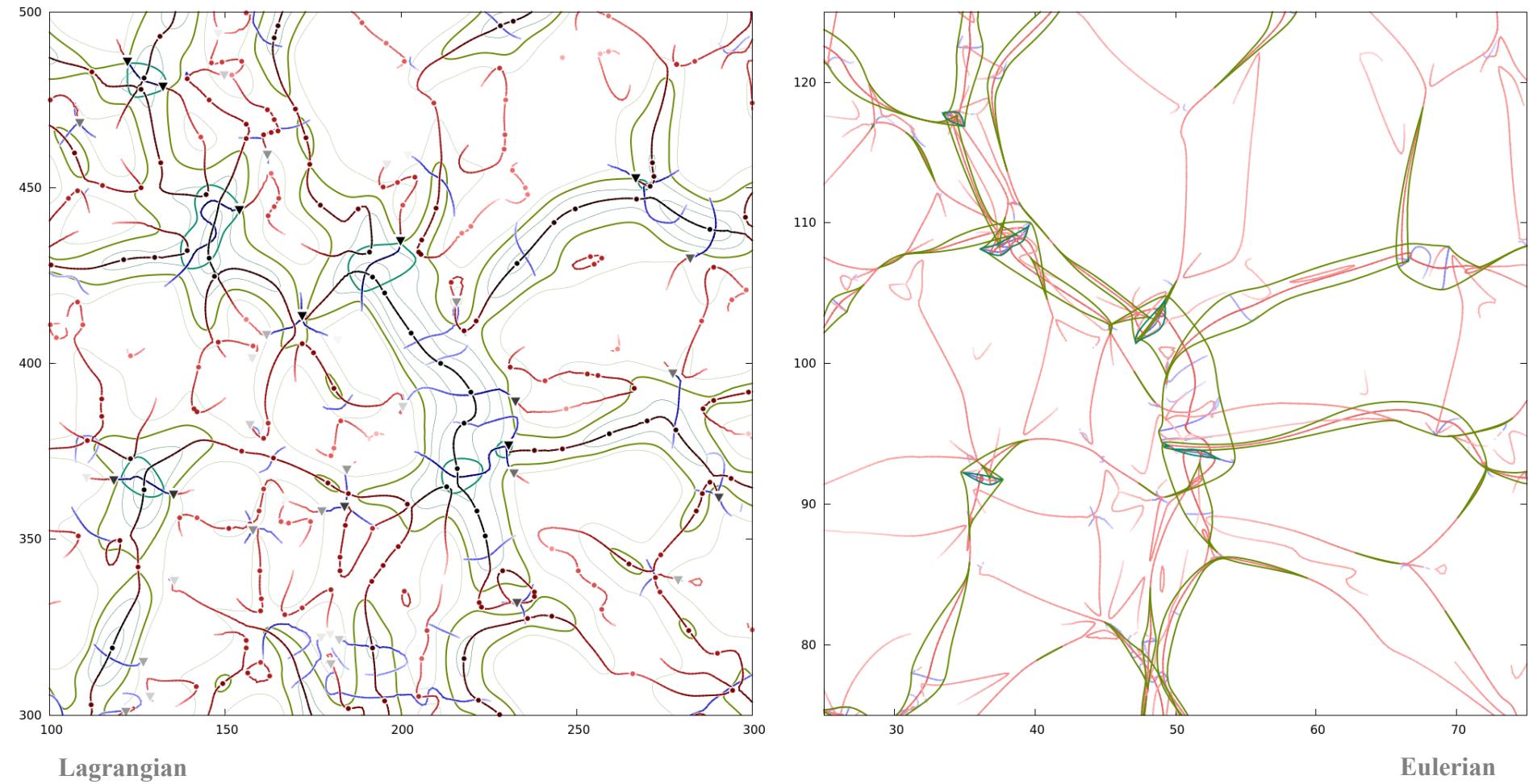
Caustic Hierarchy

Merging & Annihilation of Structures



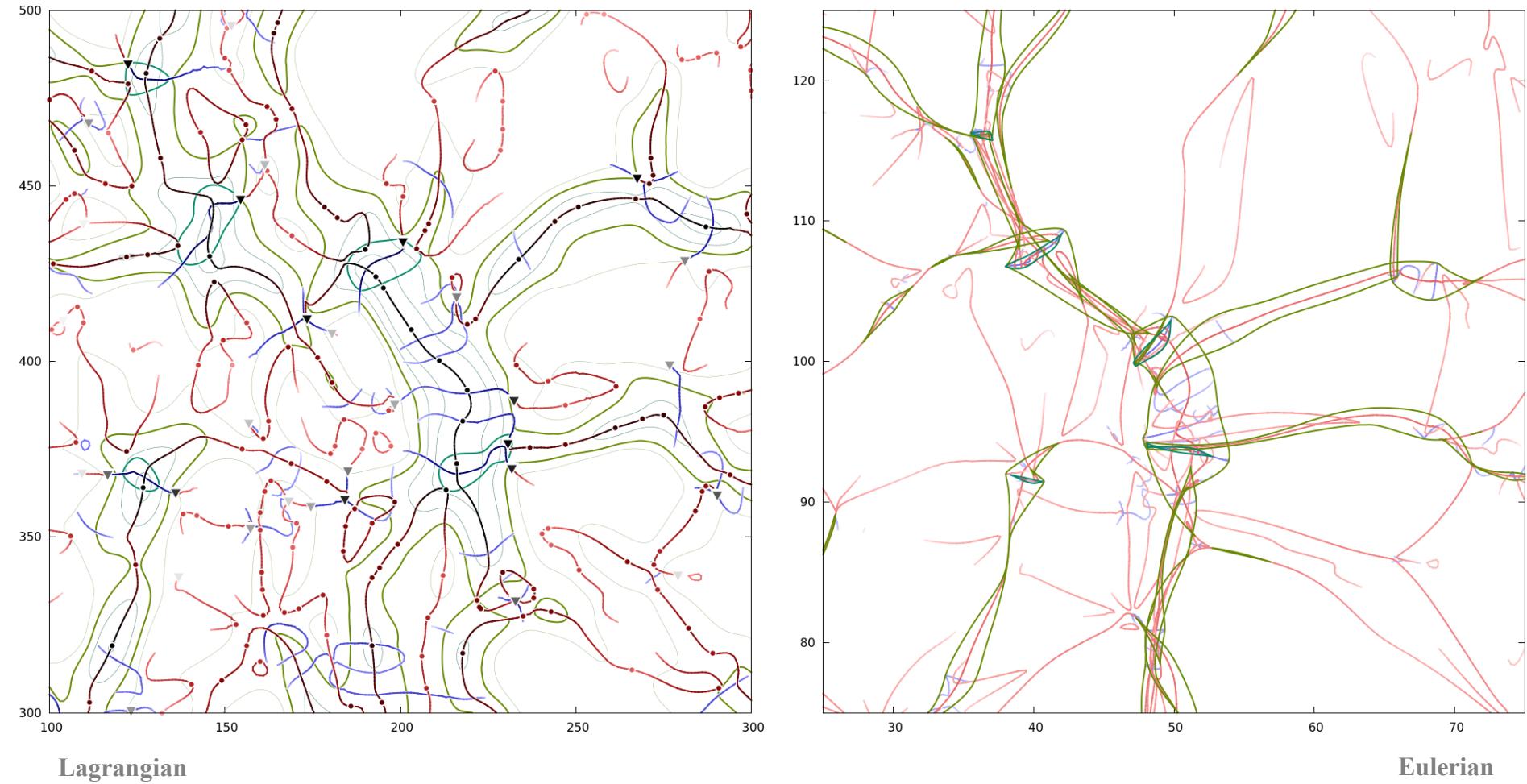
Caustic Hierarchy

Merging & Annihilation of Structures



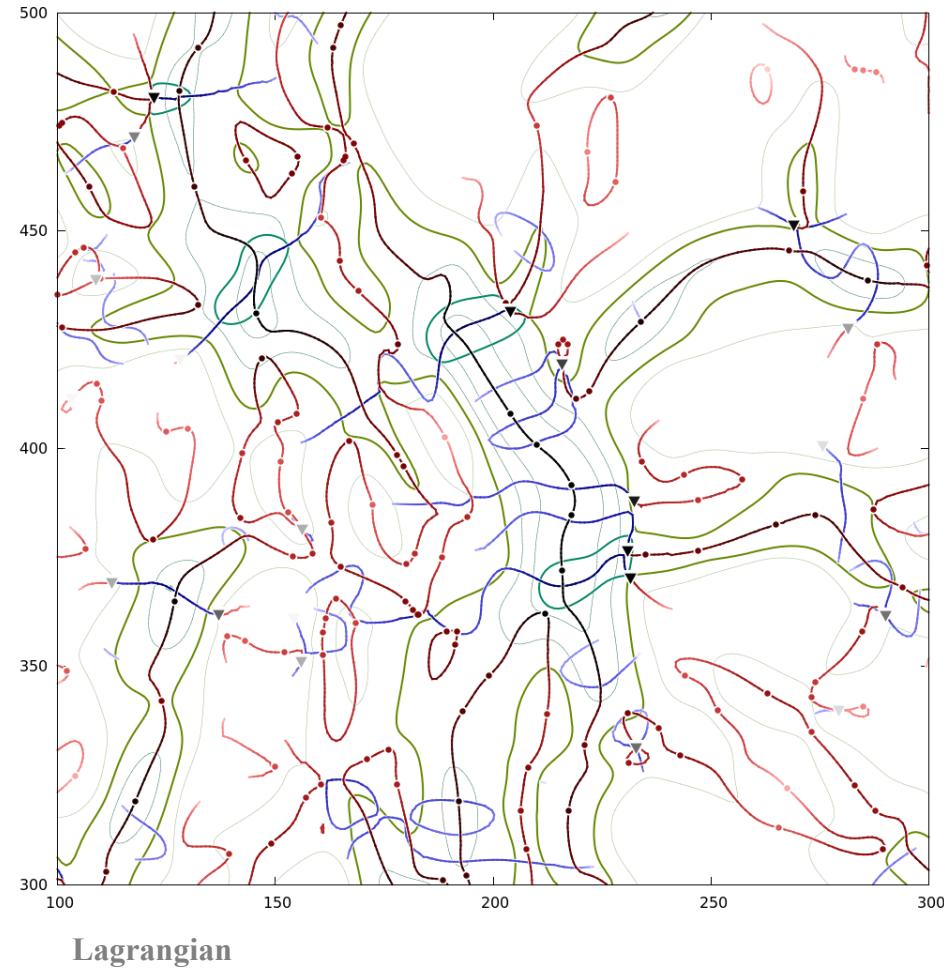
Caustic Hierarchy

Merging & Annihilation of Structures

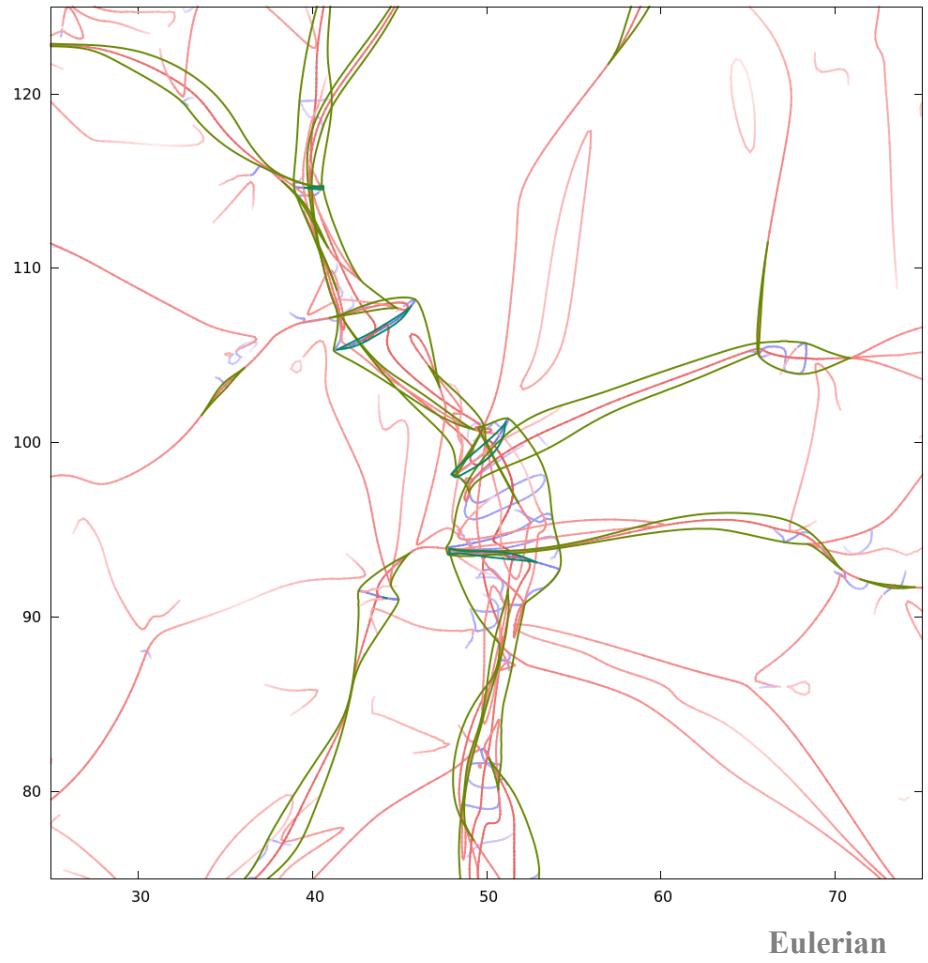


Caustic Hierarchy

Merging & Annihilation of Structures



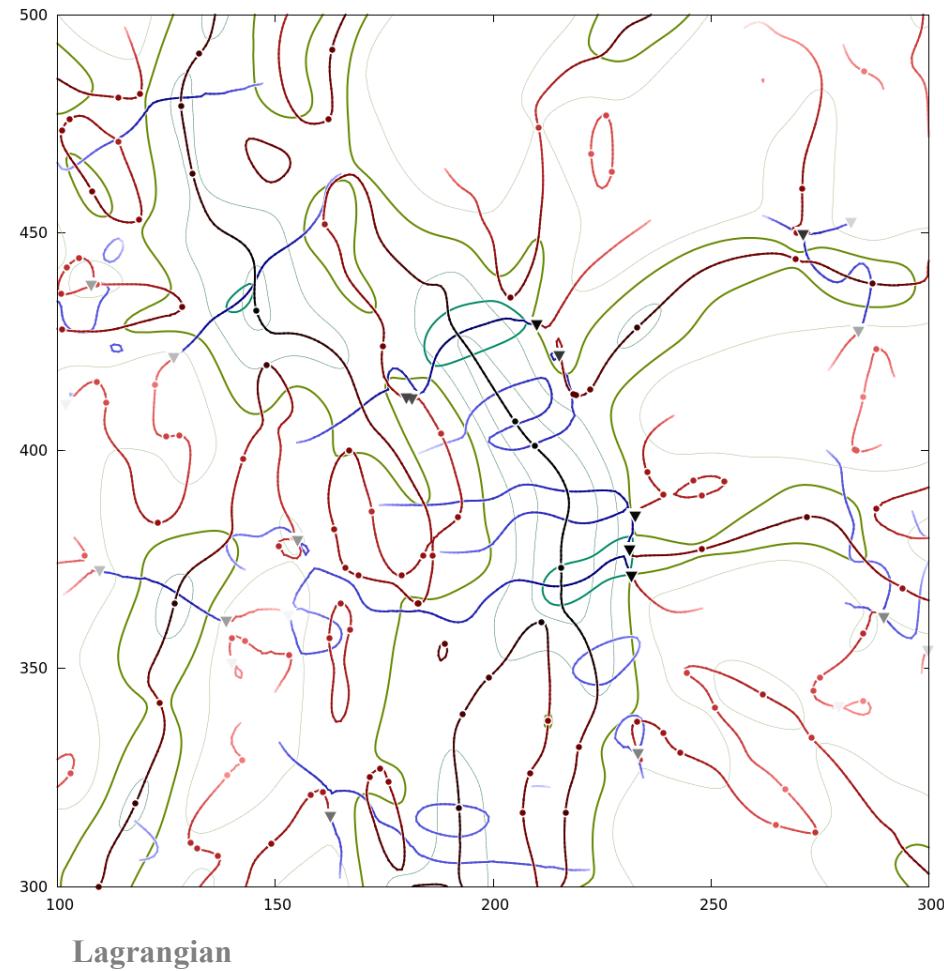
Lagrangian



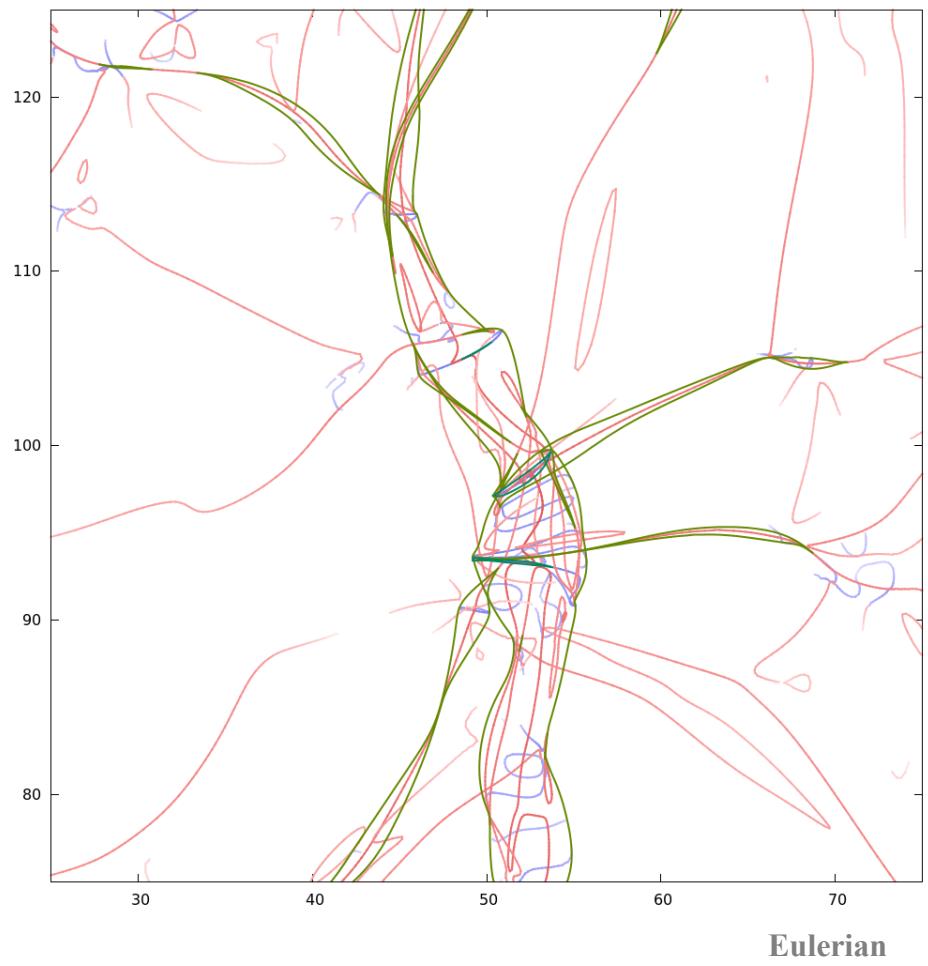
Eulerian

Caustic Hierarchy

Merging & Annihilation of Structures



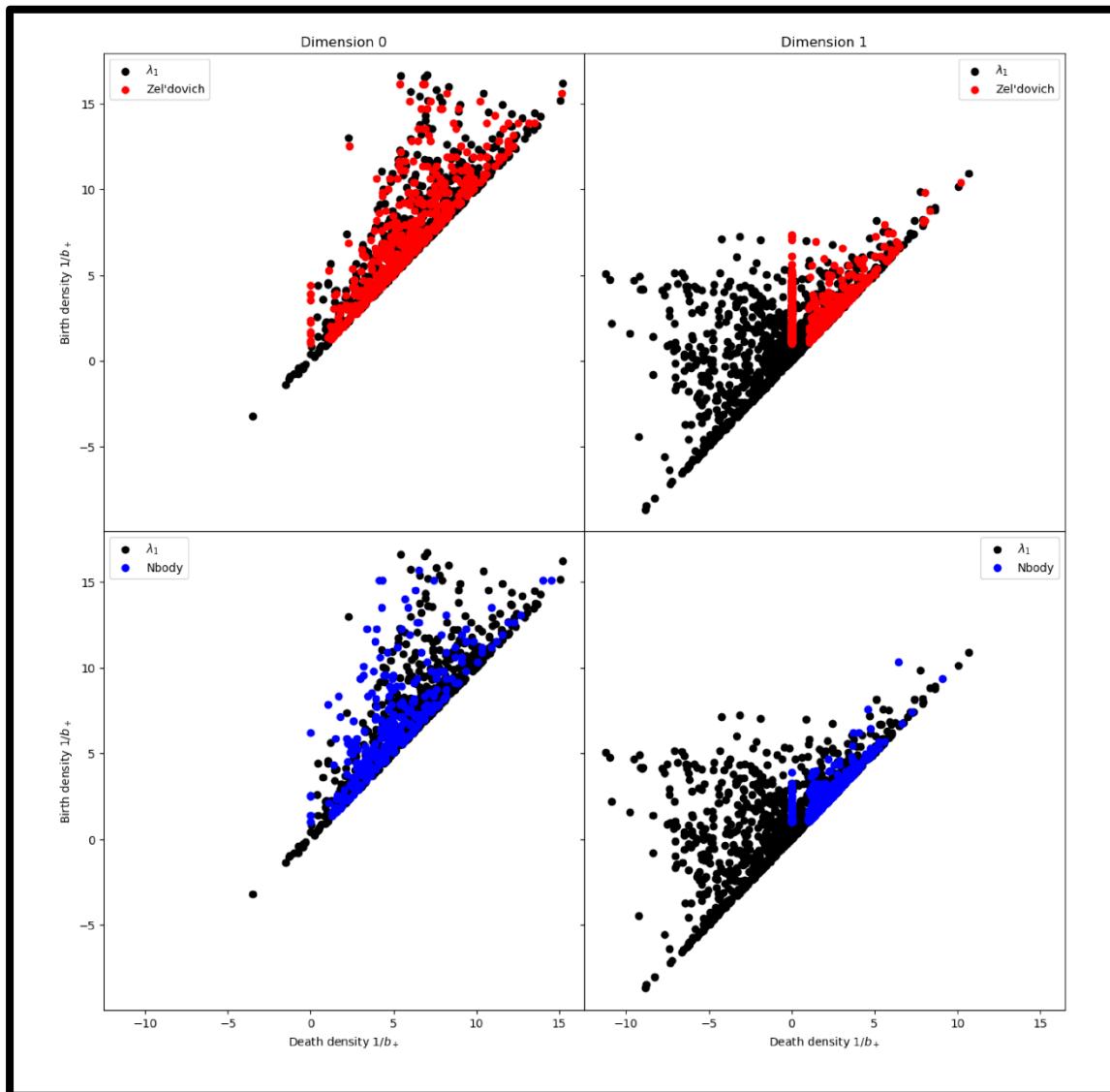
Lagrangian



Eulerian

Caustic Hierarchy & Connectivity

Deformation Persistence





Cosmic Web Persistence